

A Guide to the DiFX Software Correlator

Version 2.8

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1 Introduction

This manual is intended for many different audiences. Typically a particular reader will only need to be concerned with a small portion of this guide, but there are a number of cross-references between sections. This manual assumes some familiarity with Mark5 units, Linux, and the general way in which a VLBI correlator is used. The following topics are discussed: running DiFX, coexistence issues with the VLBA hardware correlator, explanation of various file/document types, and detailed installation instructions. This

manual covers DiFX version 2.5 and will be kept up to date with each official update to DiFX. Please report any errors that are found in this manual to wbrisken@nrao.edu.

1.1 Notation

Text written in **typewriter** font represents literal text and is to be transcribed verbatim when typing and text in *italics* is to be substituted with other text, such as the specific value of the named variable. To be consistent with this notation, all mention of programs by name or filenames (and portions thereof) are written in **typewriter** font.

2 The DiFX correlator

This document is centered around the NRAO installation of the DiFX [2] software correlator and its supporting software. Much of the contents here applies to other installations of DiFX as well, but keep in mind that not a lot of effort is made to generalize these instructions. Fig. 1 shows the general data flow-path within the DiFX software correlator system.

Past, present, and future versions of DiFX as packaged and used by VLBA operations are described in the following subsections.

2.1 NRAO-DiFX 1.0

Versions 1.0 and 1.1 based correlation on the VLBA hardware correlator job scripts – the **.fx** files. This ensures a compatibility period during which both correlators can produce visibilities with expectations of functionally identical results, a feature critical for validation. This strategy also minimizes the required software effort at its earliest phases. Version 1.0 came with the following features:

1. A complete path from **.fx** job scripts to **.FITS** files
2. A command-line only interface
3. Documentation (you are reading it now)
4. Support for VLBA and Mark IV formats
5. Correlation directly off Mark5 modules
6. Support for all projects types except those using special modes, such as pulsars, space VLBI, and near field objects
7. Spectral and time resolution bounded only by practicality

While this version should handle most observations, fast frequency switching and geodesy experiments will produce a large number of output FITS files which may be annoying to observers and the archive. Version 1.0 was available on February 6, 2008.

2.2 NRAO-DiFX 1.1

Version 1.1 builds on version 1.0 and adds the following features:

1. Used version of **mpifxcorr** that has gone through code merge with the official version
2. Blanking of data replaced by headers (Mark4 format only)
3. Proper data weights
4. Initial Mark5B support

5. Support for oversampled data through decimation
6. Multicast status information for GUI interface
7. Correlation of moving and near field objects
8. Concatenation of multiple output files into a single or multiple FITS-IDI file(s)
9. Better support for jobs with multiple configuration tables
10. Playback off Mark5 modules with missing disks
11. Support for Amazon based Mark5 units
12. Completely replaced the “Makefile” system with better integrated alternative
13. Generation of delay model polynomials rather than tables, more like VLBA HW correlator
14. u, v, w values are derived from the delay model (and hence include corrections for aberration, near field observations, and other subtle effects) and are evaluated when writing the FITS file
15. DiFX version accountability
16. Validation of data frames prior to decoding
17. Data evaluation (“sniffing”) built into FITS converter

This version was released on September 3, 2008.

2.2.1 Bugs fixed

Here are listed some of the more important bug fixes:

1. The clock offset was used with the wrong sign in the IM table.
2. Printed precision of some important numbers (RA and Dec) was increased.
3. Autocorrelations were ordered incorrectly for observations with a single polarization.
4. The Mark4 format decoder had a 1 day off bug.
5. The Mark4 format decoder had a $64 \times fanout$ sample timing offset.
6. Several causes of crashes were fixed; no known crashes remain.
7. Missing VLBA monitor data was handled badly.
8. Due to OpenMPI peculiarity, some processing nodes would get most or all of the work in some cases, which cause the work being done on other nodes to be ignored. This was fixed by looking for results in a round-robin manner.
9. Integrations that contain data from two adjacent scans are stripped when writing FITS files.
10. Allow FITS files larger than 2GiB in size.

2.2.2 Known problems

Known bugs as of the NRAO-DiFX 1.1 release:

1. The last couple (typically 2) integrations of a job (not a scan) tend to have low weight due to a premature termination of data processing.

2.3 DiFX 1.5.0

With DiFX 1.5.0 comes a name change. Past releases of this series have been known as “NRAO-DiFX”. The DiFX community has been largely receptive to the NRAO additions in support of `mpifxcorr` and it was decided that dropping the “NRAO” was appropriate. In some cases the term “VLBA DiFX” or “VLBA DiFX 1.5” may be mentioned. These are simply the deployment of DiFX 1.5.0 for the VLBA correlator with some VLBA specific features. Note that the name given to the VLBA deployment of DiFX is formally “VLBA DiFX”.

Version 1.5.0 will start allowing correlation of experiments that cannot be represented by `.fx` files and will be based on `vex` files. Version 1.5.0 builds on version 1.1 and adds the following features:

1. Support for using a wide variety of `vex` files as the basis for correlation.
2. Native ephemeris-based object trajectories are supported.
3. Pulsar gating is supported.
4. Pulsar binning is supported, but not cleanly yet.
5. A graphical user interface is available for correlator operators.
6. The multicast system is fully implemented and is used monitor and control correlation and other operations.
7. Mark5B formatted data, including its 2048 Mbps extension, is supported.
8. The VLBA DiFX Operations Plan [1] is implemented, including interface to the VLBA archive.

Non-NRAO users of DiFX 1.5.0 will still be able to use the tools provided but may not be able to take full advantage of the database back-end without some customization; it is the aim of this document to point out cases where the database is required. Many of the programs described in previous versions of this document will be upgraded or overtaken by more capable replacements.

Release of DiFX 1.5.0 was announced on June 25, 2009.

2.3.1 Bugs fixed

Here are listed some of the more important bug fixes:

1. A rounding issue in `mpifxcorr` occasionally caused the wrong source’s UVWs to be assigned.
2. Lower side band data would come out of the sniffer portion of `difx2fits` with the wrong sign for phase, rate, and delay.
3. Different rounding was used to generate start times for `.input` and `.calc` files. There are no severe consequences of this issue.
4. Scaling in pulsar gating has been made more sane.

2.4 DiFX 1.5.1

DiFX 1.5.1 is mostly a bug fix update to version 1.5.0, but with a few new features. The new features include:

1. Option to force job breaks (with the `break` parameter) has been added to `vex2difx`
2. Time/date formats other than decimal MJD are now accepted by `vex2difx`
3. Specification of data files to correlate (rather than Mark5 units) is supported in `vex2difx`

4. Specification of network parameters in `vex2difx` to allow correlation of eVLBI projects
5. `difx2fits` produces a new output file with suffix `.jobmatrix` provides the user with a better idea of the mapping of jobs into `.FITS` files
6. A `vex2difx` mode for generating DiFX files useful for determining pulsar phase has been added
7. EOP values can now be provided within the `.v2d` file
8. Upcoming FITS-IDI keyword `WEIGHTYP` populated
9. Zero-weight data is not written from `mpifxcorr`
10. New utility `checkdir` to look for oddities in Mark5 module directory files

2.4.1 Bugs fixed

Here are listed some of the more important bug fixes:

1. Concatenation of jobs in the creation of `.FITS` files does the right thing for cases where the antenna subsets change and where antenna reordering is done.
2. The Pulsar Gate Model (GM) `.FITS` file table is now correctly populated for pulsar observations.
3. Autocorrelations are written for each pulsar bin
4. The FXCORR simulator mode of `vex2difx` now selects the correct reference time for antenna clock offsets.
5. A work-around for a Streamstor problem has been added that should improve reliability in Mark5 module correlation when a change in bank is needed.
6. The sign of clock offsets in vex files has been reversed to follow the vex standard
7. Jobs are split at leap seconds
8. LBA data formats are handled more correctly in `vex2difx`
9. The model generator (`calcif2`) now respects polynomial parameters interval and order given on the command line.

DiFX 1.5.1 was made available via subversion on Sep 8, 2009.

2.5 DiFX 1.5.2

DiFX 1.5.2 is mostly a bug fix update to version 1.5.1, but with a few new features. This version of DiFX comes with the following components (and versions): `calcif2` (1.1), `calcserver` (1.2), `difx_db` (1.12; NRAO-only), `difx2fits` (2.6.1), `difx2profile` (0.1), `difxio` (2.12.1), `difxmessage` (0.7), `mark5access` (1.3.3), `mk5daemon` (1.2), `mpifxcorr` (1.5.2), `vex2difx` (1.0.2), and `vis2screen` (0.1). The new features include:

1. Support unmodulated VLBA format data with new pseudo-format “VLBN”
2. `mpifxcorr` now warns when `difxmessage` is in use so the user knows why no messages appear on the screen
3. New utility `difxcalculator` in the `difxio` package
4. eVLBI support within `vex2difx`
5. Vastly improved real-time correlation monitoring

6. New utility `diffDiFX.py` to compare two DiFX output files
7. Improved and more consistent error messages (and some of them are now documented!)
8. `vex2difx` now operates in strict parsing mode by default
9. Additional user feedback to indicate suspicious or bad `.polyco` and `.v2d` files
10. `calcif2` warns if any NaNs or Infs are produced
11. Clock adjustments are easier now with `deltaClock` and `deltaClockRate` parameters in the `vex2difx` antenna settings

2.5.1 Bugs fixed

1. Improve timestamp precision (thanks to John Morgan)
2. The `vlog` program (used at NRAO only, I think) misparsed the pulse cal information in some cases
3. Fixed memory leak in `difx2fits` when combining a large number of jobs
4. Improved FXCORR simulation mode in `vex2difx`
5. Mark5 directory reading systematically generates unique names for all scans even when two scans have the same name
6. Improve reporting of Mark5 errors during playback and change alert severity to be more appropriate
7. Don't overblank certain Mark4 modes (thanks to Sergei Pogrebenko for the bug report)
8. Vex 'data valid' period now properly respected
9. Vex clock table tolerance issue corrected
10. Changes in Mark5 mode should be safer (note that currently `vex2difx` never exercises multiple modes in a single job)
11. When making the cross spectrum sniffer plots, respect the reference antenna
12. Improved pulsar polynomial file error checking is performed
13. Amplitude-phase-delay (APD) sniffer plots always have refant first when multiple refants are supplied
14. Project name should now appear on sniffer APD plots
15. Mark5 units now send status information even when no playback is occurring (eliminating the incorrect LOST state issue as displayed in the DOI)

2.5.2 Known problems

1. Extensive use in VLBA operations has shown that occasional data dropouts of one or more antenna, sometimes in a quasi-repeatable manner, affect completeness of some jobs. It is not clear exactly what the cause is at this point, however its cure is a high priority.
2. Loss of a few FFTs of data will occur in rare circumstances.
3. Clock accountability is poor when jobs containing multiple clock models for antennas are combined.

DiFX 1.5.2 was made available via subversion on Jan 20, 2010.

2.6 DiFX 1.5.3

DiFX 1.5.3 is mainly intended as a bug fix update to version 1.5.2, though some new features have made their way into the codebase. This version of DiFX comes with the following components (and versions): `calcif2` (1.3), `calcserver` (1.3), `difx_db` (1.13; NRAO-only), `difx2fits` (2.6.2), `difx2profile` (0.2), `difxio` (2.12.2), `difxmessage` (7.2), `mark5access` (1.3.4), `mk5daemon` (1.3), `mpifxcorr` (1.5.3), `vex2difx` (1.0.3), and `vis2screen` (0.2). Many changes are motivated by issues found running DiFX full time in Socorro.

The new features include:

1. Mark5 directory (`.dir`) files can contain RT on the top line to indicate the need to play back using *Real-Time* mode.
2. `difxqueue` (NRAO only) now takes an optional parameter specifying the staging area to use.
3. New Mark5 diagnostic programs (`vsn` and `testmod`) introduced to wean off the use of the `Mark5A` program.
4. `mk5daemon` can now mount and dismount USB and eSATA disks through `mk5commands`.
5. `mk5cp` now makes the destination directory if it doesn't exist.
6. `mk5daemon` will now warn if free disk space is getting low.
7. `db2vex` (NRAO only) now allows field station logs to be provided. As of now, only media VSNs are extracted.
8. Playback off Mark5 units has been made more robust with better error reporting.
9. New utility `m5fold` that can be used to look at repeating signals in baseband data total power (e.g., switched power)
10. `vex2difx` now supports job generation in cases where upper side band was observed at one antenna and lower sideband at another.

2.6.1 Bugs fixed

1. Don't unnecessarily drop any FFTs of data.
2. Sub-integrations longer than one second could cause integer overflows.
3. Fix bug in `vex2difx` where jobs were not split at clock breaks.
4. `difx2fits` was guilty of incorrect clock accountability after a clock change at a station when merging multiple jobs. Worked around by not allowing such jobs to merge.
5. `db2vex` (NRAO only) warns when more than one clock value is found for an antenna.
6. Mark5 unit bank switches now routinely call `XLGetDirectory()` to work around a newly discovered bug in the StreamStor software.
7. A couple possible memory leaks in the `mark5access` library were fixed (thanks Alexander Neidhardt and Martin Ettl).
8. Lots of compiler warnings quashed (mostly of the "unused return value" kind).
9. Olaf Wuchnitz found two FITS file writing problems in `difx2fits` dating back to code inherited from FXCORR!

10. Two more digits are retained for the time and one more digit is retained for amplitude information in the `.apd` and `.apc` sniffer files.
11. Some bugs related to replacement of special characters by “entities” in XML messages are fixed.
12. New traps are in place in many places to catch string overruns.
13. Fix for writing `.calc` files with more than one ephemeris driven object.
14. `vex2difx` would get *very* slow due to constantly sorting a list of events. Now this list is only sorted when necessary, drastically speeding it up.
15. The `RCfreqId` parameter in the `difxdatastream` structure (in `difxio`) was used with two different meanings that are normally the same. Cases where they differed caused exceptions. Fixed in `difxio` and `difx2fits`. (Thanks to Randall Wayth for leading to the discovery)
16. `difx2fits` would assign a bogus `.jobmatrix` filename when not running the sniffer.
17. `vex2difx` could get caught in an infinite loop when making jobs where two disk modules had zero time gap.
18. `difx2fits` used a bad config index when making the pulsar GM table when multiple configs were present.
19. Within `mpifxcorr` an extra second was added to the validity period for polycos to ensure no gap in coverage.
20. `mk5dir` would add correct the date improperly for Mark4 formats after beginning of 2010.
21. Lots of fixes for building FITS files out of a subset of baseband recorded channels.
22. FITS files now support antennas with differing numbers of quantization bits.
23. Lots of Mac OS/X build issues fixed.

DiFX 1.5.3 was released on April 16, 2010.

2.7 DiFX 1.5.4

DiFX 1.5.4 is likely the last 1.5 series formal release of DiFX, though an additional release could be made if demand is there.

The new features include:

1. `difx2fits` can now produce FITS files with only a subset of the correlated sub-bands.
2. `difx2fits` can be instructed to sniff on an arbitrary timescale.
3. The `makefits` wrapper for `difx2fits` now respects a `-B` option for phase bin selection.
4. `difxio` has improved checking that prevents merging of jobs with incompatible clocks.
5. `difxio` now maintains a separate `clockEpoch` parameter for each antenna.
6. `difxStartMessage` now contains DiFX version to run, allowing queued jobs to be run under different DiFX versions.
7. The curses utilities `mk5mon` and `cpumon` now catch exceptions and can be resized without infecting the terminals they are run in.

8. New sub-library called `mark5ipc` added that provides a semaphore lock for Mark5 units.
9. The `testdifxmessagereceive` utility can now filter on message types.
10. Support for SDK9 throughout (e.g., in `tt mpifxcorr`, `mk5daemon`, and other utilities).
11. Support for new Mark5 module directory formats (Haystack Mark5 memo 81).
12. Several new Mark5 utilities to make up for Mark5A functionality that will no longer be available: `vsn`, `testmod`, `recover`, `m5erase`.
13. `mk5cp` can now copy data based on byte range.
14. Many programs directly talking to the StreamStor card of Mark5 units use WATCHDOG macros for improved diagnostics when problems occur.
15. More protection against incomplete polyco files added to `mpifxcorr` (Note: should add this to `vex2difx` as well).
16. The GUI can now spawn different DiFX versions at will through the use of `difxVersion` parameter in the `DifxStartMessage` and wrapper scripts.
17. `difx2fits` can now convert LSB to USB for matching purposed. When used, all LSB sub-bands must have corresponding USB sub-bands on one or more other antenna.
18. `mark5access`-based utilities (e.g., `mp5spec`) can now read from stdin.
19. New utility `mk5cat` can send data on a Mark5 module to stdout.

2.7.1 Bugs fixed

1. Only alt-az telescopes received the correct model. Fixed. Note that CALC and FITS-IDI don't have a good match between their sets of allowed mount types.
2. `difx2fits` now properly propagates quantization bits on a per antenna basis.
3. Logic errors in `difxio` would confuse `difx2fits` in cases where different antennas use different frequency setups. Fixed.
4. Weights are blanked in `difx2fits` prior to populating each record, preventing screwy weights for unused sub-bands.
5. `vex2difx` would sometimes hang or not converge on job generation. Fixed.
6. `vex2difx` now doesn't assume source name is same as vex source def identifier.
7. `mpifxcorr` generated corrupted weights and amplitudes when post-FFT fringe rotation was done. Fixed.

2.8 Known bugs

1. Tweak Integration Time feature of `vex2difx` often does the wrong thing.

DiFX 1.5.4 was released on October 12, 2010.

2.9 DiFX 2.0.0

DiFX 2.0.0 is based on an upgraded `mpifxcorr` that breaks `.input` file compatibility with the 1.0 series. This new version will allow more flexible correlation of mis-matched bands and correlation at multiple phase centers along with general performance improvements. Development of the 2.0 capabilities will occur in parallel with the 1.0 series features.

2.9.1 New features

1. Pulse cal extraction in `mpifxcorr`.
2. Massive multi-phase center capability.
3. New utility `zerocorr` added.
4. External pulse cal extraction utility `m5pcal` added.
5. DiFX output format is all-binary, meaning speed and disk savings

2.10 Known bugs

1. Zoom band support has multiple problems.

DiFX 2.0.0 was released on October 12, 2010.

2.11 DiFX 2.0.1

DiFX 2.0.1 is a bug fix and clean-up version in response to numerous improvements to DiFX 2.0.0. There are a number of new features as well.

2.11.1 New features

1. New utility `checkmpifxcorr` to validate DiFX input files
2. Switched power detection in `mpifxcorr`
3. Early multi-thread VDIF format support
4. RedHat RPM file generation for some packages (can extend to others on request)
5. Improvements to method of selecting which pulse cal tones get propagated to FITS
6. Initial complex sampling support
7. Improved locking mechanism for direct mark5 access (using IPC semaphores; `difxmessage`)

2.11.2 Bug fixes

1. Fix model accountability bug in `difx2fits` when combining jobs
2. Numerous fixes for zoom bands (in `mpifxcorr`, `vex2difx` & `difx2fits`)
3. Native Mark5 has improved stability for cranky modules
4. Numerous fixes for DiFX-based phase cal extraction (mostly in `difx2fits`, mostly for multi-job)
5. Fractional bit correction for a portion of lower sideband data got broken in `difx 2.0.0`. Fixed.
6. Migrate `difxcalculator` to DiFX 2; was not complete for DiFX 2.0.0

DiFX 2.0.1 was released on June 24, 2011.

2.12 DiFX 2.1

2.12.1 New features

1. Mark5-based correlation: easy access to S.M.A.R.T. data (can be viewed with `getsmart`)
2. Mark5-based correlation: emit multicast message containing drive statistics after each scan
3. VSIS interface added to `mk5daemon`
4. Support for non power-of-2 FFT lengths
5. New utilities: `mk5map` (limited functionality), `fileto5c`, `record5c`
6. Remote running of `vex2difax` from `mk5daemon`
7. Multithread VDIF support enabled for the data sources `FILE` and `MODULE`, including stripping of non-VDIF packets
8. New features added to existing utilities:
 - `mk5cp`: copy without reference to a module directory
 - `mk5cp`: ability to send data over ssh connection
 - `vsn`: get SMART data from disk drives
9. e-Control source code analysis (Martin Ettl, Wettzell)
10. Restart of correlation is now possible
11. `difax2fits`: -0 option to write minimal number of visibilities to FITS
12. `difax2fits`: write new RAOBS, DECOBS columns in source table
13. `tweakIntTime` option to `vex2difax` has been re-enabled
14. `diffDiFX.py` can now cope with two files that don't have exactly the same visibilities (i.e., some visibilities are missing from one file)
15. `plotDiFX.py` and `plotDynamicSpectrum.py` now have better plotting and more options
16. New FAKE datastream type for performance testing
17. Espresso, a lightweight system for managing disk-based correlation, has been added to the DiFX repository.
18. Option to correlate only one polarization has been added.
19. `mk5dir` can now produce `.dir` file information for VDIF formatted data.
20. Add NRAO's sniffer plotters to the repository.

2.12.2 Bug fixes

1. LBA format data now scaled roughly correctly (removing the need for large ACCOR corrections).
2. There was a bug when `xmaclength` was $> \text{nfftchan}$ for pulsar processing. This has been corrected.
3. `guardns` was incorrectly (overzealously) calculated in `mpifxcorr`.
4. `Mk5DataStream::calculateControlParams: bufferindex>=bufferbytes` bug fixed.
5. Low weight reads could result in uninitialized memory; fixed.
6. Streamstor `XLRead()` bug work-around installed several places (read at position 0 before reading at position > 0). This is thought not to be needed with Conduant SDK 9.2 but the work-around has no performance impact.
7. Fix to pulse calibration data ordering for LSB or reordered channels.
8. Pulse cal amplitude now divided by pulse cal averaging time in seconds.
9. Pulse cal system would cause crash if no tones in narrow channel. Fixed.
10. Zoom band support across mixed bandwidths (see caveat below).
11. Fix for spurious weights at end of jobs (untested...)
12. Mixed 1 and 2 bit data are handled more cleanly
13. `mpifxcorr` terminates correctly for all short jobs. Previously it hung for jobs with a number of subints between `nCores` and $4 \times \text{nCores}$
14. Correctly scale cross-correlation amplitudes for pulsar binning when using `TSYS > 0` (accounts for varying number of samples per bin c.f. nominal)
15. Lower side-band pulse cal tones had sign error. Fixed.

DiFX 2.1 was released on May 25, 2012.

2.13 DiFX 2.1.1

DiFX 2.1.1 was a minor patch release to fix a scaling issue with autocorrelations of LBA-format data in `mpifxcorr`.

DiFX 2.1.1 was committed as a patch to DiFX 2.1 on June 7, 2012.

2.14 DiFX 2.2

2.14.1 New features

1. `calcif2`: ability to estimate delay polynomial interpolation errors
2. Support for a “label” identifier for a local version of DiFX that will help discriminate exact version used.
3. Faster Mark5 directory reading
4. Faster VDIF corner turning through customized bit shifting functions
5. `mpifxcorr` can now be built without Intel Integrated Performance Primitives, though resulting in a slower correlator.

6. **vdifio**: several new VDIF manipulation and processing utilities added: **vmux**, **vsum**, **vdifd**, **vdifspec**, **vdiffold**, **vdifbstate**
7. **difxbuild**: a new installation program
8. **difxspeed**: a program to benchmark and help optimize DiFX

2.14.2 Bug fixes

1. Mutex locking bugfix for very short jobs
2. Prevent MODE errors when a datastream runs out of data well before the end of a job
3. **calcif2**: fix azimuth polynomial generation in case of wrap
4. Fix for FITS file generation for mixed sideband correlation
5. **difx2fits** now uses appropriate gain tables for S and X band in S/X experiments (Thanks to James Miller-Jones for reporting)
6. **difx2fits**: correct pcal, weather, tsys and flag data for observations crossing new year
7. Fixed scaling of autocorrelations for LBA format data
8. 0.5 ns wobble in delays for 2 Gbps Mark5B data fixed
9. Fix bug preventing subintegrations longer than 1 second. Now 2 seconds is allowed (this limit comes from signed integer number of nanoseconds).
10. Weights corrected in cases where two setups differencing only by pcal setup were correlated against each other
11. Quashed data and weight echos that would occur for about 1 integration at the beginning of each scan for datastreams that ran out of data before end of job.
12. The multicast (diagnostic) weights were low or zero in case of frequency selection (zoom band or freqId selection). Fixed.
13. **mark5access**: fix (non)blocking issue when receiving data from *stdin*

DiFX 2.2 was released on June 12, 2013.

2.15 DiFX 2.3

2.15.1 New features

1. **mpifxcorr**: LO offsets are now corrected in the time domain when fringe rotation is also done in the time domain (the usual mode), allowing considerably larger LO offsets without decorrelation
2. **mpifxcorr**: Working polarization dependent delay and phase offsets
3. **mpifxcorr**: Experimental linear2circular conversion
4. **mpifxcorr**: Complex Double sideband (RDBE/Xcube) sampling support (Note: things are not perfect here; wait for 2.4 for real use)
5. **mpifxcorr**: new file/Mark5 based VDIF/Mark5b datastream (faster and more robust)
6. **mpifxcorr**: implement work-around for buggy kernel-driver combinations; Mark5 read sizes ≥ 20 MB now allowed

7. utilities: some new command line tools for Mark5B and VDIF files (vsum, mk5bsum, vmux, mk5bfix)
8. new options for passing calibration (Walter B: memo forthcoming)
9. Hops updated to version 3.9

2.15.2 Bug fixes

1. mpifxcorr: Datastream buffer send size now calculated correctly for complex sampled data
2. mpifxcorr: Avoid very rare bug where combination of geometric delay and data commencing mid-subint meant one invalid FFT might be computed
3. mpifxcorr: multicast weights are now computed correctly for mixed-sideband correlation
4. mpifxcorr: fixed bug where some autocorrelations were not saved in a mixed-sideband correlation
5. mpifxcorr: fixed bug where send size could be computed incorrectly by 1-2 bytes for
6. Mark4/VLBA/Mark5B/VDIF formats, potentially resulting in very small amounts of data loss

DiFX 2.3 was released on January 18, 2014.

2.16 DiFX 2.4

2.16.1 New Features

1. mpifxcorr
 - Support a FAKE correlation mode for multi-threaded VDIF.
 - The mpifxcorr produced PCAL files have had a format change that allows unambiguous interpretation across all use cases.
 - Add network support (TCP, UDP and Raw Ethernet) for multi-threaded VDIF: 1. TCP and UDP variants not tested yet; 2. raw Ethernet variant is used for the VLITE project.
 - Support updated Mark5 module directories.
 - Better checking that Mark5 data being processed matches what is expected.
 - Improved Mark5B decoding: 1. Mark5B data streams are now filtered for extra or missing data; 2. packets with invalid bit (actually the TVG bit) set replace missing data; 3. this means any valid Mark5B data with the TVG bit set will not correlate.
 - Information about each Mark5 unit used in “native mode” is emitted at start of jobs so it can be logged.
 - Ultra-low frame rate VDIF data was affected by allowing a long “sort window” in the VDIF multiplexer. This has been reduced to 32 frames and seems to work fine for all bandwidths now.
2. difx2fits
 - Slightly improved compliance with the FITS-IDI convention: 1. invalid Tsys values become NaN, not 999; 2. populate DELTAT keyword in ModelComps table.
3. difxio
 - Support for X/Y polarization correlation. Many fundamental issues with linear polarization remain though: 1. this does not support in a meaningful way Linear*Circular correlations; 2. there is a terminology gap in many bits of software and file formats that confuses X/Y with H/V polarization bases; 3. the intent of this support is for short baselines (VLITE).

4. mark5access

- Support for “d2k” mode in Mark5B format (swapped sign and mag bits).
- fixmark5b() function fixed for case that fill pattern is seen at the 1 second transition.
- Make use of the TVG bit as an “invalid frame” indicator for Mark5B data.
- m5bstate: support complex sampled data

5. mk5daemon

- New utility mk5putdir: reads a binary file and replaces a Mark5 directory with it.
- mk5dir: when reading the directories, saves a copy of the binary representation in case it is needed later (perhaps via mk5putdir)
- Reworked mark5 module directory support, including support for many new variants of the directory format.
- mk5erase will save a “conditioning report” to \$MARK5_CONDITION_PATH if that environment variable is set.

6. vdifio

- Fairly large change to the API. Please read the ChangeLog for details.

7. vex2difs

- Respect the record enable bit in the SCHED block. If that value is 0 no correlation will be attempted for that antenna.
- Bug fixes preventing some LSB/zoom bands from being correlated.
- Complex data type and number of bits are now read from vex file.

2.16.2 Bug fixes

1. A “jitter” of 0.5 ns when using 2Gbps Mark5B format was fixed. A fix was back-ported to DiFX 2.3.
2. A similar jitter was corrected for high frame rate VDIF (problem identified by the VLITE project)
3. Fix case of intermittent fringes that was due to incorrect assumption about the sizeof(unsigned long): 32 bits on a 32-bit system vs. 64 bits on a 64-bit system. Some other variable types were changed for long term type safety
4. Fix off-by-one in correlation using LSB and zoom bands together.

2.17 DiFX 2.5.1

2.17.1 New features

1. Initial support for correlating Mark6. This is still much a work in progress.
2. Multiple datastreams per antenna supported via **vex2difs**
3. New delay model program: difxcalc11.? No longer requires calcserver.
4. Support for more than 6 days of EOP values.
5. “Union mode” in difx2fits allows merging of correlation output that uses different setups. Some restrictions apply. Designed for GMVA and RadioAstron use.

6. Improved VDIF support: wider range of bits/threads, support for multi-channel, multi-thread VDIF, support for complex multi-thread VDIF
7. Support for new VDIF Extended Data Version 4 which is useful for multiplexed VDIF data. See: <http://vlbi.org/vdif/docs/edv4description.pdf>
8. Python bindings for vdifio and mark5access
9. mpifxcorr: per-thread weights implemented
10. Automatic selection of arraystride by mpifxcorr if set to zero; this is done per-datastream.? Very useful for correlation of ALMA data or others with non-standard sample rates.
11. Automatic selection of xmacstride by mpifxcorr if set to zero
12. Automatic selection of guardns by mpifxcorr if set to zero
13. mpifxcorr can now operate with unicast messages instead of multicast. Useful in some situations where multicast is not supported.
14. New “dirlist” module/file directory listing format.
15. mk5cp append mode to resume interrupted copy
16. ALMA support in HOPS: non-power-of-two FFTs, up to 64 freq. channels, full linear/circular/mixed polarization support
17. HOPS improvemetns for VGOS through improved manual phase cal support
18. New package: polconvert. Used to post-correlation convert from linear to circular polariations
19. New package: autozoom. Helps a user develop .v2d file content when setting up complicated zoom band configurations.
20. New package: datasim: generate baseband data suitable for simulated correlation
21. Improved error reporting in many places

2.17.2 Bug fixes

1. fix for incorrect reporting of memory use by mpifxcorr (needed longer int sizes)
2. dataweights would sometimes be incorrect after abrupt ending of data from a datastream.
3. FITS-IDI files produced by difx2fits more standards compliant; fix problem that caused AIPS task VBGLU to fail.
4. Several segfaults across a number of programs/utils now are caught and provide useful feedback.

2.17.3 Caveats

1. Various changes made between DiFX 2.4 and 2.5 are not API-compatible. Please don't mix packages from these two releases. If you have non-DiFX software that links against the DiFX libraries, be sure to recompile them. A small number of changes may result in need to restructure such code.
2. Unlike previous DiFX releases, each tagged version will be its own SVN copy. If the number of minor releases within the 2.5 series gets large, some (reversible) pruning of the SVN repository may occur. There has been some debate about the best tagging strategy: bring any strong opinions to the Bologna meeting, where further changes to release and tagging policies can be discussed if needed.

2.18 DiFX 2.5.2

2.18.1 Bug fixes

1. Fixes for the HOPS 'rootid' rollover. The new rootcode is a conventional base-36 timestamp in seconds from the start of the new epoch (zzzzzz in the old epoch). This will last until 2087.
2. Major fixes/improvements to PolConvert for use with ALMA by the EHTC and GMVA.

2.19 DiFX 2.5.3

2.19.1 Updates

1. genmachines
 - r8264 genmachines and mark6 datastream updates
 - r8357 add mark6 activity message to mark6 datastream
 - r8409 allow multiple nodes to serve as datastream nodes for FILE based-data in the same location
2. hops 3.19 new features
 - increased the number of allowed frequency “notches” to ridiculous levels
 - an ad hoc data flagging capability to allow improved time / channel data selection for fringing
 - a capability to dump all the information on the fringe plots into ascii files for roll your own plotting
 - removed obsolete `max_parity`
 - introduced `min_weight` (to discard APs with very little correlated data in support)
 - vex2xml, a program that converts VEX (v1.5) into XML to allow easy parsing via standard XML parsers.
 - added `type_222` to save control file contents, enabled by keyword `gen_cf_record`.

polconvert to v1.7.5 (mostly minor bug fixes and robustifications)

DiFX 2.5.3 was released on March 28, 2019.

2.20 DiFX 2.5.4

2.20.1 New features

1. HOPS updated to 3.22
2. Vex2difx and difxio
 - permit H and V polarization labels (but need trunk difx2fits if these are to be propagated into enhanced FITS-IDI)
 - support the v2d parameter 'exhaustiveAutocorr'
 - support Mark6 MSNs i.e. MSNs that contain '%'
3. Difx2mark4
 - support PCal data from multi-datastream correlation
 - option '-e expt_nr' additionally propagates the experiment number into the generated root files
 - backported the options -w for mixed-bandwidth data, and -g for filtering freq groups
4. Genmachines updated to support Mark6 host auto-detection
5. Includes copies of more recent utilities: packHops.py, distFourfit.py, fplot2pdf, plotResiduals.py

2.20.2 Bug fixes

1. Mark6 native playback fixed to support other than VLBA VDIF recording
2. Vdifio
 - fix excessively long station name printout in printVDIFheader
 - fix to mk6gather to not crop a scan
 - resync code extended with further VDIF frame sizes that are common in VGOS
3. Mpifxcorr fix for complex data (VDIFC) PCal extraction for bands with other than 16 tones
4. Mpifxcorr fix for IVS-related 5/10MHz pcal issue specific to those channels where first pcal is at baseband 2.01 MHz
5. Difx2mark4
 - fix count of total bands and polarizations for multi datastream datasets, solving an issue in HOPS post-processing
 - fix parallactic angle calculation error in first t303 record
 - fix to permit Tab characters in VEX
6. Mpifxcorr no longer segfaults after “exiting gracefully” message upon missing data files
7. Tkinter add-on package renamed from tkinter to kinter_difx to avoid name collision
8. Install-difx now works under Python 3 and supports the option withmark6meta
9. Startdifx no longer piles up difxlog processes

2.20.3 Caveats

Support for Complex VDIF is incomplete in 2.5.3 and 2.5.4, both treat Complex VDIF LSB as USB data. This handling was retained in 2.5.4 for VGOS compatibility reasons in order to have all-LSB(-mislabelled) data products, and avoid postprocessing issues with mixed USB and LSB data products. One should not correlate any actual Complex VDIF LSB recordings - this produces no fringes. The issue affects only VDIFC (complex VDIF), not VDIF nor other formats.

DiFX 2.5.4 was released on August 27, 2021.

2.21 DiFX 2.5.5

2.21.1 New features

1. genmachines extended to support Mark6 with multiple expansion chassis
2. update polconvert scripts

2.21.2 Bug fixes

1. genmachines fix to support December i.e. doy 335 and later
2. vex2difx
 - fix freqClockOffs and loOffsets parameters to not expect more values than recorded frequencies
 - removed obsolete warning about 10 MHz PCal not being supported

DiFX 2.5.5 was released on October 24, 2022.

2.22 DiFX 2.6.1

2.22.1 New features

1. Improved VDIF support
 - Increased robustness in processing VDIF data with many gaps
 - Improvements in processing VDIF with frame sizes very different from 5000 bytes
 - New in-line reordering functionality via `vdifreader()` functions; allows operation on more highly skewed VDIF files
2. `mpifxcorr` `.input`, `.calc`, `.threads`, and pulsar files are now only read by the head node
3. `mpifxcorr` can be provided a new stop time via a `DifxParameter` message; results in clean shutdown at that time.
4. `mpifxcorr` can extract pulse cals with tone spacing smaller than 1 MHz
5. Support for Intel Performance Primitives version ≥ 9 (specifically IPP 2018 and 2019)
 - These newer IPP versions are more readily available than earlier versions
6. Improved support for Mark6 playback
 - Mark6 activity messages in `difxmessage`
 - Support in `genmachines` with updated `mk5daemon`
 - Support playback of Mark5B data on Mark6
 - New and improved mark6 utilities
7. `difx2fits`: populate antenna diameters and mount types for antennas known to the difxio antenna database
8. `difx2fits`: in verbose mode, explain why files are being split
9. `difx2fits`: new options for merging correlator jobs run with different clock models
10. `vex2difx`: new parameter `exhaustiveAutocorrs` can be used to generate cross-hand autocorrelations even when the two polarizations for an antenna come from different datastreams
11. `difx2mark4`: support multiple bandwidths in one pass
12. `hops`: to rev 3.19 (see notes on 2.5.3 above for details on several new and useful features)
13. `polconvert`: to rev 1.7.5 (see notes on 2.5.3 above for details)

2.22.2 Bug fixes

1. `mpifxcorr`: Retry on NFS open errors of kind: “EAGAIN Resource temporarily unavailable”
2. `mpifxcorr`: Fix weight issue when the parameter `nBufferedFFTs` > 1
3. `startdifx/genmachines`: Fixes for cases when multiple input files are provided
4. Python 2 scripts now explicitly call `python2`
5. `vex2difx`: allow up to 32 IFs (was 4) and warn when this is exceeded
6. `vex2difx`: support units in the clock rate (e.g., usec/sec); in general support time in the numerators.
7. Sun RPC is on its way out; support for “tirpc” added to `calcif2` and `calcserver`

2.22.3 Caveats

1. Moved mark6gather functions from vdifio to mark6sg; this changes the order of dependencies!
2. Various changes made between DiFX 2.5 and 2.6 are not API-compatible. Please don't mix packages from these two releases. If you have non-DiFX software that links against the DiFX libraries, be sure to recompile them. A small number of changes may result in need to restructure such code.
3. There is some suspicion that correlation of very narrow bandwidth VDIF modes on Mark6 media can result in premature termination of datastreams.
4. The `.threads` file must now exist; previously (before the change to only have manager read these files), a missing `.threads` file would cause each core process instance to have a single thread.
5. `difx_monitor` won't compile with `IPP ≥ 9`

DiFX 2.6.1 was released on August 28, 2019.

2.23 DiFX 2.6.2

2.23.1 New features

1. Python parseDiFX package added

2.23.2 Updates

1. HOPS updated to version 3.21
2. PolConvert updated to version 1.7.8
3. Former FTP access to CDDIS servers changed to FTP-SSL (`geteop.pl`)

2.23.3 Bug fixes

1. `mpifxcorr`: Fix correlation of Complex LSB data, restore fringes. Note: DiFX 2.5.x and 2.6.1 treated Complex LSB as if Complex USB, while Trunk prior to r9647 05aug2020 treated LSB nearly correctly except for a off-by-one channel bug
2. `difx2mark4`: Fix seg-fault in `createType3s.c` when a station has only a single entry in the PCAL file
3. `difx2mark4`: Remove unneeded debugging statement (calling `d2m4_pcal_dump_record()`)
4. `difx2mark4`: Update `createType3s.c` to add support for DiFX PCAL files generated from station data where each data-stream thread resides in a separate file (multi-datastream support). This separates the code reading the PCAL files from the code filling the type-3 records so that tone records from multiple data streams can be merged before populating the type-309s
5. `difx2mark4`: Update `createType3s.c` to remove support for DiFX version-0 PCAL files
6. `difx2mark4`: Add support for 10 MHz p-cal tone spacing (needed by VGOS at Yebes)
7. `difx2mark4`: Significantly increase hardcoded array sizes (`difx2mark4.h`: `NVRMAX` 8M, `MAX_FPPAIRS` 10k, `MAX_DFRQ` 800) as required for EHT2018
8. `difx2mark4`: Fix a bounds check, permit tabs in VEX file
9. `difx2fits`: Fix FITS PH table having missing or superfluous pcal records when one correlates multi-datastream antennas, or not all recorded frequencies, or multiple zooms per recorded frequency

10. mark6gather: Fix poor weights in native Mark6 correlation for VDIF frame sizes not equal to 5032 bytes
11. difxio: Fix PCal tone frequency rounding bug on some platforms
12. difxio: Cope with recorded bands that lack PCal tones, e.g., 200 MHz PCal spacing of KVN with say 32 MHz recorded bands
13. calc11: Dave Gordon provided ocean loading params at EHT stations
14. calc11: Increased the number of field rows supported in .calc files
15. vex2difx: Fix internal merge of SamplingType (real, complex) when info found in VEX and/or v2d file
16. Minor changes to oms2v2d and vexpeek
17. More IPP versions supported
18. Minor issues with vis2screen fixed
19. Fixed build failure with gcc defaults
20. Python3 support in many/most places

2.23.4 Caveats

1. difx2mark4: Some LSB-LSB baselines do not get converted in mixed-sideband correlation setups (DiFX 2.6.1, 2.6.2); if affected, use difx2mark4 2.5.3 with override-version. A bugfix is pending for DiFX 2.6.3 later this year.
2. difx2mark4: Performance regression with p-cal files, conversion of p-cal data may take noticeably longer than before
3. calcserver and difxcalc11: With the latest versions of gfortran (10.1 or newer) you will need to uncomment the line with -fallow-argument-mismatch line in the environment setup in order to compile. Users who do this should be alert to possible issues.

DiFX 2.6.2 was released on September 11, 2020.

2.24 DiFX 2.6.3

This version has never been officially released.

2.25 DiFX 2.7.1

This version has never been officially released. The 2.7 series was used by the Event Horizon Telescope. The 2.7 series is a feature branch derived from DiFX 2.6.2 with support added for outputbands.

2.26 DiFX 2.8.1

The DiFX 2.8 series collected updates and fixes from many DiFX users and is the first version universally usable on VGOS, EHT, and “traditional” VLBI data. Because the 2.7 series was never formally released, the notes below include all changes since the 2.6 series.

2.26.1 New features

1. Support for appending contiguous subbands together to create outputbands
2. Initial support for vex2
3. Complete conversion to Python 3
4. Support for CODIF format
5. mark5access programs: error output to stderr to allow piping
6. Support for IPP 2019 series
7. Experimental support for Vienna Mapping Functions in calcif2
8. Some features (within mpifxcorr and vex2difx) that provide additional options for real-time correlation
9. Espresso modification to work with singularity (or docker) image

2.26.2 Updates

1. HOPS updated to version 3.24
2. PolConvert updated to version 2.0.3
3. Several new VDIF decoders and corner turners introduced to widen range of support
4. Improved support for more than 2 bits per sample (see https://library.nrao.edu/public/memos/vlba/up/VLBASU_52.pdf)
5. Mark6: support for larger number of expansion units
6. Add a few more stations to ocean loading tables within difxcalc
7. Several new options in the tabulateddelays program
8. difxio programs (e.g., vex2difx and difx2fits) can support project codes up to 24 characters long (was 7)

2.26.3 Bug fixes

1. difx2mark4: fix a parallactic angle calculation bug
2. vdifmux() function had some logic errors causing bad performance in gappy data; fixed.
3. In subarray cases difx2fits could provide incorrect pulse cal values; fixed.
4. mark5access bug fix to prevent crash
5. mpifxcorr: fix bug affecting data weights when nBufferedFFTs > 1 and datastreams weight < 1
6. mpifxcorr could end early due to bug in receiving a DifxParameter message; fixed.
7. some of the vdif python utilities (e.g., vdifd) had errors in parsing the command line: nbit and offset were swapped

2.26.4 Caveats

1. Outputbands support can only work on one frequency setup at a time.
2. The DiFX-2.8 series may be the last to support Mark5 recordings.

Many additional non-user-visible improvements were made to the code as well as many small user-visible improvements that do not warrant specific mention. The **ChangeLog** files that are packaged with most of the DiFX software modules contain more detailed lists of code changes. DiFX 2.8.1 was released on XXX 2023. A wide array of testing has been done and this version is considered ready to be adopted by all DiFX users.

3 Features left to implement

Here is a list of other features to add to DiFX that are not directly tied to any particular version:

1. Support for K5 format
2. Pulsar bins with proper output format
3. Space VLBI support

3.1 DiFX and AIPS

Only one task in AIPS, FITLD, has to deal with the telescope/correlator specific aspect of the FITS-IDI files that the VLBA correlator and DiFX generate. The FITS-IDI variant of FITS was first documented in AIPS Memo 102 [3], and more recently in AIPS Memo 114 [4], which will be generally available shortly. It has been modified for better support support of DiFX FITS output. In general, these changes make FITLD less telescope specific so the resulting FITS-IDI files from any DiFX installation should be highly compatible with AIPS. Several changes have been made to the 31DEC08 AIPS as a result of DiFX testing:

1. Correction for digital *saturation* in auto-correlations is disabled for DiFX FITS files. See [7] for some details on this correction which is not needed for DiFX data.
2. Support for FITS-IDI files greater than 2 GiB in size.
3. Weather table was not populated properly.
4. FITS files with multiple UV tables would generate incomplete GEODELAY columns in CL tables (not relevant to DiFX).

It is recommended that your AIPS installation be kept up to date.

With the following exceptions, data reduction of DiFX correlated data should be identical to that of VLBA hardware correlator data. This includes the continued use of DIGICOR=1 in FITLD and the use of ACCOR as you would have for the hardware correlator. The exceptions are:

1. Use of FXPOL to correct data ordering in the case of *half* polar (e.g., RR and LL products) is no longer needed.
2. Use of VBGLU to concatenate data sets in the case of 512 Mbps observations is no longer needed.
3. Data is usually combined into a single FITS-IDI file with proper calibration data attached, usually implying that TBMRG is not needed to properly concattenate calibration data. This makes DiFX FITS-IDI data similar to the *pipeline-processed* VLBA data that was made available to users of the hardware correlator with the difference being that the original FITS-IDI format is retained, keeping file sizes typically 25% smaller.

These changes should make data processing easier in almost all circumstances.

4 DiFX and pulsars

DiFX supports four pulsar processing modes:

- **Binary Gating** A simple on-off pulse accumulation window can be specified with an “on” phase and an “off” phase. This can be used to boost the signal to noise ratio of pulsar observations by a factor of typically 3 to 6 and can also be used to search for off-pulse emission.
- **Matched-filter Gating** If the pulse profile at the observation frequency is well understood and the pulse phase is very well predicted by the provided pulse ephemeris, additional signal to noise over binary gating can be attained by appropriately scaling correlation coefficients as a function of pulse phase. Depending on the pulse shape, addition gains by a factor of up to 1.5 in sensitivity over binary gating are realizable.
- **Pulsar Binning** Pulsar *binning* is supported within DiFX. This entails generating a separate visibility spectrum for each requested range of pulse phase. There are no explicit limits to the number of pulse phase bins that are supported, however, data rates can become increasingly large. Currently AIPS does not support databases with multiple phase bins. Until there is proper post-processing support for pulsar binning, a separate FITS file will be produced for each pulsar phase bin.
- **Profile Mode** Profile mode is very different than the other three as it generates autocorrelations only that are used to determine the pulse shape and phase rather than generating cross correlations. This mode is enabled by placing `mode=profile` in the global scope of the `.v2d` file (conventionally near the top). The `.v2d` file can enable as many antennas as desired (they will be averaged, so if you have a single large antenna it is probably best to include only that one), but can only operate on one source at a time. The output of `mpifxcorr` can be turned into an ASCII profile with `difx2profile`. This profile can then be given to `profile2binconfig.py` to generate the `.binconfig` file that is used by the other three pulsar modes. There is some evidence that after about 10 minutes of integration the signal to noise ratio of the resultant profile stops growing. This remains to be fully understood. It could be that increasing the integration time helps; there is no reason not to use quite large integration times in this mode.

In all cases the observer will be responsible for providing a pulsar spin ephemeris, and in all cases this ephemeris must provide an accurate description of the pulsar’s rotation over the observation duration (the pulsar phase must not drift substantially with time). If gating is to be applied then the ephemeris must be additionally be capable of predicting the absolute rotation phase of the pulsar. Enabling pulsar modes incurs a minimum correlation-time penalty of about 50%. High output data rates (computed from time resolution, number of spectral channels, and number of pulsar bins) may require greater correlator resource allocations. The details of pulsar observing, including practical details of using the pulsar modes and limitations imposed by operations, are documented at http://library.nrao.edu/public/memos/vlba/up/VLBASU_32v2.pdf.

4.1 Pulse ephemeris

The use of any pulsar mode requires a pulse ephemeris to be provided by the astronomer. This is a table of one or more polynomial entries, each of which evaluates the pulsar’s rotation phase over an interval of typically a few hours. The classic pulsar program `Tempo` can be used to produce the polynomials required [8]. The pulse phase must be evaluated at the Earth center which is usually specified in `tempo` by station code 0 (zero). Many pulsars exhibit a great degree of timing noise and hence the prediction of absolute pulse phase may require updated timing observations. When submitting the polynomial for use at the VLBA correlator, please adhere to the following naming convention: *experiment-pulsar.polyco*, e.g., `BB118A-B0950+08.polyco`. Instructions for generating the polynomial file are beyond the scope of this document.

Each `.polyco` contains one or more polynomials along with metadata; an example `.polyco` file that is known to work with DiFX is shown immediately below:

```

1913+16      6-MAY-15    90748.00    57148.38041666690      168.742789 -0.722 -6.720
  6095250832.610975    16.940537786201    0    30    15  1408.000 0.7785    3.0960
  0.18914380470191894D-06  0.26835472311898462D+00 -0.10670985785738883D-02
-0.85567503020416261D-05 -0.55633960226391698D-07 -0.37190642692987219D-09
-0.58920583351397697D-12 -0.27311855964499407D-12 -0.21723579215912174D-13
  0.11968684344685061D-14  0.92517174535020731D-16 -0.28179552068141251D-17
-0.18403230317431974D-18  0.25241984130137833D-20  0.13743173681516959D-21

```

A description of the file format is available at http://tempo.sourceforge.net/ref_man_sections/tz-polyco.txt. Currently `tempo` (version 1) is well supported and `tempo2` is only supported in `tempo1` compatibility mode. Eventual support for the `tempo2 predictors` will be added. All ephemerides must be made for the virtual Earth Center observatory (i.e., XYZ coordinates 0,0,0, usually observatory code 0; DiFX versions prior to 2.5 would not accept any non-numeric code even though they are legal). Any reference frequency can be specified as the correlator takes dispersion into consideration.

Note that although `tempo` version 2 can produce usable `.polyco` files experience has shown that version 1 has fewer failure modes.

4.2 Bin configuration file

All three pulsar modes also require the preparation of a `.binconfig` file by the astronomer. The contents of this file determine which of the three pulsar modes is being used. Three pieces of information are contained within this file: the pulsar ephemeris (`polyco`) files to apply, definitions of the pulsar bins, and a boolean flag that determines whether the bins are weighted and added within the correlator. The file consists of a set of keywords (including a colon at the end) that must be space padded to fill the first 20 columns of the file and the values to assign to these keywords that start at column 21. The file is case sensitive. The pulsar bins each consist of a ending phase and a weight; each bin is implicitly assumed to start when the previous ends and the first bin starts at the end phase of the last. The phases are represented by a value between 0 and 1 and each successive bin must have a larger ending phase than the previous. Examples for each of the three pulsar modes are shown below:

4.2.1 Binary gating

```

NUM POLYCO FILES:    1
POLYCO FILE 0:      BB118A-B0950+08.polyco
NUM PULSAR BINS:    2
SCRUNCH OUTPUT:     TRUE
BIN PHASE END 0:    0.030000
BIN WEIGHT 0:       1.0
BIN PHASE END 1:    0.990000
BIN WEIGHT 1:       0.0

```

4.2.2 Matched-filter gating

```

NUM POLYCO FILES:    1
POLYCO FILE 0:      BB118A-B0950+08.polyco
NUM PULSAR BINS:    6
SCRUNCH OUTPUT:     TRUE
BIN PHASE END 0:    0.010000
BIN WEIGHT 0:       1.0
BIN PHASE END 1:    0.030000
BIN WEIGHT 1:       0.62
BIN PHASE END 2:    0.050000
BIN WEIGHT 2:       0.21

```

```

BIN PHASE END 3:    0.950000
BIN WEIGHT 3:       0.0
BIN PHASE END 4:    0.970000
BIN WEIGHT 4:       0.12
BIN PHASE END 5:    0.990000
BIN WEIGHT 5:       0.34

```

Note here that there is zero weight given to pulse phases ranging between 0.05 and 0.95.

4.2.3 Pulsar binning

```

NUM POLYCO FILES:    1
POLYCO FILE 0:       BB118A-B0950+08.polyco
NUM PULSAR BINS:     20
SCRUNCH OUTPUT:      FALSE
BIN PHASE END 0:     0.025000
BIN WEIGHT 0:        1.0
BIN PHASE END 1:     0.075000
BIN WEIGHT 1:        1.0
BIN PHASE END 2:     0.125000
BIN WEIGHT 2:        1.0
BIN PHASE END 3:     0.175000
BIN WEIGHT 3:        1.0
.
.
.
BIN PHASE END 18:    0.925000
BIN WEIGHT 18:       1.0
BIN PHASE END 19:    0.975000
BIN WEIGHT 19:       1.0

```

The primary difference is `SCRUNCH OUTPUT: FALSE` which causes each pulsar bin to be written to disk.

4.3 Preparing correlator jobs

When using `vex2difx` to prepare correlator jobs, one must associate the pulsar with a setup of its own that includes reference to the `.binconfig` file. An excerpt from a `.v2d` file is below:

```

SETUP gateB0950+08
{
    tInt = 2.000
    nChan = 32
    doPolar = True
    binConfig = BB118A-B0950+08.binconfig
}

RULE B0950+08
{
    source = B0950+08
    setup = gateB0950+08
}

```

The `.binconfig` file should be in the same path as the `.v2d` file when running `vex2difx`.

4.4 Making FITS files

For the two gating modes, preparing FITS files with `difx2fits` is no different than for any other DiFX output. FITS-IDI does not support multiple phase bins so the pulsar binning case is different and the situation is non-optimal. Each pulsar bin must be made into its own FITS file with a separate execution of `difx2fits`. The `-B` (or `--bin`) command line option takes the bin number (starting at zero as above) and writes a FITS file containing data only associated with that bin number. Be sure to systematically name output files such that the bin number is understood.

5 Conventions

5.1 Clock offsets and rates

The clock offset (and its first derivative with respect to time, the clock rate) are stored in a number of places through the correlator toolchain. The convention used by `vex` is that the clock offset is positive if the Data Acquisition System (DAS) time tick is early (i.e., the station clock is running fast) and accordingly the full name in the `vex` file is *clock_early*. In all other parts of the system the opposite sign convention is used, that is the clock offset is how late (slow) the DAS clock is. In particular, all VLBA correlator job files, the VLBA database, the DiFX `.input` file and FITS formatted output files all use the “late” clock convention. Note in particular that the *clockOffset* and *clockRate* parameters of the ANTENNA sections in the `.v2d` files use the “late” convention, not `vex`’s “early” convention.

5.2 Geometric delays and rates

The delays used to align datastreams before correlation are nominally produced by the Goddard CALC package. The calculations are done on an antenna basis using the Earth center for antenna A and the requested station for antenna B and thus results in a negative delay for sources above the horizon. The core of DiFX uses the opposite convention and thus the delays and their time derivatives (rates) as stored in the `.delay`, `.rate` and `.im` files use the “positive delay above horizon” convention. The FITS-IDI files (as produced by `difx2fits` and the VLBA hardware correlator) use the same convention as CALC, that is “negative delay above horizon”.

5.3 Antenna coordinates

Geocentric (X, Y, Z) coordinates are used universally within the software correlator and its associated files. This usage pattern extends to cover `sched`, CALC, the VLBA database and the existing hardware correlator as well. The values are everywhere reported in meters. The Z -axis points from the Earth center to the geographic North pole. The X -axis points from the Earth center to the intersection of the Greenwich meridian and the equator (geographic longitude 0° , latitude 0°). The Y -axis is orthogonal to both, forming a right handed coordinate system; The Y -axis thus points from the Earth center to geographic longitude 90°E , latitude 0° . The unit-length basis vectors for this coordinate system are called $\hat{x}$, $\hat{y}$, and $\hat{z}$. The only exception to the above stated rules is within the AIPS software package. Task FITLD flips the sign of the Y coordinate, apparently to maintain consistency with software behavior established in the early days of VLBI.

5.4 Baseline coordinates

The baseline vectors are traditionally put in a coordinate system that is fixed to the celestial sphere rather than the Earth; see the discussion in §6.4 for a discussion of coordinates that is mathematically more precise. The unit-length basis vectors for this coordinate system are called $\hat{u}$, $\hat{v}$, and $\hat{w}$. The axes are defined so that $\hat{w}$ points in the direction of the observed source tangent point, the $\hat{u}$ is orthogonal to both the vector pointing to the celestial north pole $\hat{N}$ ($\delta_{J2000} = +90^\circ$) and $\hat{w}$, and $\hat{v}$ orthogonal to $\hat{u}$ and $\hat{w}$. Note that the sign of the

$\hat{v}$ is such that $\hat{v} \cdot \hat{N} > 0$ and $\{\hat{u}, \hat{v}, \hat{w}\}$ form a right-handed coordinate system. A baseline vector $\vec{B}_{ij}$ is defined for an ordered pair of antennas, indexed by i and j at geocentric coordinates $\vec{x}_i$ and $\vec{x}_j$. For convenience, the Earth center will be denoted by $\vec{x}_0$ which has coordinate value $\vec{0}$. The (u, v, w) coordinates for baseline vector $\vec{B}_{ij}$ is given by $(\vec{B}_{ij} \cdot \hat{u}, \vec{B}_{ij} \cdot \hat{v}, \vec{B}_{ij} \cdot \hat{w})$. There are two natural conventions for defining baseline vectors: $\vec{B}_{ij} = \vec{x}_i - \vec{x}_j$ and $\vec{B}_{ij} = -\vec{x}_i + \vec{x}_j$, which will be referred to as first-plus and second-plus respectively; both conventions are used within the correlator system. Antenna-based baseline vectors are stored in polynomial form in the `.im` file and tabulated in the `.uvw` file. In all cases antenna-based baseline vectors are baseline vectors defined as $\vec{B}_i \equiv \vec{B}_{0i} = \pm \vec{x}_0 \mp \vec{x}_i$. The values stored in the `.im` and `.uvw` files adopt the first-plus convention. For antennas on the Earth surface this implies that the w baseline component is always negative for antennas that see the target source above the horizon. The `difx` format output from `mpifxcorr` contains a baseline vector for each visibility spectrum computed from the locations of the antenna pair using the first-plus convention. Note that in this file output, the reported baseline number is $256 * i + j$ and i and j are antenna indices starting at 1. The FITS-IDI format written by `difx2fits` adheres to the second-plus convention.

5.5 Visibility phase

6 Reference guide to programs and utilities

This section has usage information for the numerous programs and scripts used in the DiFX system. Basic help information for most or all of these programs can be gotten by typing the program name with either no command line arguments or with a `-h` option, depending on the program. In the usage descriptions below, arguments in square brackets `[ ]` are optional and can often include multiple different parameters. Cases where 1 or more arguments of a certain type (such as files) can be passed to the program, the usage instructions will look like `arg1 [ ... argN]`, with the implication that N arguments of this type were passed. In cases where 0 arguments of that type is also allowed, that first argument will also be in square brackets. If it is not obvious from the program name, the software package containing the program follows the section header. The package that includes each program is included in its section heading.

Note that several VLBA specific programs are discussed in this manual that are not documented here, such as `tsm`. These are preexisting programs that may be documented elsewhere and are less likely to be useful outside VLBA operations. Also note that programs from the `nrso_difx_db` package are internal to NRAO and in general are not applicable outside VLBA operations. The code for these programs can be made available upon request.

6.1 apd2antenna (package : difx2fits)

Python program `apd2antenna` will read a `.apd` file (see Sec. 7.5), which stores baseline-based fringe-fit solutions, and writes to `stdout` an antenna-based set of fringe-fit values. These are determined through a least-squares fit. A reference antenna must be specified.

Usage: `apd2antenna apdFile refAnt`

apdFile is a `.apd` file (§7.5), created by `difx2fits`

refAnt is a 1-based number or text string indicating the reference antenna to be used

Example: `apd2antenna DQ1206.apd 2`

A reference antenna is required as the least-squares solution is ill-determined otherwise. The reference antenna has its delay, rate, and phase set to zero in the process.

The output of this program, which is sent to `stdout` is documented in Sec. 7.1. It is conventional to redirect the output of this program to a file ending in `.aapd`.

6.2 avgDiFX (package : difxio)

avgDiFX is a utility packaged with **difxio**. It reads two complete DiFX filesets that need to cover the same timerange and have essentially the same array structure and produces a new complete fileset. The **.input**, **.calc**, and **.im** files, along with any extracted pulse cal data, are copied from the first of the two datasets. The visibility data from the two file sets is averaged. Certain pulsar processing may benefit from this capability.

Usage: **avgDiFX** *configFile1* *configFile2* *outputConfigFile*

Example: **avgDiFX** **pass1_01.input** **pass2_01.input** **output_01.input**

6.3 bp2antenna (package : difx2fits)

Python program **bp2antenna** will read a **.bandpass** file (see Sec. 7.6), which stores baseline-based bandpass solutions determined by **difx2fits**, and writes to **stdout** an antenna-based set of bandpasses. These are determined through a least-squares fit separately to phase and amplitude. This program requests a reference antenna be provided and a minimum of three antennas are required. This antenna's phases are fixed to be zero and all other antennas' phases are determined relative to that antenna. If a negative number is provided for the reference antenna, the antenna phases will be adjusted after antenna-based solutions are found such that the average bandpass phase at each frequency is zero. The bandpass can be smoothed using the low-pass filter **LPF** option. The value to be provided is measured in MHz; bandpass features smaller than this are smoothed out. Optionally a pulse cal data file (determined by non-DiFX program **pcalanal**) can be applied to create "absolute phase" antenna-based bandpasses. This feature is experimental. If autocorrelation bandpasses are included in the **.bandpass** file, they will be ignored.

Usage: **bp2antenna** *bandpassFile* *refAnt* [*LPF* [*pcalFile*]]

bandpassFile is a **.bandpass** file (§7.6), created by **difx2fits**

refAnt is a 1-based number or text string indicating the reference antenna to be used

LPF is the low pass filter parameter, measured in MHz

pcalFile contains pulse cal phases after removing dominant delays (FIXME: to be documented...)

Example: **bp2antenna** **DQ1206.bandpass** **2**

The output of this program, which is sent to **stdout** is documented in Sec. 7.2. It is conventional to redirect the output of this program to a file ending in **.abp**.

6.4 calcif2

Program **calcif2** evaluates the delay model, producing a delay model file (ending with **.im**) from a file containing the source, antenna and scan timing information (ending with **.calc**). The detailed calculations are performed by the Goddard **CALC** program. Prior to this writing (May 4, 2013), the only option for the calculation back-end was **CALC** version 9 with NRAO additions which add ocean loading and near-field corrections (accurate as close as a few $\times 10^5$ km). Now new options are being introduced, including **CALC** 9 with the Sekido-Fukushima near-field model (using the **--sekido** option). Note that this option requires an installation of a special version of **CALC** that is not covered in this document.

Instead of calling **CALC** for every tabulated model row, **calcif2** computes a 5th degree polynomial every 120 seconds (typically), very closely resembling the delay model generation used at the VLBA hardware correlator. These polynomials are then evaluated at each model point. This results in a tremendous speedup

at negligible loss of accuracy. By default `calcif2` will call `CALC` three times for each model point and calculates more accurate u, v, w coordinates from delay measurements made over a small patch of the sky:

$$(u, v, w) = \left(-c \frac{d\tau}{dl}, c \frac{d\tau}{dm}, c\tau \right) \quad (1)$$

where l, m are angular coordinates (in radians) relative to the delay center on the sky, τ is the delay at the delay center and c is the speed of light.

Normally `calcif2` will be called by `difxqueue`, `startdifx`, or another higher-level program if needed.

`calcif2` connects via Remote Procedure Call (RPC) to an instance of `CalcServer` which must be running on a computer identified by environment variable `$CALC_SERVER`, or by the specified computer if the `-s` option is used. If the output files (specified in the `.calc` file) exist and are current (have newer modification times than the `.calc` file), then the files will not be recreated unless the force option is used.

In addition to calculating the delay model, this program computes the baseline vectors, u, v, w (relative to Earth center on a per-antenna basis) and source elevation vs. time.

Usage: `calcif2 [ options ] { -a | calcFile1 [ calcFile2 [ ... ] ] }`

options can be:

- `-h` or `--help` : print usage information and exit
- `-a` or `--all` : run on all `.calc` files found in the current directory
- `-v` or `--verbose` : print more verbose logging/debug info
- `-q` or `--quiet` : print less verbose logging/debug info
- `-f` or `--force` : rerun even if output files exist and are current
- `-n` or `--noaber` : don't perform aberration u, v, w corrections
- `-F` or `--fit` : Instead of producing an n term polynomial from n samples, calculate more samples and perform a fit. This is not of general use as tests have shown that the improvement is negligible.
- `-z` or `--allow-neg-delay` : don't zero delays that are negative (i.e., shadowed by Earth)
- `-A` or `--noatmos` : don't include atmosphere in calculation of u, v, w
- `-s server` or `--server server` : connect to *server*, not `$CALC_SERVER`
- `-o order` or `--order order` : make polynomials with *order*+1 terms (default 5)
- `-i int` or `--interval int` : make a polynomial every *int* seconds (default 120)
- `--override-version` ignore difx version clashes

calcFile is a `.calc` file (§7.26), such as one generated by `vex2difx` (§6.102)

Example 1: `calcif2 job1420.000.calc job1421.000.calc`

Example 2: `calcif2 -s kepler job1420.000.calc`

Example 3: `calcif2 -a -i 60`

6.5 CalcServer

Program `CalcServer` contains the Goddard Space Flight Center `CALC` package version 9.1, used to compute geometric delay models for VLBI applications. It is a repackaged version of the same source code that is used to compute models on the VLBA correlator. It is configured to run as a server. All of its interactions are via RPC calls from other programs, such as `calcif2`, which could be running on the same or different computer. This program only needs to be started once on a given machine using the `startCalcServer` script. It should probably be set to start automatically upon boot of the machine on which `CalcServer` runs. Environment variable `$CALC_SERVER` should be set to the name of the computer on which `CalcServer` is running.

```
Start: startCalcServer

Test: checkCalcServer $CALC.SERVER

Stop: killall CalcServer
```

Note that `CalcServer` must be installed (with `make install`) to be usable as the paths for various files are permanently set in the executables at compile time. At this time it seems `CalcServer` cannot be compiled for 64-bit machines.

6.6 `checkdir` (package : mk5daemon)

Program `checkdir` can be used to check the integrity of one or more `.dir` files that are stored at a location pointed by environment variable `MARK5_DIR_PATH`. Even after many years of use, the Mark5 units tend to be a weak point in the reliability of correlation. Since reading the module directory and examining a bit of data from each scan are the first actions done to a module, many of the possible problems show up at this time. This utility looks for a number of possible problems, including scans that could not be decoded, overlapping or out-of-order scans, scans with illegal format parameters and others. This program makes no attempt to fix problems. It is up to the operator to determine if a problem is real or not and if further action should be taken. In cases where many scans are not properly decoded it is worthwhile to rename (or remove) the `.dir` file in question and regenerate the directory. A second directory read often succeeds when a first one does not.

Usage: `checkdir` [*options*] [*module list*]

options can be:

- h or --help : print usage information and exit
- v or --verbose : be more verbose in execution (-v -v for more)
- q or --quite : be less verbose in execution
- a or --all : run on all files in `$MARK5_DIR_PATH`
- s or --show : print the entire directory file to screen
- H or --histogram : print a histogram of record rates

Example 1: `checkdir -a`

Example 2: `checkdir NRA0-123 NRA0+266`

Example 3: `checkdir -s NRA0+233`

Either `-a` or a list of module names can be provided (but not both simultaneously). If the former, a less verbose output will be generated by default. Except in the lowest verbosity mode (the default for `-a`), module directories without any detected problems will show a one line summary consisting of the number of scans and the time range of the module.

6.7 `checkmpifxcorr` (package : mpifxcorr)

Program `checkmpifxcorr` reads the `.input` and other associated files for a DiFX job and parses them with the same logic used by `mpifxcorr` in order to determine their validity.

Usage: `checkmpifxcorr` [*options*] *configFile* ...

options can be:

-h : print usage information and exit
-f -s -e -w -i -v -d : select the verbosity level of output (options refer to “fatal”, “severe”, “error”, “warning”, “info”, “verbose”, and “debug” levels).

All of the files referenced from the provided configuration (**.input**) files are read as well (excepting any baseband files or the **.vex** file). This check has proven especially useful for pulsar processing. The default verbosity level will lead to printing of any problems at the “warning” level or worse. See Sec. 8.1 for details on the severity levels.

6.8 **cleanVDIF** (package : vdifio)

Program **cleanVDIF** loops through a VDIF file writing valid content to a new output file.

Usage: **cleanVDIF** *inputvdiffile* *outputvdiffile* *Mbps* [*options*]

inputvdiffile is the recorded VDIF file to clean

outputvdiffile is the corrected VDIF file to write

Mbps is the data rate in megabits/second

options can be:

-v or **--verbose** : be verbose in execution

Example: **cleanVDIF** bad.vdif good.vdif 256

6.9 **condition** (package : nrao-difx-db)

This is an NRAO-only program owing to its ties to the VLBA database.

Program **condition** is mainly used to extract Mark5 module conditioning reports from the database but also has the means to manually import data into the database. When querying (with the **find** action), one or more “identifiers” can be supplied which can be either the names of the Mark5 modules or serial number of individual disks (or a mix of the two!). Environment variable **VLBA_DB** must be set to point to the correct database.

Usage: **condition** [*options*] *action* + *args*

options can be:

-h or **--help** : print usage information and exit

-v or **--verbose** : be verbose in execution

action can be one of:

add *report1* [*report2* ...]

find *identifier1* [*identifier2* ...]

report is the name of a file containing one or more condition reports from **SSErase**

identifier is either a Mark5 module VSN or a hard disk serial number

Example 1: **condition** add NRAO-040

Example 2: **condition** find NRAO-042

Example 3: **condition** find NRAO+342 NRAO+270

Example 4: **condition** find Y66M3BQE

6.10 **condition_watch** (package : nrao.difx.db)

This is an NRAO-only program owing to its ties to the VLBA database.

Program **condition_watch** is meant to run as a background process on the correlator head node. Its function is to receive **Mark5ConditionMessages** emitted by a special version of **SSErase** (the module conditioning program) and stuff this data into the database. This program is automatically started by **mk5daemon** when it is supplied with the **-w** or **--condition-watch** arguments. When restarting **mk5daemon** by hand, make sure that a duplicate copy of **condition_watch** is not left running. Environment variable **VLBA.DB** must be set to point to the correct database.

Usage: **condition_watch** [*options*]

options can be:

-h or **--help** : print usage information and exit

Example: **condition_watch**

6.11 **countVDIFpackets** (package : vdifio)

Program **countVDIFpacket** loops through a VDIF file and counts number of valid and skipped frames. Packet counts are performed only on the thread ID requested.

Usage: **countVDIFpackets** *vdiffile Mbps threadId*

vdiffile is the recorded VDIF file

Mbps is the data rate in megabits/second

threadId is the threadId to report on

Example: **countVDIFpackets** *example.vdif* 256 3

6.12 **cpumon** (package : difxmessage)

Program **cpumon** is a program that listens for **difxLoad** messages multicast from the Mark5 units and displays the information; updating the display as new messages are received.

Usage: **cpumon**

Make sure the terminal is at least 60 characters wide and is at least as tall as there are computers that may transmit information. To quit, use ctrl-C. The columns displayed are:

1. Computer name
2. CPU load averaged over 10 seconds
3. Memory usage / Total memory
4. Network receive rate (Mbps)
5. Network transmit rate (Mbps)
6. Number of CPU cores

6.13 diffDiFX.py (package : vis2screen)

Program `diffDiFX.py` generates a context-sensitive difference of two DiFX output files for detailed version testing. Corresponding visibility records are differenced and statistics on the differences are accumulated and printed at the end of the processing.

Usage: `diffDiFX.py [ options ] { difxfile1 difxfile2 }`

options can be:

- `-h` or `--help` : print usage information and exit
- `-f` *FREQ* or `--freq=FREQ` : Only look at visibilities from this *FREQ* index
- `-b` *BASELINE* or `--baseline=BASELINE` : Only look at visibilities from this *BASELINE* num
- `-t` *THRESHOLD* or `--threshold=THRESHOLD` : Display any difference that exceeds *THRESHOLD*
- `-e` *EPSILON* or `--epsilon=EPSILON` : Display any differences that exceeds allowed numerical error *EPSILON*
- `-s` *SKIPRECORDS* or `--skiprecords=SKIPRECORDS` : Skip *SKIPRECORDS* records before starting comparison
- `-m` *MAXRECORDS* or `--maxrecords=MAXRECORDS` : Stop after comparing *MAXRECORDS* (if $\neq 0$) records
- `-p` *PRINTINTERVAL* or `--printinterval=PRINTINTERVAL` : Print a summary every *PRINTINTERVAL* records
- `-c` *MAXCHANNELS* or `--maxchannels=MAXCHANNELS` : The length of the array that will be allocated to hold vis results
- `-v` or `--verbose` : Turn verbose printing on
- `-i` *INPUTFILE* or `--inputfile=INPUTFILE` : Parse *INPUTFILE* for the correlation setup
- `--matchheaders` : On seeing a header mismatch, skip through file 2 looking for next match

difxfile1 is the first difx file to compare

difxfile2 is the second difx file to compare

Example: `diffDiFX.py -i example_1.input example_1.difx/DIFX_55523_025239.s0000.b0000 comparison_1.difx/DIFX_55523_025239.s0000.b0000`

If the error for any record exceeds the specified threshold a verbose error message is printed. Summary statistics are printed at the end of the file. Warnings are printed if the headers do not match between the two files.

6.14 difx2fits

Program `difx2fits` creates a FITS output file from the native output format created by `mpifxcorr` and several other files carrying information about the observation. Multiple input file sets can be specified. A separate output FITS file is created for each unique frequency setup encountered. When run, `difx2fits` requires the following files to be present for each DiFX file set being converted:

1. *baseFilename.difx/*
2. *baseFilename.input*
3. *baseFilename.calc*

4. *baseFilename.im*

Several other files are optional and are typically used to populate calibration and ancillary tables:

1. *baseFilename.flag*
2. *flags*
3. *pcal*
4. *tsys*
5. *weather*
6. *\$GAIN_CURVE_PATH/*
7. *.difx/*.history*

With the exception of the gain curve files, all the input files to `difx2fits` are expected to be in the current working directory or in the place indicated by the `.input` file. As the visibility file (`.difx`) is read, any records that are all zero are omitted.

Usage: `difx2fits [ options ] { -d | baseFilename1 [...] baseFilenameN } [outFile]`

options can be:

- `-h` or `--help` : print usage information and exit
- `-n` or `--no-model` : don't write model (ML) table
- `-s scale` or `--scale scale` : scale visibility data by *scale*
- `-t interval` or `--deltat interval` : generate `.jobmatrix` file with time intervals of length *interval* seconds
- `--difx-tsys-interval interval` : the Difx-derived tsys interval (sec) (default 30.0s averaging)
- `--difx-pcal-interval interval` : the Difx-derived pcal interval (sec) (default 30.0s averaging)
- `-S` or `--sniff-all` : sniff all bins and phase centers, not just the first
- `-T interval` or `--sniff-time interval` : use *interval* as the sniffer time resolution
- `-v` or `--verbose` : increase verbosity of output; use twice or thrice to get even more
- `-d` or `--difx` : run on all `.difx` files found in the directory
- `-k` or `--keep-order` : don't sort the antennas by name
- `-1` or `--dont-combine` : make a separate FITS file for each input job
- `-x` or `--dont-sniff` : don't generate sniffer output files
- `-0` or `--zero` : don't put visibility data in FITS file
- `--bin b` : Select on this pulsar bin number
- `--phasecentre p` : (U.S. spelling okay too) Create a FITS file for all the p^{th} phase centers (default 0)
- `--override-version` : ignore difx version clashes
- `--bandpass` : write the `.bandpass` file (see Sec. 7.6)
- `-m nJob` or `--max-jobs nJob` : split into more FITS files after reaching *nJob* input files.
- `--eop-merge-mode mode` : sets conditions for allowing jobs with different EOPs to be merged or not; options are `strict` (default), `drop`, `relaxed`

`--clock-merge-mode mode` : sets conditions for allowing jobs with different clock models to be merged or not; options are `strict` (default) or `drop`

`--antpol` : use antenna-based polarization labels as in VEX. Note: fits-idi file will violate original specifications and abide extended specifications.

`--polxy2hv` : re-labels all polarizations XY to HV. Requires `--antpol` option.

baseFilename is the prefix of the jobfile to convert; it is okay to use the `.difx` filename instead

outFile is the name of the FITS file to produce; if not provided one will be made based on the project code

Example 1: `difx2fits dq109_1 DQ109.FITS`

Example 2: `difx2fits -v -v -d`

Environment variables respected:

- `DIFX_GROUP_ID` : if set, run `difx2fits` with `umask(2)`.
- `DIFX_LABEL` : the local name of the difx install. Used to verify matching versions and put inside FITS file; if not set, `DIFX_VERSION` will be used instead.
- `DIFX_MAX_SNIFFER_MEMORY` : maximum amount of memory (bytes) to allow sniffer to use.
- `DIFX_VERSION` : the difxbld version name.
- `GAIN_CURVE_PATH` : directory containing gain files.
- `TCAL_PATH` : a directory containing Tcal value files.
- `TCAL_FILE` : if `TCAL_PATH` is not set, use the file pointed to by this env. var.

Unless adjusted with the `--difx-pcal-interval` and `--difx-tsys-interval` parameters, the respective PC and TY data will be time averaged to the default of 30 seconds. For geodetic VLBI or other observations with very short scans you may want to shorten the averaging time.

Unless disabled with the `--dont-sniff` or `-x` flag, four “sniffer” output files (`.acb`, `.apd`, `.wts` and `.xcb`) will be written for each `.FITS` file produced. These files are used by `difxsniff` and its associated programs to produce data plots that are used to assess data quality.

Unless disabled by setting *interval* to a non-positive number with the `-t` or `--deltat` option, an output file with suffix `.jobmatrix` will be produced. This file contains an ASCII art diagram of which jobs contributed to each `.FITS` file produced as a function of both time and antenna.

Unless augmented with option `--antpol` the produced FITS files are fully compliant with the original FITS-IDI specifications. In case of mixed mode polarization (XY against RL) or certain feed types (HV in particular) the option `--antpol` allows to force output of noncompliant but polarization correct FITS files; FITS-IDI numerical parameter `STK_1` is set to the new value of -9, indicating to post-processing software that it should refer to the existing FITS-IDI character parameters `Antenna1Feed1` and `Antenna2Feed1` (populated from VEX) for the polarization details.

If submitting a bug report for `difx2fits`, please include in it the full output of `difx2fits -v -v` and the `.input` and `.calc` files.

`difx2fits` displays several diagnostics during the conversion process, separately for each output FITS file. The size of each FITS table is printed; a zero size indicates that table is not produced. For the visibility table, input files contributing to the output are printed. Scan information is printed at increased verbosity levels. Not all DiFX output visibilities are written to the FITS file. Accounting of the disposition of the visibilities is provided. Invalid records are those containing infinite or NaN values and indicate a likely bug in the software. Flagged records are those identified by `vex2script` (or other DiFX file set generation programs)

in the `.flag` file as being illogical, such as cases where a particular baseline during a job belongs to a different subarray. When using integration times longer than 1 second it is possible for one visibility to span two scans. Such records are dropped. Finally any visibilities, produced outside a normal scan start/stop time are dropped; this should not occur unless the `.calc` file is modified between correlation and FITS creation.

If there are any files matching `.difx/*.history` for `.difx/` output being converted to `.FITS`, the contents of these files will be inserted into the FITS HISTORY table.

6.15 difx2mark4

Program `difx2mark4` creates a Mark4 output file set from `mpifxcorr` input and output files. When run, `difx2mark4` requires the following files to be present for each file set being converted:

1. `baseFilename.difx/`
2. `baseFilename.input`
3. `baseFilename.im`

as well as the `.vex` file referenced in the `.input` file, which may be common to many DiFX file sets.

Usage: `difx2mark4 [ options ] baseFilename1 [...baseFilenameN]`

options can be:

- `-h` or `--help` : print usage information and exit
- `-v` or `--verbose` : increase verbosity of output; use twice or thrice to get even more
- `-d` or `--difx` : run on all `.difx` files found in the directory
- `-k` or `--keep-order` : don't sort the antennas by name
- `--override-version` : ignore difx version clashes
- `-r` or `--raw` : suppresses normalization of amplitudes
- `-p` or `--pretend` : do a dry run
- `-e` or `--experiment-number n` : set the experiment number to *n* which must be a 4 digit number (default is 1234)
- `-b code flo fhi` : Override freq band codes. Frequencies are in MHz. Multiple parameters of this kind can be specified.

baseFilename is the prefix of the jobfile to convert *without* the underscore and job number

Example `difx2mark4 dq109`

6.16 difxarch (package : nrao_difx_db)

Program `difxarch` is a simple script that moves FITS files produced by `makefits` from the correlation queue staging area (defined by the `DIFX_QUEUE_BASE` environment variable) to the archive staging area (defined by environment variable `DIFX_ARCHIVE_ROOT`). A process running on the archive computer will periodically monitor new files in this staging area and will then copy them to the actual archive. In order to prevent premature pick-up of these files, they are first moved into a directory with a name beginning with a period (`.`). This directory is renamed without the period once all files to be archived are copied.

Usage: `difxarch [ options ] passName1 [...passNameN]`

options can be:

- h or --help : print usage information and exit
- v or --verbose : increase verbosity of output
- p or --pretend : generate SQL and bash commands but don't execute them
- override-version ignore DiFX version clashes

passName is the name of a correlator pass; a file called *passName.fitslist* is expected to be present

Example: `difxarch -v clock`

6.17 difxbuild

Program `difxbuild` aids in the installation of DiFX onto a cluster. Full documentation on the install process can be found in §??, so details will not be shown here. Command syntax is as follows:

Usage: `difxbuild [ options ] command [ command arguments ]`

options can be:

- h or --help : print usage information and exit
- d or --documentation : print full in-line documentation to screen
- t or --todo : print developer's to-do list
- v or --verbose : increase verbosity of output
- q or --quiet : decrease output verbosity
- p or --pretend : generate SQL and bash commands but don't execute them
- V or --version : print version and quit

command is one of the `difxbuild` commands, such as `build` or `svn`; the program help information will list all options

command arguments are options for some commands

6.18 difxcalc11

6.19 difxcalculator (package : difxio)

Program `difxcalculator` looks at a set of DiFX input files (`.input`, `.calc`, etc.) and reports/calculates key operating parameters. This program is inspired by the `difx_calculator.xls` spread sheet available at <http://www.atnf.csiro.au/vlbi/dokuwiki/doku.php/difx/calculator>.

Usage: `difxcalculator [ options ] baseName [speedUp]`

options can be:

- h or --help : print usage information and exit

baseName is the base name of a correlator job

speedUp is the expected processing rate relative to real-time

Example: `difxcalculator mt933_01`

Known bugs:

1. Does not take into consideration multiple phase centers or zoom bands.

6.20 difxclean (package : nrao.difx_db)

Program **difxclean** simply deletes all data from `$DIFX_QUEUE/` for a particular project. It also removes all jobs with status not equal to `COMPLETE` for this project from the `DIFXQUEUE` table of the database. It is intended that it be run at the same time the project is released, meaning data has been correlated and archived. The user does not have to be in any particular directory when running this program.

Usage: **difxclean** [*options*] *project*

options can be:

- h or --help : print usage information and exit
- p or --pretend : don't actually do the erasure

project is the name of the project to be “cleaned” out

Example: **difxclean** mt917

6.21 difxcopy (package : misc_utils)

Python program **difxcopy** is used to copy DiFX input (and other) files to a different directory. In the process, explicit references to other files that are being copied are changed to reflect their new file system path. For a given file prefix, *prefix*, the following files are copied if they exist: *prefix.input*, *prefix.calc*, *prefix.flag*, *prefix.delay*, *prefix.uvw*, *prefix.rate*, and *prefix.im*. The *.vex* file referenced within the *.calc* file is also copied.

Usage: **difxcopy** [*options*] *jobPrefix1* [...*jobPrefixN*] *destDir*

options can be:

- h or --help : print usage information and exit
- v or --verbose : possibly increase verbosity of output

jobPrefixN is the file prefix of a job, e.g., `mt911.04` would be the prefix corresponding to input file `mt911.04.input`

destDir is the directory in which the copied and modified files will be placed

Example: **difxcopy** mt911.02 mt911.03 mt911.04 /home/difx/queue/MT911

6.22 difxdynamicmon (package : difxmessage)

Program **difxdynamicmon** listens for multicast messages of the `difxStatus` variety and simply prints their contents to the terminal. This is mainly useful for debugging `mpifxcorr`. Diagnostic information that is produced includes status of internal buffers, memory usage, execution time, data throughput and lost subintegrations.

Usage: **difxdynamicmon** [*options*]

options can be:

- h or --help : print usage information and exit

6.23 difxlog (package : difxmessage)

Program **difxlog** can be used to collect DiFX multicast messages for a particular correlator job and write them to a file. A new instance of **difxlog** must be started for each job being run. Normally this will be done automatically if implemented in the particular deployment of DiFX. Both **startdifx** and **mk5daemon** can start DiFX and will instantiate a **difxlog** process as needed. Only messages of type **DifxAlertMessage** and **DifxStatusMessage** are collected and written to the output file.

Usage: **difxlog** *jobIdentity* *outFile* [*logLevel* *pidWatch*]

jobIdentity : the name of the job being run (specifically, it should match the **identifier** field in the **DifxMessage** being sent).

outFile : the name of the output file containing log information.

logLevel : the minimim message severity to retain (see §8.1).

pidWatch : the program id (in the Unix sense) of the mpifxcorr process running.

Example: **difxlog** mt911.04 mt911.04.difxlog 4 1243

Unless a *pidWatch* value is specified, **difxlog** will run until killed. If a *pidWatch* value is provided, **difxlog** will quit as soon as that process stops running. The *loglevel* parameter can be used to select the maximum severity level to write to the log. The possible values and their meanings are:

0	Fatal	mpifxcorr cannot continue because of the noted problem
1	Severe	an internal error that should never happen happened (likely bug)
2	Error	a problem was encountered in the data processing
3	Warning	something suboptimal was noted
4	Informative	a note containing progress information
5	Verbose	more detailed progress information
6	Debug	values probably of use only to software developers

6.24 difxqueue (package : nrao_difx.db)

This is an NRAO-only program owing to its ties to the VLBA database.

Python program **difxqueue** is a program used to maintain the DiFX correlator queue. There are two main responsibilities in doing so: copying or deleting files in the correlator queue directory (which is project specific: **\$DIFX_QUEUE_BASE/projectName**) and maintaining the database entries for each queued job. In the VLBA context, this program is the main interface between the analysts and the correlator operators. This program is mainly intended to work on one *job pass* at a time rather than single jobs or whole projects. In some cases one job pass could be one job, or it could be a whole project (or both), but in many cases there will be multiple passes per project with possibly multiple jobs per pass. It is possible for **difxqueue** to operate on individual jobs when a list of job numbers is provided. The first command line argument describes the action to perform. Each subsequent argument is then context dependent; see the examples or run with the **-h** command line parameter to get a feel for the variety of options allowed. Once a job has been correlated successfully, its status will be **COMPLETE**. There is no need to **delete** a job from the queue once it is complete. Doing so will require recorrelation if the results of that job are still needed. Each job in the queue has a priority. The smaller the priority, the lower the number. By default a queued job will have priority 2.

Usage: **difxqueue** [*options*] *action* [*args*]

options can be:

-h or **--help** : print usage information and exit

-p *priority* or **--priority** *priority* : set the priority of jobs to *priority*
-q *queuedir* or **--queuedir** *queuedir* : manually set the staging directory
-v or **--verbose** : increase verbosity of output
-d or **--db-only** : do not copy/delete/move files; operate only on database
--override-version : ignore DiFX version clashes

action : the action to perform

add : add job(s) to the queue, usually a whole pass at a time. Default priority is 2; use the **-p** option to set the priority if a different priority is required.

Example 1: `difxqueue add clock`

Example 2: `difxqueue add mt911 1 2 3 4`

Example 3: `difxqueue -p 3 add geodesy`

bump : increase the priority of queued job(s)

Example: `difxqueue bump clock`

del : remove job(s) from the queue

Example 1: `difxqueue del clock`

Example 2: `difxqueue del mt911 2 3`

list : list all jobs within a pass

Example: `difxqueue list mt911`

listall : list all incomplete jobs in the queue; note that this is not restricted even to any particular project. If one or more projects is specified, all jobs, complete or not, for those projects will be listed. If no segment code is appended to a project name, then all matching proposal codes will be listed.

Example 1: `difxqueue listall`

Example 2: `difxqueue listall BX123 BY321`

Example 3: `difxqueue listall BR138A`

log : list all correlations that have happened for a given project. This simply searches the DIFXLOG database table and dumps it to the screen in a readable fashion.

Example 1: `difxqueue log BX123`

prod : print production queue list, possibly sending to a file

Example 1: `difxqueue prod`

Example 2: `difxqueue prod queue.txt`

set : set the status of queued job(s)

Example 1: `difxqueue set tc015d COMPLETE`

Example 2: `difxqueue set tc015d QUEUED 3 4`

slide : decrease the priority of queued job(s)

Example: `difxqueue slide mt911 6`

args : *action* dependent arguments, usually a pass name and possible list of job numbers

Note that except for the **listall** and **prod** actions, the current working directory must contain the `.joblist` file for a project.

6.25 difxsniff (package : SniffPlots)

Program `difxsniff` is a reimplementaion of the VLBA analysts' program `sniff.pd` to be more appropriate for software correlation where the sniffer data is generated at the same time as the FITS files. It uses the same underlying set of plotting programs (`plotwt`, `plotbp`, and `plotapd`) as `sniff.pd` did. It should be run in a project directory as it will create a subdirectory (if not existing already) which by default is called `sniffer/refant` within the current directory. All files created by `difxsniff` will be placed in this directory, overwriting existing files with the same filenames. Unlike `sniff.pd`, `difxsniff` is a purely non-interactive command line program. Note that although `.FITS` files are provided to `difxsniff`, it is the associated files ending in `.apd`, `.wts`, `.acb` and `.xcb` that are actually read.

Usage: `difxsniff [ options ] refants FITS1 [ ... FITSN ]`

options can be:

`-h` or `--help` : print usage information and exit

refants is a list reference antennas, separated by spaces

FITS is a FITS file created by `difx2fits`; multiple FITS files can be specified together

Example 1: `difxsniff LA *.FITS`

Example 2: `difxsniff NL FD *.FITS`

6.26 difxspeed (package : vex2difx)

Program `difxspeed` does processing benchmarking, possibly over a range of parameters, of DiFX. To ensure that data playback (reading from files, Mark5 modules or network) are not limiting performance, the FAKE mode of DiFX (see §??) is used; thus the output data are meaningless. `difxspeed` takes a `.speed` (§7.14) file as input. This file contains various parameters, many of which are identical to those in the `.v2d` (§7.42) files. An important difference with the parameters specified in `.speed` files is that multiple values can be provided for many of the parameters. In the benchmarking process, a separate run of DiFX for each combination of the supplied parameters is performed. The first combination is run twice, with the first being labeled a *dummy* run. This is because the timing of the first execution can vary depending on recent usage of the correlator.

Usage: `difxspeed [ options ] inputFile [ numIterations ]`

options can be:

`-h` or `--help` : print usage information and exit

inputFile is the `.speed` file describing the series of benchmarks to run

numIterations is the number of times to execute all test combinations

Each run of `difxspeed` will append a new column of data to a file called `inputFile.out` ; if the file does not exist, a new file will be created. Documentation of this output file format can be found in §7.15.

6.27 difxusage (package : nrao_difx_db)

This is an NRAO-only program owing to its ties to the VLBA database.

Program `difxusage` mines the VLBA database for correlator usage statistics. Usage is as follows:

Usage: `difxusage [ options ] mjdStart mjdStop`

options can be:

- h or --help : print usage information and exit
- v or --verbose : be more verbose in execution
- l or --list : print all matching jobs
- a or --all : select jobs in all states
- c or --complete : select only complete jobs (default)
- k or --killed : select only killed jobs
- u or --unknown : select only unknown jobs

mjdStart is the start time (in Modified Julian Days) to look for jobs

mjdStop is the stop time (in Modified Julian Days) to look for jobs

Note that environment variable VLBA_DB must be set to point to the postgres database in question.

6.28 difxvmf (package : calcif2)

Note: this is coming in DiFX 2.7 series...

difxvmf takes a DiFX fileset (including the **.im** file) and modifies the wet and dry troposphere values based on the Vienna Mapping Functions. This program retrieves the needed external data from <http://vmf.geo.tuwien.ac.at>. The **.im** file will be replaced with an updated version.

Usage: **difxvmf** [*options*] *filebase1* [*filebase2* ...]

options can be:

- h or --help : print usage information and exit
- v or --verbose : be more verbose in execution
- w or --usewx : use metrology data from each site rather than defaults

filebasen is the **.input** file or prefix to be processed; multiple can be provided.

If **--usewx** is specified, files of the form *project.station.weather* will be looked for in the local directory and used to supply metrology data, overriding defaults.

Environment variables:

DIFX_VMF_DATA : a writable directory for caching downloaded VMF data

DIFX_VERSION : to enforce DiFX version compatibility

6.29 difxwatch (package : difxmessage)

Program **difxwatch** can be used to monitor progress of ongoing DiFX jobs and kill jobs that appear to be hung.

Usage: **difxwatch** [*options*]

options can be:

- h or --help : print usage information and exit
- version : show program's version number and exit
- i *idletime* or --idle-time *idletime* : maximum number of seconds a job is allowed to be idle before it is killed.

6.30 DiFX Operator Interface

The DiFX Operator Interface (DOI) is a java-based application to monitor and control the correlation of DiFX jobs. Jobs to be correlated can be selected with a file browser or retrieved with a database request. A separate manual [?] will be made available with instructions for its use. The only specific detail that will be mentioned here is the contorted path the starting of a job takes:

1. The job to run is selected.
2. The DOI determines which resources (Mark5 units and processor nodes) are required.
3. If the intended output file already exists, a dialog will ask the operator whether to overwrite this file or not.
4. The DOI allocates resources.
5. The DOI assembles a `DifxStartMessage` XML document and multicasts it with the correlator head node as the recipient.
6. The `mk5daemon` process running on the head node captures this message.
7. `mk5daemon` fork(s); the child process changes its `userId` to `difx` and spawns an `mpirun` process via `ssh` to ensure the proper environment variables are set.
8. The `mpirun` process starts a copy of `mpifxcorr` on each of the Mark5 units and processing nodes that is requested.
9. `mk5daemon` fork(s) again; the child process changes its `userId` to `difx` and spawns a `difxlog` process via `ssh` to ensure the proper environment variables are set.
10. All processes continue until job end is reached or the job is killed.
11. When the first fork(ed) `mk5daemon` process ends, the `difxlog` process stops automatically, causing the second fork(ed) process also to stop.
12. The DOI receives messages suggesting the job has ended and frees the allocated resources.

6.31 e2ecopy (package : nrao.difx.db)

Program `e2ecopy` copies files one directory to another, changing the ownership to `$DIFX_ARCH_USERNAME` in the process. This program must be setuid root; the person installing the program must run `chmod +s e2ecopy` after installation if `make install` is not run by root. Normally this program is run by `difxarch` (see §6.16). This is a VLBA-centric program, but could be used by others.

Usage: `e2ecopy [ options ] fromDir toDir file1 [ ... fileN ]`

options can be:

`-h` or `--help` : print usage information and exit
`-v` or `--verbose` : be more verbose

fromDir : the source directory of the file(s) to copy

toDir : the destination directory

file : a file to copy (multiple files may be provided)

Note: “e2e” is NRAO terminology for “End to End”, a philosophy of providing user software covering the full project lifecycle from proposal handling to archive access. In this particular case the name arose due to the location of the archive staging area at NRAO.

6.32 **errormon** (package : difxmessage)

Program **errormon** listens for multicast messages of the **difxError** variety and simply prints their contents to the terminal. It is effectively the same as **difxlog** except that log data is sent to *stdout* rather than a systematically named file.

Usage: **errormon** [options] [*maxSeverity*]

options can be:

-h or **--help** : print usage information and exit

maxSeverity : maximum severity level to display (default = 8)

See § 8.1 for a list of severity codes. If no *maxSeverity* is provided, the default level of 8 will cause no selection to occur; all messages will be printed.

A similar program, **errormon2**, does nearly the same thing, but defaults to a less verbose output, and sends output to *stderr* rather than *stdout* (so use with **grep** or other *nix tools is more cumbersome. It also writes its output to a log file.

See documentation for **difxlog** to see the list of alert levels.

6.33 **extractSingleVDIFThread** (package : vdifio)

This program has been superceded by **vmux** (see Sec. 6.105).

6.34 **extractVDIFThreads** (package : vdifio)

This program has been superceded by **vmux** (see Sec. 6.105).

6.35 **fakemultiVDIF** (package : vdifio)

Note: the input file for **fakemultiVDIF** must have a single thread or unpredictable results will occur.

6.36 **fileto5c** (package : mark5daemon)

6.37 **filterVDIF** (package : vdifio)

6.38 **generateVDIF** (package : vdifio)

6.39 **genmachines** (package : mpifxcorr)

Program **genmachines** uses the information in a **.input** file and a file containing information about the members of the compute cluster (such as the file pointed to by **\$DIFX_MACHINES**) to produce a **.machines** file (§7.31) needed by **mpifxcorr**. Note that **genmachines** is not intended to be run by hand anymore as **startdifx** does this, if necessary. If playback directly off Mark5 units is to be done, **genmachines** will send a multicast request to all Mark5 units on the correlator requesting an inventory of loaded Mark5 modules. The **mk5daemon** process on each unit will respond with another multicast message containing the loaded modules and the status of the unit, i.e., whether busy or available to be used. This information is collected by **genmachines** which will look for availability of all the modules and detect conflicts (i.e., two needed modules loaded in the same unit). If all needed modules are found and enough resources remain for the computations, a **.machines** file and a **.threads** file are written. Note that the **.machines** file contains a certain number of comment lines so that the use of Unix command **wc -l** can be used to determine exactly how many processes will be started. It is suggested to run this program immediately before starting the software correlator to minimize the chance that the Mark5 units change their status or that information about the modules whereabouts becomes stale; it is thus discouraged to run with ***.input**.

Usage: **genmachines** [*options*] *input1* [$\cdots$ *inputN*]

options can be:

- h or --help : print usage information and exit
- v or --verbose : be more verbose
- o or --overheadcores *ohc* : leave at least *ohc* on each compute node unscheduled
- m *file* or --machinesfile *file* : use *file* instead of \$DIFX_MACHINES
- n or --no-threads : don't write a .threads file.
- d or --difxdb : lookup module locations in a database.

input is a .input file; multiple files can be specified, each producing its own .machinesfile

6.40 **getshelf** (package : nrao-difx-db)

This is an NRAO-only program owing to its ties to the VLBA database.

Program **getshelf** retrieves the shelf location of specified modules from the legacy VLBA database and prints them to the screen. While possibly useful, this program is not required for the software correlation process.

Usage 1: **getshelf** [*options*] *module1* [*module2* [$\cdots$]]

Usage 2: **getshelf** [*options*] *shelfFile*

options can be:

- h or --help : print usage information and exit
- v or --verbose : print the database query string as well

moduleN is the volume serial number (VSN) of a module; multiple modules can be specified

shelfFile is a .shelf file (§7.37), as may be written by **db2vex** Default is the current working directory if none is provided.

Example 1: **getshelf** NRA0+267

Example 2: **getshelf** bx123a.skd.shelf

6.41 **jobdisks** (package : mpifxcorr)

Program **jobdisks** looks through job files to see which modules (disks) are needed for correlation. It reads through .input files, as used by **mpifxcorr**, to get the needed information. There are two modes of operation. By default, a matrix of all modules for all stations is displayed, with a -- symbol indicating that a particular station is not used in a particular job. An asterisk (*) indicates a module change. The second mode, instigated with command line argument -c, summarizes only module changes. Running without any arguments will cause **jobdisks** to look at job files within the current directory, prioritizing on .input files if any exist and falling back on .fx files otherwise. Listings for a subset of jobs can be made by specifying particular files.

Usage: **jobdisks** [*options*] [*file1*] $\cdots$ [*fileN*]

options can be:

- h or --help : print usage information and exit

`-c` or `--changes` : print module changes only

file is a `.fx` or `.input` file; mixed types are not supported. Multiple input files may be supplied.

Example 1: `jobdisks`

Example 2: `jobdisks job1420*.input`

Example 3: `jobdisks *.fx`

Example 4: `jobdisks -c`

Known bugs:

1. Program should make sure datastream type is MODULE.

6.42 **joblist** (package : mpifxcorr)

Program `joblist` prints useful information about DiFX correlator jobs to *stdout*. Six columns of output are produced:

1. Job file base filename
2. File indicator, showing a particular character for each one of the files associated with that job that is found within a pair of square brackets, []:

`c` `.calc` file (§7.11)

`m` `.machines` file (§7.31)

`t` `.threads` file (§7.38)

`i` `.im` file (§7.25)

`v` `.difx` file (§7.12)

3. Band code of first scan in file
4. Observation duration of correlation (in minutes)
5. Recording mode triplet; three integers(data rate(Mbps), number of baseband channels & quantization bits) separated by dashes
6. Comma separated list of antennas

One line is printed for each `.input` file found in the list of directories provided (or current directory if not listed).

Usage: `joblist [ options ] [ dir1 ] ... [ dirN ]`

options can be:

`-h` or `--help` : print usage information and exit

dir is a directory for which to print job information (default is current shell directory). Multiple directories can be specified.

Example 1: `joblist`

Example 2: `joblist $JOB_ROOT/*`

6.43 `jobstatus` (package : mpifxcorr)

Warning: As of DiFX 2.0.1, this utility has not yet been updated to work with DiFX 2 output format.

Program `jobstatus` lists the current correlation progress for each DiFX job in one or more directories. This program is normally run without any command line arguments from within the project directory. For each job, the base filename is listed with 5 or 6 additional columns of data. These columns are

1. Observation duration (minutes)
2. Record mode triplet (*Mpbs-nChan-nBit*)
3. Number of stations in job
4. Speed up factor (ratio of correlation time to observe time), or zero if correlation has not yet begun.
5. Percentage complete
6. Number of minutes remaining (only if Percentage complete isn't 0% or 100%)

Below these lines, five more lines containing information about the group of jobs as a whole is are presented. The contents of these lines are:

1. Total job time : Minutes of observe time in listed jobs
2. Fraction complete : Percentage in time through the entire project
3. Job time remaining : Minutes of observation left to be correlated
4. Wall time remaining : Minutes of real time needed to complete jobs
5. Average speedup : Ratio of total correlation time to run time, up to current point

Note that the speedup and time remaining values are estimates and don't include model calculation, conversion to FITS, and job startup time.

Usage: `jobstatus` [*options*] [*dir1*] ... [*dirN*]

options can be:

`-h` or `--help` : print usage information and exit

dir is a directory for which to print job information (default is current shell directory). Multiple directories can be specified.

Example 1: `jobstatus`

Example 2: `jobstatus $JOB_ROOT/*`

Known bugs:

1. This program has not been updated to work with DiFX 2.0 output

6.44 **listcpus** (package : mk5daemon)

Python program **listcpus** uses ssh to connect to each machine listed in a file (usually `$DIFX_MACHINES`) and peaks at the list of CPUs on that machine and prints to stdout. Only the first column of this file is used and any content after a `#` is ignored. For each CPU on the machine, the model name, which usually also contains the CPU speed, is listed. For multi-core CPUs, each core will appear as its own CPU.

Usage: **listcpus** [*options*]

options can be:

- h or --help : print usage information and exit
- v or --verbose : increase output verbosity
- m *file* or --machines *file* : use *file* instead of `$DIFX_MACHINES` for list of machines to probe

Multiple directories can be specified.

Example 1: **listcpus**

Example 2: **listcpus -m myCPUs.list**

6.45 **makefits** (package : difx2fits)

This program is tuned for NRAO use; some modifications may be required for use at other sites. In particular, this program requires that jobs were queued to be run in `$DIFX_QUEUE_BASE//experiment/`.

Program **makefits** is basically a wrapper for **difx2fits** (§6.14) that does some sanity checking and ensures that files end up in the proper places with the proper names. This program is intrinsically *pass-based* and it bases its functionality on the `.joblist` (§7.27) file that is written by **vex2difx** (§6.102). One must run this program on the software correlator head node (`swc000` in the current VLBA DiFX implementation). Upon successful completion, FITS-IDI files are created in the same directory in the correlator job staging area (`$DIFX_QUEUE_BASE/projectName`) and the sniffer output files are left in a subdirectory of the current working directory. An additional output file is left in the current working directory called `passName.fitslist`. This file has a list of the FITS files that are to be archived once the data for this pass are deemed valid.

The checks that **makefits** performs will by default not allow an incomplete set of FITS files to be produced. This can be overridden with a special command line argument (below). This is part of an accountability chain that aims to ensure that nothing gets omitted.

Usage: **makefits** [*options*] *passName*

options can be:

- h or --help : print usage information and exit
- v or --verbose : increase output verbosity
- override-version : ignore potential difx version conflicts
- allow-partial : bypass check for complete set of correlated output and proceed

Multiple directories can be specified.

Example: **makefits clock**

6.46 **makemark4** (package : difxdb)

This program is tuned for NRAO use; some modifications may be required for use at other sites.

Program **makemark4** is essentially a wrapper for **difx2mark4** (§6.15) that does some sanity checking and ensures that files end up in the proper places with the proper names. This program is intrinsically *pass-based* and it bases its functionality on the **.joblist** (§7.27) file that is written by **vex2difx** (§6.102). One must run this program on the software correlator head node (**swc000** in the current VLBA DiFX implementation). Upon successful completion, Mark4 file sets are created in the same directory in the correlator job staging area (**\$DIFX.QUEUE.BASE/projectName**). An additional output file is left in the current working directory called *passName.mark4list*. This file has a list of the Mark4 file sets that are to be archived once the data for this pass are deemed valid.

Usage: **makemark4** [*options*] *passName*

options can be:

- h or --help : print usage information and exit
- v or --verbose : increase output verbosity

Example: **makemark4** **rdv95**

6.47 **m5bstate** (package : mark5access)

Program **m5bstate** will perform state counts on a baseband data file.

Usage: **m5bstate** *file* *format* *nFrames* [*offset*]

file is the file to decode

format is the format of the data

nFrames is the number of data frames (typically a few kB in size) to decode

offset is the number of bytes into the file to start decoding

Example: **m5bstate** **sample.vlba** **VLBA1_2-256-8-2** **100**

Notes:

1. See documentation for **m5b** for details on specifying the data format.
2. Only real-sampled data with 1 or 2 bits per sample is supported at this time.
3. In the case of VDIF data, only single thread data with 2^n channels is supported. For equivalent functionality in the multi-thread VDIF case see **vdifd**.

6.48 **m5d** (package : mark5access)

Program **m5d** is a very simple example program using the mark5access decoding library. It turns out to be useful enough as a stand-alone program to be separately documented. This program takes as command line input the name of a file containing (or thought to be containing) VLBI baseband data, the expected format of the data, and the number of samples per baseband to decode. Optionally a starting file offset can be supplied. If the data can be decoded correctly, information about the data will be printed to the screen along with a table of decoded data. The output values, -3, -1, 1, or 3 for valid data, are printed in *nchan* columns. Data that cannot be decoded (either due to data replacement headers, data fill pattern replacing the actual data after unloading from a Mark5 module, or identified via the VDIF invalid bit) will show as 0. It should be invoked with the following parameters:

Usage: `m5d file format n [ offset ]`

file is the file to decode

format is the format of the data

n is the number of samples to decode

offset is the number of bytes into the file to start decoding

Example 1: `m5d sample.vlba VLBA1_2-256-8-2 24`

Example 2: `m5d sample.mk4 MKIV1_4-128-2-1 600 200`

Example 3: `m5d sample.5b Mark5B-512-16-2 1200`

The format parameter is constructed from four parts as *type-rate-nchan-nbit* where:

type is the type of format and should be one of VLBA1_1, VLBA1_2, VLBA1_4, MKIV1_1, MKIV1_2, MKIV1_4, Mark5B, or VDIF

rate is the data rate in Mbps

nchan is the number of baseband channels

nbit is the number of bits per recorded sample

See the usage examples above for some explicit values. Note for the VLBA and MKIV format types the fanout is appended as this affects the decodability of the files.

Notes:

1. In the case of VDIF data, only single thread data with 2^n channels is supported. For equivalent functionality in the multi-thread VDIF case see `vdifd`.

6.49 m5findformats (package : mark5access)

Program `m5findformats` attempts to determine which format a baseband data file may be. Currently it searches over 16 to 2048 Mbps data rates in factors of 2 and checks only for MKIV, VLBA and Mark5B types.

Usage: `m5findformats filename`

- *filename* is the name of the baseband data file.

Run with no command line arguments to get help information.

6.50 m5fold (package : mark5access)

Program `m5fold` takes a baseband data stream and integrates the power (formed by squaring the voltage) in a number of time bins that equally divide a given period. This is a simplified version of “folding” such as is used in pulsar processing. A typical use of such functionality would be to investigate the waveform of the switched power injected into the receiver for calibration. This program has found considerable utility in determining time offsets between the sample clock and formatter time (modulo the period of the calibration cycle). In the case of 2-bit sampling a non-linear correction is applied before results are written to a file. This correction takes the form

$$P = \frac{1}{\left(\operatorname{erf}^{-1}\left(\frac{\hat{P}-v_{\text{high}}^2}{1-v_{\text{high}}^2}\right)\right)^2}, \quad (2)$$

where P is a value proportional to true power and $\hat{P}$ is the value obtained by calculating $\langle \hat{v}^2 \rangle$ when the bitstream is reproduced with values $\hat{v} \in (-v_{\text{high}}, -1, 1, v_{\text{high}})$. This non-linear correction can be turned off by setting $nbin$ to a negative value. Note that this program is not useful for 1-bit quantized data. The program should be used as follows:

Usage: `m5fold infile format nbin nchunk freq outfile [ offset ]`

infile is the file to decode

format is the format of the data

nbin is the number of bins to calculate per period; if negative, power correction is not performed and the absolute value of *nbin* is used

nchunk is the number of 10000 sample chunks to operate on

freq is the reciprocal of the period to be observed (Hz)

outfile is the name of the output file

offset (optional) is the number of bytes into the file to start decoding

Example: `m5fold sample.vlba VLBA1_2-256-8-2 128 10000 80 sample.fold`

See the documentation for `m5d` for information on specifying the data format.

The output file will contain $nchan+1$ columns where $nchan$ is the number of baseband channels in the data stream. The first column contains the time (seconds) within the period. Each remaining column is folded power for one baseband channel. If $nbin$ is positive and the data is 2-bit quantized, the scaling is such that $\langle v^2 \rangle = \sigma^2$ yields a power reading of 1.0, for sampler threshold σ . Optimal signal to noise ratio occurs for a value of about 1.03. For non 2-bit quantization, the power will be in units of reconstituted counts².

In the case of VDIF data, only single thread data with 2^n channels is supported. For equivalent functionality in the multi-thread VDIF case see `vdiffold`.

6.51 m5pcal (package : mark5access)

Program `m5pcal` can be used to extract pulse cal tones from baseband data in Mark4, VLBA, Mark5B and single-thread VDIF formats.

Usage: `m5pcal [ options ] infile format freq1 [ freq2 [...] ] outfile`

options can be:

- h or --help : print usage information and exit
- c *n* or --chunksize *n* : use a fixed rather than automatic chunk size
- v or --verbose : increase output verbosity
- q or --quiet : decrease output verbosity
- n *n* : loop over *n* chunks of data (default is 1000)
- N *N* : perform *N* outer loops, each yielding a result set
- o *o* or --offset *o* : jump *o* bytes into file
- i *i* or --interval *i* : use pulse cal comb interval of *i* MHz (default is 1)
- e *e* or --edge *e* : don't use channels closer than *e* MHz from band edges when computing delay (default is 1/8 of bandwidth)

infile is the file to decode

format is the format of the data‘

freq1 ... is/are the frequencies (MHz) relative to baseband of the first tone to detect; there should be one *freq* specified per baseband channel

outfile is the name of the output file

Note: The position of the first tone in a baseband channel (*freq1* for baseband 1, and so on) must not be larger than the tone interval (set with *-i i*). All tones are extracted from each baseband channel. The tone interval is allowed to exceed the bandwidth of a baseband channel in which case *freqN* will effectively select just a single tone from the baseband.

See the documentation for **m5d** for information on specifying the data format.

6.52 **m5slice** (package : mark5access)

6.53 **m5spec** (package : mark5access)

Program **m5spec** is an example program using the mark5access decoding library that is a bit more advanced than the **m5d** program is. It forms total power spectra for each baseband channel in the data, including cross spectra for polarization pairs, assuming data is in alternating polarization pairs (if not, the cross spectra should make no sense, but they are formed anyway). The results are written to a text file with the following columns: Column 1 is the frequency offset from baseband for each channel; Columns 2 to *nchan*+1 are the total power spectra for each baseband channel; Columns *nchan*+2 to $4\times nchan+1$ contain, in pairs, the amplitude and phase of the cross spectra for each pair of channels. It should be invoked with the following parameters:

Usage: **m5spec** *infile format npoint n outfile* [*offset*]

infile is the file to decode

format is the format of the data

npoint is the number of points to calculate for each spectrum

n is the number of FFT frames to include in the calculation

outfile is the name of the output file

offset (optional) is the number of bytes into the file to start decoding

Example: **m5spec** *sample.vlba* VLBA1_2-256-8-2 256 1000 *vlba.spec*

See the documentation for **m5d** for information on specifying the data format.

In the case of VDIF data, only single thread data with 2^n channels is supported. For equivalent functionality in the multi-thread VDIF case see **vdifspec**.

6.54 **m5test** (package : mark5access)

Program **m5test** is an example program using the mark5access decoding library that works its way through a VLBI baseband data stream attempting to decode data and header information to look for problems. Every million samples (per baseband channel) a summary line containing frame number, decoded date and time, and counts of valid and invalid frames are shown. After 20 invalid frames are encountered the program will stop. Otherwise the program will run until end of file or until interrupted by the user. Usage is as follows:

Usage: **m5test** *infile format* [*offset*]

infile is the file to decode

format is the format of the data

offset (optional) is the number of bytes into the file to start decoding

Example: `m5test sample.vlba VLBA1_2-256-8-2`

See the documentation for `m5d` for information on specifying the data format.

6.55 `m5time` (package : mark5access)

Program `m5time` decodes the time of the beginning of a Mark4, VLBA, or Mark5B datastream and prints the result in integer MJD and UT hours, minutes, seconds to the screen.

Usage: `m5time infile format`

infile is the file to decode

format is the format of the data'

6.56 `m5timeseries` (package : mark5access)

Program `m5timeseries` produces a power measurements for each of channel of a baseband data file, averaging over a specified time interval.

Usage: `m5timeseries infile format tint ntime outfile [ offset ]`

infile is the file to decode

format is the format of the data'

tint is the integration time per sample in milliseconds

ntime is the number of samples to generate

outfile is the name of the output file

offset (optional) is the number of bytes into the file to start decoding

Example: `m5timeseries sample.vlba VLBA1_2-256-8-2 6.25 8000 sample.series`

The output file contains *nchan*+2 columns of data where *nchan* is the number of channels in the data file. The first column is sample number. The second column is time since beginning of series, in seconds. The remaining columns are power measurements for the channels.

6.57 `m5tsys` (package : mark5access)

6.58 `mk5cat` (package : mk5daemon)

This program sends data on a module to standard out. See additional documentation under `mk5cp` which operates on similar principles (`mk5cat` is `mk5cp` writing to *stdout* Note that the executable for `mk5cat` is identical to that for `mk5cp` and only the name of the program actually differs.

Usage: `mk5cat [ options ] { bank | VSN } scans`

options can be:

`-h` or `--help` : print usage information and exit

`-v` or `--verbose` : increase verbosity, e.g., print directory to screen

bank is either A or B

VSN is a valid 8-character VSN of a loaded module

scans is one or more scan numbers (starting at 1) with scan numbers separated by commas.

Many of the other baseband data utilities documented here such as `m5d`, `m5spec` and `vmux` can take input from *stdin* and thus can be mated with `mk5cat`. Usually a single hyphen (-) as the name of the input file indicates this to these programs.

Example: `mk5cat B PT_BB241_No0111 | m5spec - Mark5B-2048-16-2 128 10000 methanol.spec`

6.59 `mk5control` (package : `mk5daemon`)

`mk5control` is a program that sends XML messages of type `DifxCommand` to the `mk5daemon` programs that run on the software correlator cluster members. This program is a superset of `mk5take` and `mk5return`, allowing any allowed command to be sent.

Usage: `mk5control [ options ] command unit1 ... unitN`

options can be:

`-h` or `--help` : print usage information and exit.

command is the (non-case-sensitive) command to be executed; see list below.

unit is the number of a correlator Mark5 unit, a range, `all` for all software correlator cluster members, `mark5` for all Mark5 units, or `swc` for all software correlator compute nodes.

Example 1: `mk5control stopmark5a 07 08 09 11 14`

Example 2: `mk5control resetmark5 14-24`

Example 3: `mk5control startmark5a mark5`

The list of supported *command* types is below. All commands are not case sensitive.

- **GetVSN** Request a `Mark5Status` XML document to be multicast from the *unit*
- **ResetMark5** Execute `SSReset` and `ssopen`; this cures many/most mark5 hangs
- **Clear** Clear the stat of the Mark5 unit and get the VSNs, can be dangerous if other programs are currently using the Streamstor card
- **Reboot** Reboot the machine
- **Poweroff** Shut down the machine
- **StopMk5Daemon** Stop the `mk5daemon` program; you probably never need to do this
- **GetDir** Extract the directory from the modules in both banks and save to files in `$MARK5_DIR_PATH`
- **GetDirA** Same as above, but look only at bank A
- **GetDirB** Same as above, but look only at bank B
- **StopDir** Stop a directory read that is in progress
- **KillMpifxcorr** Send `sigkill` (like `kill -9`) to all processes on machine called `mpifxcorr`

- **Copy** Copy data from a module to files in a provided directory. At least three parameters must be provided that match the parameters of **mk5cp**. Because of the way **mk5control** parses the command line, the word **copy** and the parameters must all be enclosed in quotes.
- **StopCopy** Stop a data copy process
- **GetVer** Request send of a **DifxMessageMk5Version** XML message
- **mountXX** Cause Linux device **/dev/sdXX** to be mounted on **/mnt/usb**
- **umount** Cause **/mnt/usb** to be unmounted
- **Test** Used in debugging — for developers only

6.60 **mk5cp** (package : mk5daemon)

Program **mk5cp** copies baseband VLBI data from a module to a file somewhere on the operating system filesystem, perhaps an external USB disk. This program is often started using **mk5control** to tell the instance of **mk5daemon** running on the desired Mark5 unit to start the copy. Status information is multicast via a **Mark5StatusMessage** document.

Usage: **mk5cp** [*options*] { *bank* | *VSN* } *scans outputDirectory*

options can be:

- h or --help : print usage information and exit
- v or --verbose : increase verbosity: print directory to screen

bank is either A or B

VSN is a valid 8-character VSN of a loaded module

scans is one or more scan numbers (starting at 1) with scan numbers separated by commas. No spaces are allowed in the list. A range can be specified with a hyphen (see examples). Alternatively, a scan name, or portion thereof, can be specified. When a partial scan is provided, any scan name matching that partial scan name will be copied. The *scans* parameter can also take either a time range (two floating point modified Julian Days connected with an underscore, or a byte range (must be a multiple of 4) via two integers separated by an underscore, or a byte start and a length can be specified with two integers separated by a plus sign.

outputDirectory is the directory to which files will be copied. Make sure the destination directory exists before running this program and make sure sufficient free space remains on that filesystem. If the *outputDirectory* parameter is set to - (the hyphen), data will go to stdout. Another utility called **mk5cat** is derived from this behavior.

Example 1: **mk5cp** NRA0-123 4 /mnt/usb/WaltersProject

Example 2: **mk5cp** A 1,2,3 /tmp

Example 3: **mk5cp** B BC120A /mnt/usb/BC120A/PT

Example 4: **mk5cp** NRA0+255 6-12 /tmp

Example 5: **mk5cp** NRA0+255 123123100.124123100 /tmp

Example 6: **mk5cp** NRA0+255 54327.13124.54327.15124 /tmp

Example 7: **mk5cp** NRA0+322 123123100+1000000 /tmp

6.61 **mk5daemon** (package : mk5daemon)

mk5daemon is a program that started automatically at boot time on all of the software correlator cluster nodes (not only the Mark5 units!) that performs a number of operations in support of the software correlator.

The functions that **mk5daemon** performs are:

- **Logging**

All received multicast messages, significant internal functions, and interactions of the **Mark5A** program are logged to human readable log files. These log files are restarted at the beginning of each day. By default these log files are saved in `/tmp`.

- **High level control of Mark5 units**

The **Mark5A** program (written by Haystack) is the principle program used to access the Mark5 systems at the VLBA stations and the hardware correlator. DiFX directly accesses the Streamstor card via a library level programming interface. Since only one program is allowed to do this (or face a crash of varying degree of seriousness), access to the Streamstor card must be carefully managed. One function of **mk5daemon** is to maintain knowledge of who “owns” the Streamstor card at a given time to prevent conflicts. The starting and stopping of the **Mark5A** program can be requested by two messages of type **DifxCommand**: **startmark5a** and **stopmark5a**. When these commands are received by **mk5daemon**, the requested action is taken unless Streamstor conflict is likely. This type of command and others can be sent to **mk5daemon** with the **mk5control** program (§6.59).

- **Low level control of Mark5C units**

This program exposes a VLBI Standard Interface (VSI) over TCP port 2620 that very closely emulates equivalent functionality of the Mark5C Data Recording System (DRS) program provided by Haystack observatory. The implementation of the DRS command set is not complete but is sufficient for monitor and control at record time. At the time of writing this program is used in lieu of DRS at the two Mark5C units provided by USNO.

- **CPU, memory, and network monitoring**

Every 10 seconds, **mk5daemon** extracts data from the `/proc` directory to get information about the CPU load, memory usage, and network traffic. These numbers are multicast in a **DifxLoad** message and logged.

- **Module VSN and state determination**

Receipt of a multicast **getvsn** command will result in **mk5daemon** multicasting out a **Mark5Status** message containing information on the VSNs of the inserted modules as well as the state of the Mark5 unit. When **Mark5A** is running, a socket is opened to this program and the **bank_set?** query is issued, which returns the VSNs, regardless of the activity. When **Mark5A** is not running, **mk5daemon** either directly determines the VSNs through a Streamstor API library call if the Mark5 unit is idle, or doesn't respond if the Mark5 unit is busy. With each **Mark5Status** message that is multicast from **mk5daemon** the state of the Mark5 unit is included. See §8 for details on these XML messages.

- **Starting of mpifxcorr**

If **mk5daemon** is started with the `-H` or `--head-node` option, it will be allowed to start new correlations. A correlation will be started when a **difxStart** message is received if it passes some minor sanity checks. Since **mk5daemon** runs as root, it has the capability of changing file ownerships. By default, the output files from **mpifxcorr** and **difxlog** will have their ownership and permissions changed to match those of the `.input` file.

Normally **mk5daemon** is started automatically, either by `/etc/rc.local` or by a script in `/etc/init.d`. The command line options supported are:

Usage: `mk5daemon [ options ]`

options can be:

- h or --help : print usage information and exit
- q or --quiet : be less verbose and don't multicast state
- H or --head-node : give head-node permissions
- m or --isMk5 : force mk5daemon on this host to act as Mark5 regardless of hostname (default is mark5fx??)
- u *userID* or --user *userID* : use *userID* when executing remote commands (default is 'difx')
- l *logPath* or --log-path *logPath* : put logs in directory *logPath*, not /tmp

Please be sure not to have multiple instances of `mk5daemon` running at any one time on any individual Mark5 or correlator unit!

6.62 `mk5dir` (package : mark5daemon)

Program `mk5dir` extracts the directory from a module. Normally one would not call this program directly but would use the `getdir` option of `mk5control`. By default this program will change the disk module state to `played`. There is a danger that if this is done with an SDK 9 unit and the disk is later needed in an SDK 8 unit that it will no longer be readable in the later without a full reset of its VSN. As of April 2014 this program supports decoding of all directory types described by Mark5 Memos 81 (http://www.haystack.mit.edu/tech/vlbi/mark5/mark5_memos/081.pdf) and 100 (http://www.haystack.mit.edu/tech/vlbi/mark5/mark5_memos/100.pdf).

Usage: `mk5dir [ options ] { bank | VSN }`

options can be:

- h or --help : print usage information and exit
- v or --verbose : increase verbosity: print directory to screen
- f or --force : force directory read even if current
- F or --fast : get format details from new-style directory (Mark5B format only)
- n or --nodms : get directory but don't change disk module state
- s or --safedms : only change disk module state if SDK 8 unit or new dir type
- d or --dmsforce : proceed with change of module state even if this makes module unreadable in SDK 8 units
- b *b* or --begin *b* : begin with scan number *b*
- e *e* or --end *e* : end with scan number *b*
- w *file* or --write *file* : write directory file *file* to the module

bank is one of A, B or AB

VSN is a valid 8-character VSN of a loaded module

The resultant directory file will be saved in a file called `VSN.dir` in the directory pointed to by environment variable `MARK5_DIR_PATH`.

This program responds to the value of environment variable `DEFAULT_DMS_MASK`. This variable should be an integer representing the state of three bits. `mk5dir` only responds to the setting of bit 1 (value 2); if this bit is set, the disk module state will not be updated on directory reading. It is recommended to set this environment variable at recording stations so auto-erasure of modules does not occur.

In the mode where a specified `.dir` file is to be written to a Mark5 module directory the VSN must be provided explicitly (i.e., selecting by bank is not allowed)

6.63 **mk5erase** (package : mark5daemon)

Program **mk5erase** replaces the functionality of **SSErase**. It is used to either erase or condition a Mark5 module. It supports SDK9 and earlier revisions of the Conduant API and either legacy or new (see Mark5 memop 81) module directories. Conditioning results are multicast upon conclusion of conditioning, to be received by **condition_watch**. Conditioning (which is started with the **-c** option) causes an entire read/write cycle of the entire module to be performed. This can require a good fraction of 24 hours to complete. Status updates and progress are sent every 10 seconds as well. By default the original directory version will be restored on the module, with zero scans. The version of directory to use can be forced with either the **-l** or **-n** options.

Usage: **mk5erase** [*options*] *VSN*

options can be:

- h** or **--help** : print usage information and exit
- v** or **--verbose** : increase verbosity: print directory to screen
- f** or **--force** : force directory read even if current
- c** or **--condition** : Do full conditioning, not just erasing
- r** or **--readonly** : Perform read-only conditioning
- w** or **--writeonly** : Perform write-only conditioning
- d** or **--getdata** : Save the performance data to a file called *VSN.timedata*
- l** or **--legacydir** : Put an empty legacy directory on the module when complete
- n** or **--newdir** : Put an empty new style directory on the module when complete
- 0** or **--nodir** : Put a zero-size directory on the module when complete

VSN is a valid 8-character VSN of a loaded module

Note that this program will not run without specifying a legal mounted module VSN. If you wish to erase a module that has no VSN set, use the **vsu** program first.

Control-C can be used to safely abort conditioning early. The directory will be left in an indeterminate state, so use caution when doing this; if conditioning is stopped before completion use the **vsu** program to assess and possibly modify the current module state.

6.64 **mk5mon** (package : difxmessage)

Program **mk5mon** is a program that listens for **mark5Status** messages multicast from the Mark5 units and displays the information; updating the display as new messages are received.

Usage: **mk5mon**

Make sure the terminal is at least 110 characters wide and is at least as tall in characters as there are Mark5 units that may transmit information. To quit, use ctrl-C. The columns being displayed are:

1. Mark5 unit name
2. VSN of module in Bank A
3. VSN of module in Bank B
4. State of the Mark5 unit
5. Playback rate, if playing, in Mbps
6. Playback position, in bytes from beginning of module
7. Scan number of data being played, if playing
8. Scan name of data being played, if playing

6.65 **mk6cp** (package : mark6sg)

mk6cp is a wrapper around **mk6gather** which makes the operation of copying multiple files from a Mark6 module easier.

Usage: **mk6cp** [*options*] *filematch1* [*filematch2* ...] *destination*

options can be:

- h or --help : print usage information and exit
- v or --verbose : increase output verbosity
- r or --resume : don't copy files that exist in the destination path

filematch1 is a shell-style pattern for matching scan names

destination is the output path to place files; needs to start with . or / or end with /

6.66 **mk6gather** (package : mark6sg)

mk6gather extracts data from a Mark6 module, assembling as necessary the data scattered across the disks in the module.

Usage: **mk6gather** [*options*] *template*

options can be:

- h or --help : print usage information and exit
- v or --verbose : increase output verbosity
- a or --append : append to existing file; this continues a previous gather
- o *outfile* or --output *outfile* : send output to *outfile* (default is **gather.out**)
- b *bytes* or --bytes *bytes* : stop writing after *bytes* are written
- s *bytes* or --skip *bytes* : skip first *bytes* of file

template specifies a file list to match, in a shell-style wildcard pattern, such as the name of a scan (e.g., BB407A_LA.No0001)

If more than one file is to be copied, it is best to use the **mk6cp** (6.65) program instead.

6.67 **mk6ls** (package : vdifio)

mk6ls looks in the standard Mark6 mountpoint locations (either in **/mnt/disks/*/*/data** or in the path pattern set by environment variable **\$MARK6_ROOT** for VDIF formatted files. Found files will be probed and summarized.

Three printing levels are supported:

- *short* : Simply prints the file names. The portion of the filename corresponding to the mountpoint location is excised.
- *long* : Prints the file names (same as for *short*), the number of actual files making up the virtual Mark6 file (e.g., scattered across multiple mountpoints), the full size of the virtual file, and an indication of the completeness of the virtual file set.
- *full* : Prints the same information as for *long* followed by details of the files, such as details of the Mark6 file version, block sizes, packet numbers. This mode is mainly useful for developers with access to the vdifio source code.

Usage: **mk6ls** [*options*]

- h or --help : print usage information and exit
- s or --short : print long form output (default)
- l or --long : print long form output
- f or --full : print full information for each file

6.68 mk6mon (package : difxmessage)

6.69 mk6summary (package : mark6sg)

6.70 mk6vmux (package : vdfio)

6.71 mpifxcorr

The core of the DiFX software correlator is the program called **mpifxcorr**. This program uses Message Passing Interface (MPI) to exploit parallel computing to make correlation practical on a cluster of ordinary computers. This program runs on all the machines listed in the **.machines** file that is passed to **mpirun**, the program that starts **mpifxcorr**. It should be initiated from the cluster head node from within the project directory. The usage line below is appropriate for use with OpenMPI (§??) and within the DiFX context; other incantations may provide better results depending on the setup. See the OpenMPI documentation for more details.

Usage: **mpirun -np nProcess --bynode --hostfile** [*otherOptions*] *machinesFile* **mpifxcorr** *inputFile* [*options*]

nProcess is the number of processes to start; found with **wc -l machineFile**

machinesFile the **.machineFile**

inputFile the **.input** to run; the full path to this file needs to be given, so prepending the file with **'pwd' /** is typical

otherOptions can be any additional option to **mpirun**; **startdifx** uses the **--mca btl ^udapl,openib --mca mpi_yield_when_idle 1** to suppress some warning messages and be less aggressive on networking

options are additional options that **mpifxcorr** can take which include:

- MmonHostname:[monSkip]** : hostname of a machine serving as a monitor data server, with optional value indicating how many records to skip between sends.
- rstartSec** : start *startSec* seconds into the job, writing a new set of files into the visibility directory (**.difx/**)
- vgoscomplex** : flips sideband of all Complex VDIF stations by complex conjugating the unpacked Complex VDIF data

Within the DiFX framework, the user should never have to directly start **mpifxcorr** as this is done more simply with **startdifx** or via the DiFX Operator Interface in conjunction with **mk5daemon**.

6.72 oms2v2d (package : vex2difx)

The VLBI scheduling program **sched** generates a file with extension **.oms** which is used to populate some fields in the VLBA database. These fields are usually used to feed the dynamic scheduler but can also be used to reduce the tedium of transferring information from the **sched** input file (**.key**) to the **vex2difx** input file **.v2d**. For simple experiments this resulting **.v2d** file can be used unedited, but for more typical experiments editing will be required. The resulting file will have the same file prefix as the input file and will end with **.v2d**.

Usage: **oms2v2d** [*options*] *file.oms*

file.oms is an **.oms** file written by **sched**

options can be:

-h or **--help** : print usage information and exit

-f or **--force** : allow overwrite of output file

Example: **oms2v2d** **bx123.oms**

Note: **sched** now produces a **.tv2d** file (template **vex2difx**) that can contain useful information for some projects (especially multi-phase-center projects), however, this file is not tied to any particular version of DiFX (or more importantly, version of **vex2difx**) and thus cannot be guaranteed to be legal. It is thus suggested to use **oms2v2d** and transfer over needed information by hand after the fact.

6.73 padVDIF (package : vdifio)

Program **padVDIF** takes an input VDIF file and inserts additional packets as needed to create a contiguous without gaps. Newly inserted frames will have the invalid bit set.

Usage: **padVDIF** *infile outfile Mbps* [*newStartMJD*]

infile is the input VDIF file

outfile is the output VDIF file

Mbps is the data rate in megabits/second

newStartMJD is the MJD (with fractional component) to overwrite the times with

Example: **padVDIF** **raw.vdif smooth.vdif 2048**

6.74 plotapd (package : SniffPlots)

Program **plotapd** takes a text file containing *sniffer* fringe fit results and makes plot files. Separate plots for Amplitude, Phase and Delay (hence the name suffix “apd” are made for each baseline in the resultant file. A plot of delay rate is also produced.

This is an interactive command line program; running **plotapd** will prompt the user for inputs. The program **difxsniff** will run **plotapd** and its sister programs automatically, so usually it won’t be necessary to run by hand.

The **PGPLOT_FONT** environment variable must be set, otherwise all plot text will be missing.

6.75 **plotbp** (package : SniffPlots)

Program **plotbp** takes a text file containing *sniffer* bandpass output files and creates plots. This program can produce both auto- and cross-correlation data plots.

This is an interactive command line program; running **plotbp** will prompt the user for inputs. The program **difxsniff** will run **plotbp** and its sister programs automatically, so usually it won't be necessary to run by hand.

The `PGPLOT_FONT` environment variable must be set, otherwise all plot text will be missing.

6.76 **plotwt** (package : SniffPlots)

Program **plotwt** takes a text file containing *sniffer* data weights. In this context, a data weight of 0 indicates complete loss of data and a weight of 1 indicates complete data. Each plotted data point will usually span many correlator integration periods. A solid dot will be plotted for the mean of these points, and "error bars" will indicate the range of weights over the averaging period. Note that in some cases where awkward integration times are used it may be possible for the weight to occasionally exceed 1, as long as the long running average never does.

This is an interactive command line program; running **plotwt** will prompt the user for inputs. The program **difxsniff** will run **plotwt** and its sister programs automatically, so usually it won't be necessary to run by hand.

The `PGPLOT_FONT` environment variable must be set, otherwise all plot text will be missing.

6.77 **printDiFX.py** (package : vis2screen)

Program **printDiFX** prints a summary of the visibility information in a DiFX output file. It loops through all the records printing some basic info about frequencies, baselines, polarizations, times etc., plus a couple of selected visibility values from the start and middle of the record, to the screen.

Usage: **printDiFX** *difxfile inputfile*

difxfile is the full name of the visibility file in the `.difx` directory

inputfile is the path to the input file used to generate this difx output

Example: **printDiFX** `example_1.difx/DIFX.55523.025239.s0000.b0000` `example_1.input`

6.78 **printVDIF** (package : vdifio)

Program **printVDIF** loops through a VDIF file inspecting each packet header and printing some basic summary info (time etc.) to the screen.

Usage: **printVDIF** *vdiffile Mbps*

vdiffile is the recorded VDIF file to inspect

Mbps is the data rate in megabits/second

Example: **printVDIF** `example.vdif` `256`

6.79 **printVDIFgaps** (package : vdifio)

6.80 **printVDIFheader** (package : vdifio)

Program **printVDIFheader** is a powerful diagnostic tool that prints details of each VDIF header found in a VDIF file. All fields of each header, including information in known Extended VDIF Headers (see Sec. ??) are printed. Three different print formats are allowed: compact **short** version, detailed **long**, and hexadecimal **hex**.

Usage: **printVDIFheader** *inputFile* [*inputFrameSize* [*printLevel*]]

inputFile is the input multi-thread VDIF file, or - for *stdin*

inputFrameSize is the size of one thread's data frame, including header (for RDBE VDIF data this is 5032)

printLevel describes what is to be printed; one of **short**, **long**, or **hex**

Run this program with no arguments to get some additional explanation.

6.81 **psrflag** (package : difxio)

psrflag is a program that reads one or more DiFX filesets and produces a flag table readable by AIPS UVFLG that contains flags for times when the fringe rate, on a per-baseline basis, resonates with the pulse period, allowing DC bias to correlate. The pulsar **.binconfig** file is used to determine the harmonic content of the pulsar profile.

Usage: **psrflag** [*options*] *inputFile* ...

inputFile is the input multi-thread VDIF file; multiple may be provided

options can be:

- h or --help : print help information and quit
- v or --verbose : be more verbose in execution

6.82 **record5c** (package : mark5daemon)

6.83 **recover** (package : mk5daemon)

recover is a program that wraps the XLRRecover call for convenient use. This replaces the functionality of the **recover=** command of the **Mark5A** program.

Usage: **recover** [*options*] *type bank*

type is the type of recovery to attempt. See below.

bank should be either A or B and is the bank containing the module to address.

options can be:

- h or --help : print usage information and exit
- f or --force : allow overwrite of output file
- v or --verbose : be more verbose in execution

Example: **recover -v 2 A**

There are three possible modes of operation that are selected with the *type* argument:

- 0 Fix directory if power failed during recording
- 1 Allow access to data that might have been overwritten
- 2 Unerase the module

These recovery attempts will not always be successful.

6.84 **reducepoly** (package : difxio)

Program **reducepoly** takes one or more DiFX file sets as input and will modify each fileset's delay model (.im file) to have polynomial representations with fewer terms. All polynomials in the .im file will be reduced, including the baseline vectors, atmospheric components, azimuth, and elevation. The original .im files will be overwritten. The main purpose of this program is to evaluate the impact of using different polynomial orders,

Usage: **reducepoly** [*options*] *inputFile* ...

inputFile is the input multi-thread VDIF file, or - for *stdin*; multiple may be provided

options can be:

- h or --help : print help information and quit
- 2 : reduce polynomials to two terms
- 3 : reduce polynomials to three terms
- 4 : reduce polynomials to four terms
- 5 : reduce polynomials to five terms

6.85 **searchVDIF** (package : vdfio)

6.86 **splitVDIFbygap** (package : vdfio)

6.87 **startdifx** (package : mpifxcorr)

Starting **mpifxcorr** generally requires a lengthy command. This inspired the creation of **startdifx** which vastly simplifies use of the DiFX correlator. In addition to spawning the **mpifxcorr** processes, **startdifx** can orchestrate some of the preparatory work (for example running **calcif2** and **genmachines**) and optionally run **difx2fits** to create a .FITS file for each job. This program is meant to work within the DiFX environment and would probably require modification to be useful in other situations.

Usage: **startdifx** [*options*] [*startDelay*] *input1* [*input2* ...]

options can be:

- h or --help : print usage information and exit
- v or --verbose : be more verbose
- q or --quiet : be quieter
- f or --force : proceed on files even if correlator output already exists and is up to date
- a or --automachines : run **genmachines** only if no .machines file exists
- g or --genmachines : run **genmachines** unconditionally (default)
- n or --nomachines : don't run **genmachines**

`-d` or `--dont-calc` : don't run `calcif` even if needed – will skip file
`-m` or `--message` : start `mpifxcorr` by sending `difxStart` message to `mk5daemon`
`-F` or `--fits` : run `difx2fits` on output of each job separately
`-l` or `--localhead` : use the current host as the headnode, not `$DIFX_JEAD_NODE`
`--override-version` : ignore potential difx version conflicts
`-A agent` or `--agent=agent` : call `mpirun` through *agent* with filebase as only argument
`-M machinesFile` or `--machines-file=machinesFile` : use *machinesFile* instead of the one expected based on the job name

startDelay is an optional number of seconds to jump into the job upon start

inputN is a `.input` file, or its prefix

Example 1: `startdifx job1420.000.input`

Example 2: `startdifx -f -n job1420.000 job1421.000`

Example 3: `startdifx -F *.input`

Environment variables respected:

- `DIFX_MESSAGE_GROUP` : multicast group to use (when using `-m` option), overriding default 224.2.2.1
- `DIFX_MESSAGE_PORT` : multicast port to use (when using `-m` option), overriding default 50200
- `DIFX_HEAD_NODE` : when using `-m` option, this must be set, which specifies which machine will serve as the head node
- `DIFX_MPIRUNOPTIONS` : used to pass options to the `mpirun` command
- `DIFX_CALC_PROGRAM` : can be used to change the delay model program (default is `calcif2`); only needed if model calculations have not been done
- `DIFX_CALC_OPTIONS` : used to override options to the delay model program

6.88 **statemon** (package : difxmessage)

Program **statemon** listens for multicast messages of the `difxStatus` variety and simply prints their contents to the terminal. This is mainly useful for debugging `mpifxcorr` and any programs responsible for launching it.

Usage: `statemon [ options ]`

options can be:

`-h` or `--help` : print usage information and exit

Environment variables respected:

- `DIFX_MESSAGE_GROUP` : multicast group to use, overriding default 224.2.2.1
- `DIFX_MESSAGE_PORT` : multicast port to use, overriding default 50200

6.89 stopmpifxcorr (package : mpifxcorr)

If software correlation is in progress and it is desired to stop it, it is best to gently stop it rather than killing it abruptly. In most circumstances this can be accomplished with **stopmpifxcorr**. This program must be run on the machine running the *manager* process of the software correlator (usually this is **swc000** for the VLBA). If multiple **mpifxcorr** processes are found running on a machine, **stopmpifxcorr** will not proceed unless the **-f** option is used.

Usage: **stopmpifxcorr** [*options*]

options can be:

- h** or **--help** : print usage information and exit
- f** or **--first-pid** : send stop message to the numerically first process ID found
- q** or **--quiet** : don't produce much output

6.90 stripVDIF (package : vdifio)

Program **stripVDIF** strips network headers from a VDIF format baseband data file (e.g., captured from Wireshark) and dumps a pure VDIF stream.

Usage: **stripVDIF** *infile* *outfile* [*skipbytesfront* [*skipbytesback* [*skipbytesinitial*]]]

infile is the input VDIF file

outfile is the output VDIF file

skipbytesfront is the number of bytes to skip over before each frame (default is 54)

skipbytesback is the number of bytes to skip over after each frame (default is 4)

skipbytesinitial is the number of bytes to skip over only once after opening the file (default is 28)

Example: **stripVDIF** *vdif.wireshark* *vdif.pure* 54 4 28

6.91 tabulateddelays (package : difxio)

Program **tabulateddelays** takes one or more DiFX filesets and produces a text file containing, for each scan, a table of interferometer delays (μ s) and rate (μ s/s) based on the values in the *.im* file at a cadence of one entry every 8 seconds. Based on command line options values other than the delay can be extracted and tabulated. The details of the output files are documented in comments at the top of the output file.

Usage: **tabulateddelays** [*options*] *inputFile* ...

inputFile is the input multi-thread VDIF file, or - for *stdin*; multiple may be provided

options can be:

- h** or **--help** : print help information and quit
- az** : print azimuth (deg) and azimuth rate (deg/s) instead of delay, rate
- el** : print elevation (deg) and azimuth rate (deg/s) instead of delay, rate
- dry** : print the dry component of atmospheric delay (μ s) instead of delay, rate
- wet** : print the wet component of atmospheric delay (μ s) instead of delay, rate
- uvw** : print the antenna-based baseline coordinates (*x, y, z*) (meters) instead of delay, rate

--clock : print the clock offset (μ s) and rate (μ s/s) instead of delay, rate
--perint : print values at the center of every integration rather than every 8s
--addclock : include clock offset/rate in printed delay/rate values

This program reads through one or more difx datasets and evaluates delay polynomials in the `.im` files on a regular time grid (every 8 seconds). Delays and rates are both calculated. Output should be self explanatory. Plotting utilities such as `gnuplot` can be used directly on the output.

When operating without **--perint**, the entirety of the delay polynomials are plotted, even exceeding the time range of the scans to which they belong. Comments in the output separate scans cleanly. When **--perint** is used, only the time covered by the scans is output.

Sign conventions:

- Delay: a positive delay indicates wavefront arrival at the station before wavefront arrival at earth center. The delay includes contribution from wet and dry atmosphere components.
- Rate: simply the time derivative of Delay.
- Clock Offset: sign convention is opposite that of `.vex` "clock_early" parameter; a positive clock offset indicates slow station clock. The sum of Clock Offset and Delay is the total correlator delay.
- Clock Rate: simply the time derivative of Clock Offset.

6.92 testdifxinput (package : difxio)

This program was intended mainly for helping debug parsing of `.input` files and associated other files. It turned out to be useful as a general tool to investigate the contents of such files. When multiple input files are provided on the command line merging of the resultant data structures is attempted. Two output files are created when run: `input.out` and `calc.out`. These files should closely resemble the input files if the parsing was done properly.

Usage: `testdifxinput [ options ] inputFilePrefix1 [ inputFilePrefix2 ...]`

options can be:

-v or **--verbose** : be a bit more verbose
-h or **--help** : print help information and quit

inputFilePrefix is the base name of an input file

6.93 testdifxmessagereceive(package : difxmessage)

Test program `testdifxmessagereceive` captures multicast DiFX messages and prints them to the screen. Both the raw XML is shown and the decoded values. It is mostly useful as a tool for examining the correctness of the multicast messages that are broadcast and is not intended to be part of an operational system. There is a special binary mode which instead listens for the multicast high time resolution autocorrelations. In this mode, only a terse summary of what is received is printed (see the source code for more information).

Usage: `testdifxmessagereceive [ options ] [ type ]`

options can be:

-h or **--help** : print help information
-b or **--binary** : generate output based on binary records

type is the kind of message to capture (not for use with binary records):

1. DifxLoadMessage
2. DifxAlertMessage
3. Mark5StatusMessage
4. DifxStatusMessage
5. DifxInfoMessage
6. DifxDatastreamMessage
7. DifxCommand
8. DifxParameter
9. DifxStart
10. DifxStop
11. Mark5VersionMessage
12. Mark5ConditionMessage
13. DifxTransientMessage

If *type* is not provided, all message types will be captured.

6.94 **testmod** (package : mk5daemon)

testmod is a program that is used to perform read and write tests on Mark5 modules. It is meant as a replacement of the **ResetModule** program that relies on the **Mark5A** program which is being phased out of VLBA operations. Read-only tests can be performed without risk of erasing astronomical data recorded on the disks. The more invasive write-read tests (which are the default) will erase all preexisting data! A matrix of numbers similar to what is produced by **ResetModule** or **SSerase** in condition mode, but with statistics from a much smaller volume of reading/writing is produced. Usually badly performing disks will occur in pairs with both bad disks belonging to the same bus (e.g., disks 0 and 1, 2 and 3, 4 and 5, or 6 and 7). Badly performing drives should have their directory files **.dir** (see § 7.17) updated by hand to include **RT** at the end of the top line.

Usage: **testmod** [*options*] *bank*

options can be:

- h or --help : print usage information and exit
- v or --verbose : produce more informative/diagnostic output; -v -v for even more
- f or --force : continue to produce files despite warnings
- r or --readonly : Perform read-only test
- R or --realtime : Switches to real-time mode (see below)
- d or --skipdircheck : Disable directory checking (see below)
- S or --speed : Disable correctness testing to better test throughput
- n *n* or --nrep *n* : Repeat the test *n* times (default is 2)
- s *s* or --blocksize *s* : Read and write *bytes* at a time (default is 10 MB)
- b *b* or --nblock *b* : Perform *b* reads per test (default is 50)
- p *p* or --pointer *p* : Start read-only tests at byte position *p*
- o *file* or --dirfile *file* : Write the module directory to file *file*

bank is the Mark5 bank containing the disk to be studied (A or B)

Many modules being tested are perhaps damaged in some way and may require the **-R** and/or **-d** options above. Is is generally safe to use these options, but the diagnostic power of this program may be reduced in cases where some drives are intrinsically slow, but still produce valid data.

6.95 testseqnumbers (package : difxmessage)

Program **testseqnumbers** is a utility to listen for DiFX multicast messages and identify any that come with a sequence number that is not sequential. This is a good way to identify possible packet loss or duplication on a DiFX cluster network.

Usage: **testseqnumbers** [*options*]

options can be:

- h or --help : print usage information and exit
- v or --verbose : produce more output; -v -v for even more

If run without the -v option, only unexpected packets will be noted. If run with one -v flag, each received packet will be identified with a period being written to the screen. If run with 2 -v flags, each packet received will have its source and sequence number printed.

6.96 vdif2to8 (package : vdifio)

Program **vdifb2to8** takes a 2-bit sampled VDIF stream and reencodes it as 8-bit samples. It should work on any form of VDIF with 2 bits per sample (2+2 bits complex). It is anticipated that any 8-bit decoder will be performed with linear level spacings, unlike the case for 2-bit samples which use a high to low ratio of 3.3359 to minimized quantization noise. Given this anticipation the levels chosen in the output 8-bit stream correspond to levels of 35.5 and 118.5 counts for the low and high states respectively. These levels lead to a ratio as close to 3.3359 as possible. The 0.5 offset comes about from assuming that the 256 output states are centered on zero and thus range from -127.5 ... -0.5, 0.5 ... 127.5. Extra bytes in the input stream (not corresponding to valid VDIF frames) will be excised but invalid frames (as marked with the invalid bit) will be retained and encoded to 8 bits.

Usage: **vdif2to8** *inputFile* *frameSize* *outputFile*

inputFile is the file to read (2 bits per sample)

frameSize is the size of the VDIF frames including frame headers (5032 for VLBA or VLA VDIF data)

outputFile is the output file (8 bits per sample)

Example: **vdif2to8** input.vdif 5032 output.vdif

6.97 vdifbstate (package : vdifio)

Program **vdifbstate** will perform state counts on a multi-thread VDIF baseband data file.

Usage: **vdifbstate** *file* *frameSize* *dataRate* *threadlist* *nFrames* [*offset*]

file is the file to decode

frameSize is the size of the VDIF frames including frame headers (5032 for VLBA or VLA VDIF data)

dataRate is the file data rate, measured in Mbps, not including frame headers

threadList is a comma-separated list of thread ids to decode

nFrames is the number of data frames (typically a few kB in size) to decode

offset is the number of bytes into the file to start decoding

Example: `vdifbstate sample.vdif 5032 1024 1,2,3,4 100`

Notes:

1. See documentation for `m5b` for details on specifying the data format.
2. Only real-sampled data with 1 or 2 bits per sample is supported at this time.
3. If a non-power-of-two number of threads is requested, extra channels will be invented to pad out the next power of two. Data in these extra channels has undefined qualities.

6.98 `vdifChanSelect` (package : `vdifio`)

6.99 `vdifd` (package : `vdifio`)

Program `vdifd` takes a multi-thread VDIF file and decodes some samples. It is implemented as a python script that makes use of `vmux` and `m5d` to do most of the work. It is thus a good program to study to understand how `vmux` can be used. This program takes as command line input the name of a file containing (or thought to be containing) multi-thread VDIF baseband data, information about the VDIF stream, and the number of samples per baseband to decode. Optionally a starting file offset can be supplied. If the data can be decoded correctly, information about the data will be printed to the screen along with a table of decoded data. The output values, -3, -1, 1, or 3 for valid data, are printed in *nchan* columns. Data that cannot be decoded (either due to data replacement headers, data fill pattern replacing the actual data after unloading from a Mark5 module, or identified via the VDIF invalid bit) will show as 0. It should be invoked with the following parameters:

Usage: `vdifd file frameSize dataRate threadlist n [ offset ]`

file is the file to decode

frameSize is the size of the VDIF frames including frame headers (5032 for VLBA or VLA VDIF data)

dataRate is the file data rate, measured in Mbps, not including frame headers

threadList is a comma-separated list of thread ids to decode

n is the number of samples to decode

offset is the number of bytes into the file to start decoding (default is 0)

Example: `vdifd sample.vdif 5032 1024 1,2,3,4 24`

In the above example, a stream consisting of 4 channels, each with 64 MHz bandwidth and 2 bits per sample and with thread ids 1, 2, 3 and 4 are to be decoded. If the thread ids were to be permuted, the decoded output data would be permuted in the same manner.

Notes:

1. At this time only 2-bit real-valued data can be properly decoded.
2. If a non-power-of-two number of threads is requested, extra channels will be invented to pad out the next power of two. Data in these extra channels has undefined qualities.

6.100 **vdiffold** (package : vdifio)

Program **vdiffold** takes a baseband data stream and integrates the power (formed by squaring the voltage) in a number of time bins that equally divide a given period. This is a simplified version of “folding” such as is used in pulsar processing. A typical use of such functionality would be to investigate the waveform of the switched power injected into the receiver for calibration. This program has found considerable utility in determining time offsets between the sample clock and formatter time (modulo the period of the calibration cycle).

In the case of 2-bit sampling a non-linear correction is applied before results are written to a file. This correction takes the form

$$P = \frac{1}{\left(\operatorname{erf}^{-1} \left(\frac{\hat{P} - v_{\text{high}}^2}{1 - v_{\text{high}}^2} \right) \right)^2}, \quad (3)$$

where P is a value proportional to true power and $\hat{P}$ is the value obtained by calculating $\langle \hat{v}^2 \rangle$ when the bitstream is reproduced with values $\hat{v} \in (-v_{\text{high}}, -1, 1, v_{\text{high}})$. This non-linear correction can be turned off by setting $nbin$ to a negative value. Note that this program is not useful for 1-bit quantized data.

The program should be used as follows:

Usage: **vdiffold** *infile* *frameSize* *dataRate* *threadlist* *nbin* *nchunk* *freq* *outfile* [*offset*]

infile is the file to decode

frameSize is the size of the VDIF frames including frame headers (5032 for VLBA or VLA VDIF data)

dataRate is the file data rate, measured in Mbps, not including frame headers

threadList is a comma-separated list of thread ids to decode

nbin is the number of bins to calculate per period; if negative, power correction is not performed and the absolute value of *nbin* is used

nchunk is the number of 10000 sample chunks to operate on

freq is the reciprocal of the period to be observed (Hz)

outfile is the name of the output file

offset (optional) is the number of bytes into the file to start decoding

Example: **vdiffold** *sample.vdif* 5032 1024 1,2,3,4 128 10000 80 *sample.fold*

The output file will contain $nchan+1$ columns where $nchan$ is the number of baseband channels in the data stream. The first column contains the time (seconds) within the period. Each remaining column is folded power for one baseband channel. If *nbin* is positive and the data is 2-bit quantized, the scaling is such that $\langle v^2 \rangle = \sigma^2$ yields a power reading of 1.0, for sampler threshold σ . Optimal signal to noise ratio occurs for a value of about 1.03. For non 2-bit quantization, the power will be in units of reconstituted counts².

Notes:

1. At this time only 2-bit real-valued data can be properly decoded.
2. If a non-power-of-two number of threads is requested, extra channels will be invented to pad out the next power of two. Data in these extra channels has undefined qualities.
3. The output columns are in the same order as the thread id list. Thus, you can rearrange the output order by changing the order of the thread list. This enables reordering of data so that polarization pairs occur consecutively, allowing more sensible cross-correlation columns.

6.101 **vdifspec** (package : vdifo)

Program **vdifspec** forms total power spectra for each baseband channel in the data, including cross spectra for polarization pairs, assuming data is in alternating polarization pairs (if not, the cross spectra should make no sense, but they are formed anyway). The results are written to a text file with the following columns: Column 1 is the frequency offset from baseband for each channel; Columns 2 to $nchan+1$ are the total power spectra for each baseband channel; Columns $nchan+2$ to $4 \times nchan+1$ contain, in pairs, the amplitude and phase of the cross spectra for each pair of channels. It should be invoked with the following parameters:

Usage: **vdifspec** *infile* *frameSize* *dataRate* *threadlist* *npoint* *n* *outfile* [*offset*]

infile is the file to decode

frameSize is the size of the VDIF frames including frame headers (5032 for VLBA or VLA VDIF data)

dataRate is the file data rate, measured in Mbps, not including frame headers

threadList is a comma-separated list of thread ids to decode

npoint is the number of points to calculate for each spectrum

n is the number of FFT frames to include in the calculation

outfile is the name of the output file

offset (optional) is the number of bytes into the file to start decoding

Example: **vdifspec** **sample.vdif** 5032 1024 1,2,3,4 256 1000 **vlba.spec**

Notes:

1. At this time only 2-bit real-valued data can be properly decoded.
2. If a non-power-of-two number of threads is requested, extra channels will be invented to pad out the next power of two. Data in these extra channels has undefined qualities.
3. The output columns are in the same order as the thread id list. Thus, you can rearrange the output order by changing the order of the thread list. This enables reordering of data so that polarization pairs occur consecutively, allowing more sensible cross-correlation columns.

6.102 **vex2difx**

vex2difx is a program that takes a **.vex** files (such as one produced by **sched** with various tables based on observe-time data appended, probably by **db2vex** in the case of VLBA operations) and a **.v2d** configuration file (see §7.42) and generates one or more **.input** and **.calc** file pairs for use with the DiFX correlator. Note specifically that **.ovex** files, as used at many/most Mark4 correlators, are not supported. **vex2difx**, along with **calcif2**, supercedes the functionality of **vex2config** and **vex2model**, two programs that were widely used but never fully integrated into the VLBA's software chain. Don't forget that **oms2v2d** can be used to create a valid baseline **.v2d** file from the **.oms** file made by **sched**, perhaps saving some time.

The following guiding principles drove the design of **vex2difx**:

1. The output files should never need to be hand edited
2. Simple experiments should not require complicated configuration
3. All features implemented by **mpifxcorr** should be accessible
4. All experiments expressible by **vex** should be supported

5. The configuration file should be human and machine friendly
6. Command line arguments should not influence the processing of the vex file

Note that not all of these ideals have been completely reached as of now. It is not the intention of the developer to guess all possible future needs of this program. Most new features will be easy to implement so send a message to the difx-users mailing list or file a JIRA [9] bug tracking ticket for requests.

Usage: `vex2difx [ options ] v2dFile`

options can be:

- `-h` or `--help` : print usage information and exit
- `-o` or `--output` : write a configuration file called `v2dFile.params` (see §7.34) as output
- `-v` or `--verbose` : produce more informative/diagnostic output; `-v -v` for even more
- `-d` or `--delete-old` : delete old output from same `.v2d` file
- `-f` or `--force` : continue to produce files despite warnings
- `-s` or `--strict` : treat some warnings as errors and quit (default)

v2dFile is a `.v2d` file (see §7.42) that controls the operation of this program; the filename cannot contain underscore characters

Example: `vex2difx bx123.v2d`

6.102.1 VDIF issues

Unlike for the other formats, VDIF does not make use of the `$TRACKS` section for numeric assignment of channel. Instead the channels as listed in the `$FREQ` section are sorted alphabetically by their link name (usually something like `Ch01`). The alphabetical list is matched against the thread-channel order where the threads are listed in numeric order; the “thread index” takes precedence over the “channel index”. The `track_frame_format` parameter of the `$TRACKS` section is still required.

6.102.2 Mark5B issues

The Mark5B format, including its 2048 Mbps extension, is now supported by `vex2difx`. The `.vex` file track assignments for Mark5B format has never been formally documented. `vex2difx` has adopted the track assignment convention used by Haystack. Formally speaking, Mark5B has no tracks. Instead it stores up to 32 bitstreams in 32 bit words. The concept of fanout is no longer used with Mark5B. Instead, the equivalent operation of spreading one bitstream among 1 or more bits in each 32 bit word is performed automatically. Thus to specify a Mark5B mode, only three numbers are needed: Total data bit rate (excluding frame headers), number of channels, and number of bits per sample (1 or 2). The number of bitstreams is the product of channels and bits.

The `$TRACKS` section of the vex file is used to convey the bitstream assignments. Individually, the sign and magnitude bits for each channel are specified with `fanout_def` statements. In unfortunate correspondence with existing practice, 2 is the first numbered bitstream and 33 is the highest. In 2-bit mode, all sign bits must be assigned to even numbered bitstreams and the corresponding magnitude bit must be assigned to the next highest bitstream. To indicate that the data is in Mark5B format, one must either ensure that a statement of the form

```
track_frame_format = MARK5B;
```

must be present in the appropriate `$TRACKS` section or

```
format = MARK5B
```

must be present in each appropriate `ANTENNA` section of the `.v2d` file. As a concrete example, a `$TRACKS` section may resemble:

```

$TRACKS;
def Mk34112-XX01_full;
  fanout_def = A : &Ch01 : sign : 1 : 02;
  fanout_def = A : &Ch01 : mag : 1 : 03;
  fanout_def = A : &Ch02 : sign : 1 : 04;
  fanout_def = A : &Ch02 : mag : 1 : 05;
  fanout_def = A : &Ch03 : sign : 1 : 06;
  .
  .
  .
  fanout_def = A : &Ch15 : mag : 1 : 31;
  fanout_def = A : &Ch16 : sign : 1 : 32;
  fanout_def = A : &Ch16 : mag : 1 : 33;
  track_frame_format = MARK5B;
enddef;

```

6.102.3 Media specification

`vex2difx` allows `.input` file generation for two types of media. A single `.input` file can have different media types for different stations. Ensuring that media has been specified is important as antennas with no media will be dropped from correlation. The default media choice is Mark5 modules. The `TAPELOG.OBS` table in the input vex file should list the time ranges valid for each module. Jobs will be split at Mark5 module boundaries; that is, a single job can only support a single Mark5 unit per station. All stations using Mark5 modules will have `DATA SOURCE` set to `MODULE` in `.input` files. If file-based correlation is to be performed, the `TAPELOG.OBS` table is not needed and the burden of specifying media is moved to the `.v2d` file. The files to correlate are specified separately for each antenna in an `ANTENNA` block. Note when specifying filenames, it is up to the user to ensure that full and proper paths to each file are provided and that the computer running the datastream for each antenna can see that file. Two keywords are used to specify data files. They are not mutually exclusive but it is not recommended to use both for the same antenna. The first is `file`. The value assigned to `file` is one or more (comma separated) file names. It is okay to have multiple file keywords per antenna; all files supplied will be stored in the same order internally. The second keyword is `filelist` which takes a single argument, which is a file containing the list of files to read. The file pointed to by `filelist` only needs to be visible to `vex2difx`, not the software correlator nodes. This file contains a list of file names and optionally start and stop MJD times. Comments can be started with a `#` and are ended by the end-of-line character. Like for the `file` keyword, the file names listed must be in time order, even if start and stop MJD values are supplied. An example file as supplied to `filelist` is below:

```

# This is a comment. File list for MK for project BX123
/data/mk/bx123.001.m5a 54322.452112 54322.511304
/data/mk/bx123.002.m5a 54322.512012 54322.514121 # a short scan
/data/mk/bx123.003.m5a 54322.766323 54322.812311

```

If times for a file are supplied, the file will be included in the `.input` file `DATA TABLE` only if the file time range overlaps with the `.input` file time range. If not supplied, the file will be included regardless of the `.input` file time range, which could incur a large performance problem.

A few sample `ANTENNA` blocks are shown below:

```

ANTENNA MK
{
  filelist=bx123.filelist.mk
}

```

```
ANTENNA OV { file=/data/ov/bx123.001.m5a,
              /data/ov/bx123.002.m5a,
              /data/ov/bx123.003.m5a }
```

```
ANTENNA PT { file=/data/pt/bx123.003.m5a } # recording started late here
```

6.102.4 Pulsars

Some information, including example `.v2d` sections, on setting up pulsar correlation can be found in §4.

You may find additional information at <http://www.atnf.csiro.au/vlbi/dokuwiki/doku.php/difx/vex2difx>.

6.103 `vexpeek` (package : vex2difx)

Program `vexpeek` takes a vex file as input and sends to `stdout` the experiment name, segment, and a list of antennas and the MJD times that they were included in the observation. This program is mainly intended to be called from python program `db2vex` which needs to know a little about the file before appending the `CLOCK` and `TAPELOG_OBS` tables. The VLBA operations system relies on such functionality but there is no reason other operations couldn't use this. This program uses the same parsing infrastructure as `vex2difx` so the warnings that may be produced and sent to `stderr` in running `vex2difx` will also do so with `vexpeek`. Thus, when `db2vex` is run some of these error messages may be seen.

Usage: `vexpeek [options] vexFile`

options can be:

- h or --help : print usage information and exit
- v or --verbose : print decoded version of *vexFile* to screen
- f or --format : add per-antenna format description to output
- b or --bands : print list of bands used by *vexFile*
- u or --diskusage : add per-antenna disk usage to output
- s or --scans : print list of scans and antennas used by each

vexFile is the vex format file to be inspected.

6.104 `vlog` (package : vex2difx)

Program `vlog` takes as input a calibration file (`cal.vlba`, §7.10). This file is parsed to produce four files containing formatted arrays that are convenient for use in the construction of FITS tables: `flag`, `pcal`, `tsys`, and `weather` (§7.24, §7.35, §7.39 and §7.40). This program is named after AIPS task `vlog` that does nearly the same thing.

Usage: `vlog calFile [antennaList]`

calFile is the `cal.vlba` file produced by `tsm` to be processed.

antennaList is an optional comma-separated list of antennas to process. If omitted, all antennas with calibration data will be processed.

Running with no command line arguments will print usage information to the terminal and exit.

6.105 **vmux** (package : vdifio)

Program **vmux** takes a VDIF file with multiple threads of one channel each and multiplexes the data into a new VDIF file consisting of a single thread containing all the channels. In the case that a non-power-of-two number of channels are contained in the input file (or equivalently, a non-power-of-two number of threads are specified on the command line), the next power of two will be selected as the number of channels in the output thread and any unused channel slots will contain random data.

Usage: **vmux** [*options*] *inputFile* *inputFrameSize* *framesPerSecond* *threadList* *outputFile* [*offset* [*chunkSize*]]

inputFile is the input multi-thread VDIF file, or - for *stdin*

inputFrameSize is the size of one thread's data frame, including header (for RDBE VDIF data this is 5032)

framesPerSecond is the number of frames per second in the input file for each thread (and is thus the number of output frames per second as well)

threadList is a comma-separated list of integers in range 0 to 1023; the order of the numbers is significant and dictates the order of channels in the output data

outputFile is the name of the output, single-thread VDIF file, or - for *stdout*

offset is an optional offset into the input file (in bytes)

chunkSize is (roughly) how many bytes to operate on at a time (default is 2000000)

options can be:

- h or --help : print usage information and exit
- v or --verbose : be more verbose in execution
- q or --quiet : be less verbose in execution
- f *f* or --fanout *f*: set fanout factor to *f*
- e or --EDV4 : convert VDIF extended data version to EDV4 (default)
- n or --noEDV4 : don't convert VDIF extended data version to EDV4

The concept of fanout applies to certain variants of VDIF data where one logical sampled channel is interleaved multiple threads. Certain modes of the DBBC3 makes use of this mode. See Sec. ?? for details.

VDIF Extended Data Version (EDV) 4 is used to contain per-channel validity flags within a multi-channel VDIF file. See Sec. ?? for details.

6.106 **vsn** (package : mk5daemon)

Program **vsn** is used to check or set the Volume Serial Number (VSN), the write protect state, and the Disk Module State (DMS) of a module. This program can not be used when the Mark5 unit is being used for something else. You must be logged into the Mark5 unit that contains the module to inspect/change. In addition to displaying the VSN of the module, this utility will list information about each disk in the module. The columns displayed are:

1. *Disk number*: in the range 0 to 7.
2. *Drive model*: the model number of the disk.
3. *Serial number*: the serial number of the disk (in parentheses).
4. *Drive model revision number*: addition model information.

5. *SMART capable*: 1 indicates SMART information is available; 0 otherwise.
6. *SMART state*: (only valid if SMART capable) 1 indicates good health.

Note that the drive model, serial number and revision number can have spaces making it hard to tell when one field start and the next begins. Thus the serial number is contained within parentheses to make this clear.

Usage: **vsn** [*options*] *bank* [*newVSN*]

options can be:

- h or --help : print usage information and exit
- f or --force : proceed without asking
- v or --verbose : be more verbose in operation
- p or --played : set DMS to played
- r or --recorded : set DMS to recorded
- e or --erased : set DMS to erased
- w or --writeprotect : set write protection
- u or --unwriteprotect : clear write protection
- s or --smart : Get S.M.A.R.T. data from disks and write to file *VSN.smart*)

bank is the Mark5 unit bank to look at (must be A or B)

newVSN is the new name to assign to the module and must be a legal VSN

Example 1: show VSN: **vsn** A

Example 2: set VSN: **vsn** A NRA0+456

If you get a message such as “Watchdog caught a hang-up executing ...” that means access to the module failed. This could indicate a bad module. The module should be reinserted (perhaps in a different bank or unit) and the unit rebooted before coming to a firm conclusion.

6.107 **vsum** (package : vdifio)

Program **vsum** prints a summary of one or more VDIF files, printing such information as list of threads found, collectively, in the first and last few MB of the file, the VDIF epoch and other time information, and packet size. If the data is not recognized as VDIF, an error code will be printed; see below for the list of codes and their meanings. Legacy format VDIF data is not supported.

Usage: **vsum** [*options*] *file1* [*file2* [...]]

file1 ... is/are the VDIF file(s) to summarize

options can be:

- h or --help : print usage information and exit
- s or --shortsum : print one-line summary per file
- 6 or --mark6 : interpret provided file names as Mark6 scans
- allmark6 : summarize all files found on mounted Mark6 modules

If the `-6` or `--mark6` or `--allmark6` option is used, it is assumed that the files are to be found in their expected location, which could be altered by an environment variable. See sec ?? for more details.

If the summary operation failed for a file, one of the following error codes will be returned:

- 1 The file size could not be determined
- 2 The file could not be opened
- 3 Memory allocation failure
- 4 File read failed
- 5 Frame size could not be determined
- 6 First frame could not be found
- 7 Seek to near-end-of-file failed
- 8 Read at end of file failed
- 9 A valid frame at the end of the file could not be found

The `-s` or `--shortsum` option produces output that can be used directly by `vex2difx` as a file list.

6.108 **zerocorr** (package : mark5access)

Program **zerocorr** is intended to cross correlate data with zero time delay (and a window determined by the spectral resolution) between two recordings made at the same station. It is possible to correlate data observed in different formats and even mis-matched bandpasses, through the construction of a file describing the details of the single sub-band that is to be correlated.

Usage: **zerocorr** [*options*] *confFile*

options can be:

- `-h` or `--help` : print usage information and exit
- `-v` or `--verbose` : be more verbose in operation

confFile is a file describing the correlation parameters

Example: **zeroconf** `td006.zc`

See documentation on *confFile* (or `.zc` file) in section 7.46. Output data documentation can be found for `.vis` files in section 7.45 and for `.lag` files in section 7.29.

7 Description of various files

In the descriptions that follow, the locations of some files is given as `/home/vlbiobs`, meaning the directory `/home/vlbiobs/astrometry/mmmmyy/project` or one of its subdirectories (this is VLBA-centric). Here *mmmyy* is the month and year of the project's observation (i.e., `jan08`) and *project* is the full project name, with segment, in lower case, such as `bw088n`. In what follows, the “software correlator project directory” (sometimes “project directory”) refers to the directory from which software correlation is to proceed. File names beginning with a period (e.g., `.acb`) represent file name extensions, typically (but not always) to job file bases, such as `job121.000`.

7.1 **.aapd**

The program **apd2antenna** (see Sec. 6.1) takes the `.apd` file (Sec. 7.5 created by **difx2fits** (Sec. 6.14) and performs least-squares fits to reference the phase, delay, and rate measurements to a specified reference antenna.

The first line in the file is **obscode**: followed by the observation code, e.g., `MT831`.

Each subsequent line has the same format with the following fields:

Key	Units/allowed values	Comments
<i>MJD</i>	integer ≥ 1	MJD day number corresponding to line
<i>hour</i>	$\geq 0.0, < 24.0$	hour within day
<i>source name</i>	string	name of source; no spaces allowed
<i>antenna number</i>	integer ≥ 1	antenna table index for first antenna
<i>antenna name</i>	string	name of antenna 1; no spaces allowed
<i>n_{BBC}</i>	integer ≥ 1	number of baseband channels, n_{BBC} The next four columns are repeated n_{BBC} times
<i>delay</i>	ns	the fringe fit delay
<i>phase</i>	degrees	phase of fringe fit peak
<i>rate</i>	Hz	the fringe fit rate

7.2 .abp

When run with the `--bandpass` option, `difx2fits` will create a `.bandpass` file. This file can be run through `bp2antenna` (Sec. 6.3 to convert the baseline-based bandpasses to antenna-based bandpasses.

It is possible that line comments, starting with `#`, are present in the file.

The first line of the file is `obscode:` followed by the project code.

A header line starting with `Bandpass` will signify the start of the data for one antenna for one baseband channel. Such a line has exactly eight space-separated fields:

1. The keyword `Bandpass`
2. Zero-based antenna number for the antenna
3. Uppercase antenna code for the antenna
4. Baseband channel number (zero-based)
5. Number of points in the bandpass, N_{point}
6. Signed Sum of Local Oscillator (SSLO) for the baseband channel frequency (MHz)
7. Baseband channel bandwidth (MHz)
8. Polarization (typically R, L, H, or V)

Following a header line should be the specified number, N_{point} , of data lines, each with three space-separated values:

1. Sky frequency (MHz)
2. Real part of bandpass at this frequency
3. Imaginary part of bandpass at this frequency

Note that the bandpass is represented as the measurements. The correction factor to flatten the bandpass would be the complex-valued reciprocal of these measurements.

7.3 .acb

When generation of sniffer output files is not disabled, each `.FITS` file written by `difx2fits` will be accompanied by a corresponding `.acb` file. This file contains auto-correlation spectra for each antenna for each source. In order to minimize the output data size, spectra for the same source will only be repeated once per 15 minutes. The file contains many concatenated records. Each record has the spectra for all baseband channels for a particular antenna and has the following format. Note that no spaces are allowed within any field. Values in **typewriter** font without comments are explicit strings that are required.

Line(s)	Value	Units	Comments
1	timerange: <i>MJD</i> <i>start time</i> <i>stop time</i> obscode: <i>observe code</i> chans: <i>n_{chan}</i> x <i>n_{BBC}</i>	integer ≥ 1 string string string ≥ 1 ≥ 1	MJD day number corresponding to line e.g., 13h34m22.6s e.g., 13h34m52.0s e.g., MT831 number of channels per baseband channel number of baseband channels
2	source: <i>source name</i> bandw: <i>bandwidth</i> MHz	string MHz	e.g., 0316+413 baseband channel bandwidth
3 to 2+n _{BBC}	bandfreq: <i>frequency</i> GHz polar: <i>polarization</i> side: <i>sideband</i> bbchan: <i>bbc</i>	GHz 2 chars U or L 0	band edge (SSLO) frequency of baseband channel e.g. RR or LL for upper or lower sideband Currently not used but needed for conformity
3+n _{BBC} to 2+n _{BBC} (n _{chan} + 1)	<i>antenna number</i> <i>antenna name</i> <i>channel number</i> <i>amplitude</i>	≥ 1 string ≥ 1 ≥ 0.0	antenna table index = chan + (bbc - 1) · n _{chan} for chan, bbc ≥ 1

The above are repeated for each auto-correlation spectrum record. This file can be plotted directly with `plotbp` or handled more automatically with `difxsniff`.

7.4 .apc

This file type is nearly identical to the better known `.apd` file; the name acronym refers to Amplitude Phase Channel. The amplitude, phase, and rate for the brightest channel is determined for each IF for each solution interval. When generation of sniffer output files is not disabled, each `.FITS` file written by `difx2fits` will be accompanied by a corresponding `.apc` file. This file contains *channel-based* fringe fit solutions typically every 30 seconds for the entire experiment. These solutions are not of calibration quality but are sufficient for use in evaluating the data quality.

The first line in the file is the observation code, e.g., MT831 .

Each subsequent line has the same format with the following fields:

Key	Units/allowed values	Comments
<i>MJD</i>	integer ≥ 1	MJD day number corresponding to line
<i>hour</i>	$\geq 0.0, < 24.0$	hour within day
<i>source number</i>	integer ≥ 1	source table index
<i>source name</i>	string	name of source; no spaces allowed
<i>ant1 number</i>	integer ≥ 1	antenna table index for first antenna
<i>ant2 number</i>	integer ≥ 1	antenna table index for second antenna
<i>ant1 name</i>	string	name of antenna 1; no spaces allowed
<i>ant2 name</i>	string	name of antenna 2; no spaces allowed
<i>nBBC</i>	integer ≥ 1	number of baseband channels, n_{BBC} The next four columns are repeated n_{BBC} times
<i>channel</i>	$\geq 1, \leq n_{\text{chan}}$	the strongest channel
<i>amplitude</i>	≥ 0.0	the amplitude of the peak channel
<i>phase</i>	degrees	phase of the peak channel
<i>rate</i>	Hz	the channel phase rate

7.5 .apd

When generation of sniffer output files is not disabled, each `.FITS` file written by `difx2fits` will be accompanied by a corresponding `.apd` file. This file contains Amplitude, Phase, Delay (hence the name) and rate results from fringe fit solutions typically every 30 seconds for the entire experiment. These solutions are not of calibration quality but are sufficient for use in evaluating the data quality.

The first line in the file is the observation code, e.g., `MT831`.

Each subsequent line has the same format with the following fields:

Key	Units/allowed values	Comments
<i>MJD</i>	integer ≥ 1	MJD day number corresponding to line
<i>hour</i>	$\geq 0.0, < 24.0$	hour within day
<i>source number</i>	integer ≥ 1	source table index
<i>source name</i>	string	name of source; no spaces allowed
<i>ant1 number</i>	integer ≥ 1	antenna table index for first antenna
<i>ant2 number</i>	integer ≥ 1	antenna table index for second antenna
<i>ant1 name</i>	string	name of antenna 1; no spaces allowed
<i>ant2 name</i>	string	name of antenna 2; no spaces allowed
<i>nBBC</i>	integer ≥ 1	number of baseband channels, n_{BBC} The next four columns are repeated n_{BBC} times
<i>delay</i>	ns	the fringe fit delay
<i>amplitude</i>	≥ 0.0	the amplitude of fringe fit peak
<i>phase</i>	degrees	phase of fringe fit peak
<i>rate</i>	Hz	the fringe fit rate

7.6 .bandpass

When run with the `--bandpass` option, `difx2fits` will create a `.bandpass` file. The data in this file is created after applying the results of the fringe fit process and then averaging over all data. This option should only be used when it is expected that all of the data being processed is on a source strong enough for valid fringe fit solutions. The contents of the file complex-valued bandpasses determined on each of the baselines. In its current form (2023/04/24) only cross-correlations are considered, but the option remains open to write real-valued autocorrelations as well.

It is possible that line comments, starting with `#`, are present in the file.

The first line of the file is `obscode:` followed by the project code.

A header line starting with **Bandpass** will signify the start of the data for one baseline for one baseband channel. Such a line has exactly ten space-separated fields:

1. The keyword **Bandpass**
2. Zero-based antenna number for first antenna
3. Zero-based antenna number for second antenna
4. Uppercase antenna code for first antenna
5. Uppercase antenna code for second antenna
6. Baseband channel number (zero-based)
7. Number of points in the bandpass, N_{point}
8. Signed Sum of Local Oscillator (SSLO) for the baseband channel frequency (MHz)
9. Baseband channel bandwidth (MHz)
10. Polarization (typically R, L, H, or V)

Following a header line should be the specified number, N_{point} , of data lines, each with three space-separated values:

1. Sky frequency (MHz)
2. Real part of bandpass at this frequency
3. Imaginary part of bandpass at this frequency

Note that the bandpass is represented as the measurements. The correction factor to flatten the bandpass would be the complex-valued reciprocal of these measurements.

7.7 .binconfig

The **.binconfig** file is a file created by the user of DiFX and referenced by the **.input** file to specify pulsar options. The file uses the standard DiFX input file format and has the following parameters:

Key	Units/allowed values	Comments
NUM POLYCO FILES	integer ≥ 1	Number of polyco files to read ($nPoly$) The next row is duplicated $nPoly$ times
POLYCO FILE p	string	Name of p^{th} polynomial file
NUM PULSAR BINS	integer ≥ 1	Number of pulse bins to create ($nBin$)
SCRUNCH OUTPUT	boolean	Sum weighted bins? If not, write all bins The next rows are duplicated $nBin$ times
BIN PHASE END b	float 0.0-1.0	Pulsar phase where bin ends
BIN WEIGHT b	float ≥ 0.0	Weight to use when scrunching

The start of one bin is equal to the end of the previous bin; bins wrap around phase 1.0. The BIN PHASE END parameters must be listed in ascending phase order. See Sec. 4 for example usage of **.binconfig** files.

7.8 .bootstrap

The `difxbuild` installer program begins its process by building an environment based on the contents of a `.bootstrap` file. In the simplest case only three parameters are required (`version`, `headnode`, and `difxbase`), however installations can be customized through the use of several other parameters. The `.bootstrap` file is a text file containing zero or one *key = value* statements on each line. Comments begin with a `#`.

The parameters specified can include:

- **version:** which version of difx to install. Currently supported values are DIFX-DEVEL, DIFX-2.1, and DIFX-2.2. The `DIFX_VERSION` environment variable will reflect this value. This parameter is required.
- **headnode:** the computer that will be singled out as the head node. The `DIFX_HEAD_NODE` environment variable will reflect this value. This parameter is required.
- **difxbase:** the top level directory for DiFX software. DiFX-version-independent files will be placed directly beneath this directory. By default DiFX version specific files will be installed in a subdirectory of this (see information about the root parameter below). It is okay (and encouraged) to use the same difxbase for all installed versions as this allows common third-party software to be used. This parameter is required.
- **root:** the base directory for DiFX version/label specific files for the primary platform. Secondary platforms will use the same but with a provided extension (see `altplatformX` below). If not provided, this parameter will default to *difxbase* (or *label* if specified).
- **ipproot:** path to the base of the Intel Performance Primitives library. This is IPP version dependent and may require a bit of trial and error to get right. If this is set to none then an IPP-free DiFX will be installed. This requires FFTW to be installed. Each architecture can have its own ipproot value. ipproot specifies the the default; architecture-dependent overrides are specified with a parameter such as `ipprooti686` or `ipprootx86_64`.
- **label:** a label used to identify an installation of DiFX. By default it is set equal to the specified version. Setting it to an alternate value allows multiple installations of the same DiFX version to be later identifiable. The `DIFX_LABEL` environment variable will reflect this value.
- **calcserver:** the computer to send RPC model requests to. If not specified, this will default to the value of the headnode parameter. The `CALC_SERVER` environment variable will reflect this value.
- **cflags:** default c and c++ compiler flags to use. If not specified, the default of `-O2 -Wall -march=core2` will be used.
- **pathextra:** extra binary search paths to add to the `PATH` environment variable that is set in the `setup_difx` script.
- **ldextra:** extra paths to be added to the `LD_LIBRARY_PATH` environment variable that is set in the `setup_difx` script.
- **wrapper:** an optional wrapper program that can be used to spawn `mpifxcorr`. This value gets coded into the `runmpifxcorr.label` launcher script. For example, `valgrind` could be used as the wrapper program if memory leak checking is desired. Use this parameter with caution.
- **mca:** parameters to add to the `/etc/openmpi-mca-params.conf` file. If not provided, no such file will be created. This can be useful to include or exclude certain network interfaces. You can set this on a per platform basis. To do this, for example, set `mcai686` and `mcax86_64` separately.

- **primaryarch**: Normally bootstrapping needs to be done on a machine running on the primary architecture. If **primaryarch** is set, the bootstrapping step can be run on any machine. This should be set to **i686**, **x86_64**, or whatever **"uname -m"** returns on the primary architecture.
- **altplatformX**: Here *X* is a number from 1 to 9. This parameter gives a sub-label to each non-primary platform. Examples might be **SDK8** and **SDK9** for Mark5 units using two different Conduant library versions. For each specified alternate platform the following three additional parameters are needed ...
 - **altplatformXarch**: The CPU architecture, as determined by **"uname -m"**, that this platform is based upon.
 - **altplatformXhost**: A representative computer making use of this platform. This is used when spawning a parallel build process.
 - **altplatformXtest**: A bash conditional expression used to determine if the computer running the **setup_bash** script belongs to this platform. An example is: **x'pkg-config --modversion streamstor' < "x9.0"**

7.9 .cablecal

Cable calibration is used to measure the electrical pathlength of the oscillator signals being sent from the control building to the antenna vertex room. The VLBA was designed to minimize pathlength variations over time, but inevitably some temperature change and stretching due to antenna motion is to be expected. At the VLBA, the program **db2cc** is used to pull cable cal from the VLBA monitor database and format it in a manner that can be read by **difx2fits**. There will be one file per antenna which is given filename of *exp.stn.cablecal*, where *exp* is the experiment code and *stn* is the antenna station code.

The **.cablecal** files are text format files. Comments within the file start with **#**. Valid data lines have four space-separated columns of data:

- station code (in capital letters, usually two characters long)
- MJD timestamp
- integration period for the measurement (measured in seconds, or zero if not specified)
- the cable cal round-trip pathlength (measured in picoseconds)

7.10 cal.vlba

Monitor data that gets attached to FITS files is extracted by **tsm** into a file called *projectcal.vlba* where *project* is the name of the project, i.e., **bg167** or **bc120a**. A single file contains the monitor data for all VLBA antennas, maybe also including GBT, Effelsberg and Arecibo, for the duration of the project. The file is left in **/home/vlbiobs** and is compressed with **gzip** after some time to save disk space, resulting in additional file extension **.gz**. A program called **vlog** (sec §6.104) reads this file and produces files called **flag**, **pcal**, **tsys**, and **weather** in the software correlator project directory. This file type can be read by AIPS task **ANTAB**.

7.11 .calc

The main use of the **.calc** file is to drive the geometric model calculations but this file also serves as a convenient place to store information that is contained in the **.fx** file but not in the **.input** file and is needed for **.FITS** file creation. In the DiFX system, one **.calc** file is created by **vex2difx** (§6.102) for each **.input** file. This file is read by **calcif2** (§6.4) to produce a tabulated delay model, *u, v, w* values, and estimates of atmospheric delay contributions.

In brief, the parameters in this file that are relevant for correlation include time, locations and geometries of antennas, pointing of antennas (and hence delay centers) as a function of time and the Earth orientation parameters relevant for the correlator job in question. Additional parameters that are stuffed into this file include spectral averaging, project name, and information about sources such as calibration code and qualifiers. In the NRAO application of DiFX, source names are faked in the actual `.input` file in order to allow multiple different configurations for the same source. A parameter called *realname* accompanies each source name in the `.calc` file to correctly populate the source file in `.FITS` file creation.

The syntax of this file is similar to that of the `.input` file. The file consists entirely of key-value pairs separated by a colon. The value column is not constrained to start in column 21 as it is for the files used by `mpifxcorr`. There are five sections in the `.calc` file; these sections are not separated by any explicit mark in the file.

The first section contains values that are fixed for the entire experiment and at all antennas; all data in this section is scalar. In the following table, all numbers are assumed to be floating point unless further restricted. The keys and allowed values in this section are summarized below. Optional keys are identified with a $\star$. Deprecated keys that will likely be removed in an upcoming version are identified with an $\times$.

Key	Units/allowed values	Comments
JOB ID	integer ≥ 1	taken from <code>.fx</code> file
$\star$ JOB START TIME	MJD + fraction	start time of original <code>.fx</code> file
$\star$ JOB STOP TIME	MJD + fraction	end time of original <code>.fx</code> file
$\star$ DUTY CYCLE	float ≤ 1	fraction of the job contained within scans
OBSCODE	string	observation code assigned to project
$\star$ SESSION	short string	session suffix to OBSCODE, e.g., A or BE
$\star$ DIFX VERSION	string	version of correlator, e.g. DIFX-1.5
$\star$ DIFX LABEL	string	name of correlator install, e.g. DIFX-WALTER
VEX FILE	string	dir/filename of vex file used to create the job
START MJD	MJD + fraction	start time of this subjob
START YEAR	integer	calendar year of START MJD
START MONTH	integer	calendar month of START MJD
START DAY	integer	day of calendar month of START MJD
START HOUR	integer	hour of START MJD
START MINUTE	integer	minute of START MJD
START SECOND	integer	second of START MJD
$\star$ SPECTRAL AVG	integer ≥ 1	number of channels to average in FITS creation
$\star$ START CHANNEL	integer ≥ 0	start channel number (before averaging)
$\star$ OUTPUT CHANNELS	integer ≥ 1	total number of channels to write to FITS
	$> 0.0, < 1.0$	fraction of total channels to write to FITS
$\star$ TAPER FUNCTION	string	currently only UNIFORM is supported

The second section contains antenna(telescope) specific information. After an initial parameter defining the number of telescopes, there are $nTelescope$ sections (one for each antenna), each with the following six parameters. Lowercase t in the table below is used to indicate the telescope index, an integer ranging from 0 to $nTelescope - 1$. Note that in cases where units are provided under the Key column, these units are actually part of the key.

Key	Units/allowed values	Comments
NUM TELESCOPES	integer ≥ 1	number of telescopes ($nTelescope$). The rows below are duplicated $nTelescope$ times.
TELESCOPE t NAME	string	upper case antenna name abbreviation
TELESCOPE t MOUNT	string	the mount type: altz, equa, xyew, or xyns
TELESCOPE t OFFSET (m)	meters	axis offset in meters
TELESCOPE t X (m)	meters	X geocentric coordinate of antenna at date
TELESCOPE t Y (m)	meters	Y geocentric coordinate of antenna at date
TELESCOPE t Z (m)	meters	Z geocentric coordinate of antenna at date
★ TELESCOPE t SHELF	string	shelf location of module to correlate

The third section contains a table of sources. Sources are indexed from the following section describing the scans.

Key	Units/allowed values	Comments
NUM SOURCES	integer ≥ 1	number of sources ($nSource$) The rows below are duplicated $nSource$ times.
SOURCE s NAME	string	name of source (possibly renamed from <code>.vex</code>)
SOURCE s RA	radians	J2000 right ascension
SOURCE s DEC	radians	J2000 declination
SOURCE s CALCODE	string	usually upper case letters or blank
SOURCE s QUAL	integer ≥ 0	source qualifier

The fourth section contains scan specific information. Except for one initial line specifying the number of scans, $nScan$, this section is composed of nine parameters per scan. Each parameter is indexed by s which ranges from 0 to $nScan - 1$.

Key	Units/allowed values
NUM SCANS	integer ≥ 1
SCAN s IDENTIFIER	string
SCAN s START (S)	seconds
SCAN s DUR (S)	seconds
SCAN s OBS MODE NAME	string
SCAN s UVSHIFT INTERVAL (NS)	time to integrate before doing uv shifts (used mainly for multi-phase-center observ)
SCAN s AC AVG INTERVAL (NS)	averaging interval for export of fast-dump spectra (used for VFASTR)
SCAN s POINTING SRC	integer ≥ 1
SCAN s NUM PHS CTRS	integer ≥ 1
SCAN s PHS CTR p	integer ≥ 1

The fifth section contains Earth orientation parameters (EOP). Except for one initial line specifying the number of days of EOPs, $nEOP$, this section is composed of five parameters per day of sampled EOP values. Each parameter is indexed by e which ranges from 0 to $nEOP - 1$.

Key	Units/allowed values	Comments
NUM EOP	integer ≥ 1	number of tabulated EOP values ($nEOP$) The rows below are duplicated $nEOP$ times.
EOP e TIME (MJD)	MJD + fraction	time of sample; fraction almost always zero
EOP e TAL.UTC (sec)	integer seconds	leap seconds accrued at time of job start
EOP e UT1.UTC (sec)	seconds	UT1 - UTC
EOP e XPOLE (arcsec)	arc seconds	X coordinate of polar offset
EOP e YPOLE (arcsec)	arc seconds	Y coordinate of polar offset

The next (completely optional) section has a table for positions and velocities of spacecraft. Each spacecraft is indexed by s and each row thereof by r .

Key	Units/allowed values	Comments
* NUM SPACECRAFT	integer ≥ 0	number of spacecraft ($nSpacecraft$) Everything below is duplicated $nSpacecraft$ times.
SPACECRAFT s NAME	string	name of spacecraft
SPACECRAFT s ROWS	integer ≥ 1	number of data rows, $nRow_s$ for spacecraft s The row below is repeated $nRow_s$ times.
SPACECRAFT s ROW r	7 numbers	tabulated data; see below

Each data vector of data consists of seven double precision values: time (mjd), x , y , and z (meters), and $\dot{x}$, $\dot{y}$, and $\dot{z}$ (meters per second). These values should be separated by spaces.

The final section identifies the files to be produced.

Key	Units/allowed values	Comments
IM FILENAME	string	dir/filename of <code>.im</code> file to create
FLAG FILENAME	string	dir/filename of <code>.flag</code> file to create

7.12 .difx/

The SWIN format visibilities produced by `mpifxcorr` are written to a directory with extension `.difx`. Three kinds of files can be placed in this directory as described below.

Note that the formats and naming conventions of these files is not guaranteed to stay unchanged from version to version of DiFX, and hence it is not recommended to rely on these files for archival purposes.

7.12.1 Visibility files

The bulk of the output from `mpifxcorr` is usually in the form of a binary visibility file. Usually there will be a single visibility file in this directory, but there are three ways in which multiple files may be produced: 1. a restart of the correlation, 2. if there are multiple phase centers, and 3. if there are multiple pulsar bins.

The visibility files are systematically named in the form: `DIFX_day_sec.ssrc.bbin`, where `day` is the 5 digit integer MJD of the start of visibilities, `sec` is a zero-padded 6 digit number of seconds since the MJD midnight, `src` is a 4 digit zero-padded integer specifying the phase center number (starting at 0), and `bin` is a 4 digit zero-padded integer specifying the pulsar bin number (starting at 0).

These files contain visibility data records. Each record contains the visibility spectrum for one polarization of one baseband channel of one baseline for one integration time. Each starts with a binary header and is followed by binary data.

The format of the header is shown in the table below.

Key	data type	units	comments
baseline number	int		$= (a_1 + 1) * 256 + (a_2 + 1)$ for $a_1, a_2 \geq 1$
day number	int	MJD	date of visibility centroid
seconds	double	sec	vis. centroid seconds since beginning of MJD
config index	int	≥ 0	index to <code>.input</code> file configuration table
source index	int	≥ 0	index to <code>.calc</code> file scan number
freq index	int	≥ 0	index to <code>.input</code> frequency table
antenna 1 polarization	char	R, L, X, Y	
antenna 2 polarization	char	R, L, X, Y	
pulsar bin number	int	≥ 0	
visibility weight	double	≥ 0.0	data weight for spectrum; typically ~ 1
u	double	meter	u component of baseline vector
v	double	meter	v component of baseline vector
w	double	meter	w component of baseline vector

Note that for both the header and the data, the endianness is native to the machine running `mpifxcorr`, and there are currently no provisions for processing such files on a machine of different endianness.

Following the end-of-line mark for the last header row begins binary data in the form of (real, imaginary) pairs of 32-bit floating point numbers. The `.input` file parameter `NUM CHANNELS` indicates the number of complex values to expect. In the case of upper sideband data, the first reported channel is the “zero frequency” channel, that is its sky frequency is equal to the value in the frequency table for this spectrum. The Nyquist channel is not retained. For lower sideband data, the last channel is the “zero frequency” channel. That is, in all cases, the spectrum is in order of increasing frequency and the Nyquist channel is excised.

7.12.2 Pulse cal data files

Pulse calibration data can be extracted by `mpifxcorr`. Extraction is configured on a per-antenna basis. Data for each antenna is written to a separate file; if correlation is restarted, an additional pulse cal data file will be written.

The pulse cal data files are systematically named in the form: `PCAL_day_sec_ant`, where *day* is the 5 digit integer MJD of the start of visibilities, *sec* is a zero-padded 6 digit number of seconds since the MJD midnight, and *ant* is the 1 or 2 letter antenna name in capital letters. There is potential for these text files to have very long lines (more than 10,000 bytes) when many pulse cal tones are extracted.

For DiFX versions 2.3 and earlier the data format was exactly the same as documented in §7.35. This old version will be considered “version 0”.

The data format being used now is similar in spirit but more convenient for `mpifxcorr` to produce and for `difx2fits` and `difx2mark4` to digest leading to broader support (in theory complete) of the various polarization, frequency, and sideband combinations allowed by DiFX. The data format is as follows:

Comment lines begin with an octothorpe (`#`). The first few lines of comments may contain machine-readable information in the following format:

```
# DiFX-derived pulse cal data
# File version = 1
# Start MJD =
# Start seconds =
# Telescope name =
```

Data lines always contain 6 fixed-size fields:

1. *antId* : Station name abbreviation, e.g., LA
2. *day* : Time centroid of measurement (MJD, including fractional portion)
3. *dur* : Duration of measurement (days)
4. *datastreamId* : The datastream index of for this data.
5. *nRecBand* : Number of recorded baseband channels
6. *nTone* : (Maximum) number of pulse cal tones detected per band per polarization

Following these fields is a variable-length arrays of numbers. This array contains the pulse cal data and consists of *nRecBand***nTone* groups of four numbers. The groups are arranged in ascending record band index (slow index) and ascending tone number (fast index) where the tone number increases away from the reference frequency; not sure what happens with dual-sideband complex! The first member of this group is the tone frequency (MHz), or -1 to indicate there was not a measurment. The second member of this group is the polarization, one of R, L, X or Y. The third and fourth are respectively the real and imaginary parts of the tone measured at the given sky frequency.

7.12.3 Switched power files

`mpifxcorr` can be used to extract switched power from individual antennas. Extraction is configured on a per-datastream (usually the same as per-antenna) basis. Data for each data stream is written to a separate file; if correlation is restarted, an additional set of switched power files will be started.

The switched power files are systematically named in the form: `SWITCHEDPOWER_day_sec_ds`, where *day* is the 5 digit integer MJD of the start of visibilities, *sec* is a zero-padded 6 digit number of seconds since the MJD midnight, and *ds* is the datastream index as set in the `.input` file. The test lines in these files can be long (more than 1000 bytes).

The format of these files is as follows. Each line of the file represents all measurements made on one datastream at over one integration period. The lines contain the following columns:

1. *mjdstart* : The start of the integration period in mjd+fraction
2. *mjdstop* : The end of the integration period in mjd+fraction

These are then followed by 4 numbers for each recorded channel:

1. P_{on} : power in the “on” state
2. $\sigma_{P_{\text{on}}}$: uncertainty of the power in the “on” state
3. P_{off} : power in the “off” state
4. $\sigma_{P_{\text{off}}}$: uncertainty of the power in the “off” state

The magnitudes of the numbers are meaningless but their ratios have meaning. Text after a comment character (`#`) are ignored.

7.13 .difxlog

The `difxlog` program (§6.23) captures `DifxAlertMessage` and `DifxStatusMessage` message types that are sent from an ongoing software correlation process and writes the information contained within to a human readable text file. One line of text is produced for each received message. The first five columns contain the date and time in `ddd MMM dd hh:mm:ss yyyy` format (e.g., `Wed Apr 22 12:48:41 2009`). The sixth column contains a word describing the contents of the remainder of the line: Options are:

`STATUS` : The status of the process is described

`WEIGHTS` : The playback weights for each antenna are listed

other : This word represents an alert severity level (one of `FATAL`, `SEVERE`, `ERROR`, `WARNING`, `INFO`, `VERBOSE` and `DEBUG`) and is followed by the alert message itself.

7.14 .speed

The program `difxspeed` (§6.26) runs a set of performance benchmarks described by a `.speed` file as documented here. This file is a text file containing various parameters. There are 5 required parameters:

Parameter	Comments
<code>datastreams</code>	comma separated list of nodes on which to run datastream processes
<code>cores</code>	comma separated list of nodes on which to run core processes
<code>antennas</code>	list of 1 or 2 letter antenna names to process
<code>nThread</code>	one or more values listing number of threads to use (see below)
<code>vex</code>	vex file to use as descriptor of observation

The value of `nThread` applies to all core processes; if multiple comma-separated values are specified, these will result in additional runs of benchmarking.

Other parameters that can be specified, either as single values or as arrays to be used in full combination with all other value arrays, include:

Parameter	Comments
<code>nAnt</code>	number of antennas (in order of listed antennas, starting with first listed)
<code>nCore</code>	number of core processes to start, using in order <code>cores</code>
<code>tInt</code>	integration time (seconds)
<code>specRes</code>	spectral resolution (MHz)
<code>fftSpecRes</code>	resolution of transform
<code>xmacLength</code>	cross-multiply stride size
<code>strideLength</code>	fringe rotation stride size
<code>numBufferedFFTs</code>	number of FFTs to process in one go
<code>toneSelection</code>	pulse cal tone selection

Notes:

1. `datastreams` and `cores` lists can repeat hostnames.
2. Some combinations of parameters is illegal; at the moment it is up to the user to ensure all combinations of values are allowed.
3. Additional parameters can be easily added to the program on demand.

7.15 .speed.out

The output of `difxspeed` (§6.26) is a file usually ending in `.speed.out`. The first many lines are comments describing to a human reader the fixed parameters of the benchmark trials and a table describing the meanings of the columns of the uncommented data lines that follow. In summary, the columns in the lines that follow are:

Column(s)	value(s)
1 to N	values of variable parameters as described by comments above
$N + 1$	The average execution time of all trials run with the combination of parameters
$N + 2$	The RMS scatter in execution time
$N + 3$ to $N + 2 + R$	List of execution times from all trials

In the above table, N is the number of parameters taking on multiple values and R is the number of times `difxspeed` was run.

7.16 \$DIFX_MACHINES

This section describes the format of a file used through DiFX-2.2. For more recent versions please see documentation on the DiFX wiki <http://www.atnf.csiro.au/vlbi/dokuwiki/doku.php/difx/start/>.

Environment variable `DIFX_MACHINES` should point to a file containing a list of machines that are to be considered elements of the software correlator. Program `genmachines` (§6.39) uses this file and information within a `.input` file to populate the `.machines` file needed by `mpifxcorr`. Because usually only one node in a cluster has direct access to a particular Mark5 module (or data from that module), the ordering of computer names in the `.machines` file is important. Rows in the `$DIFX_MACHINES` file contain up to three items, the last one being optional. The first column is the name of the machine. The second column is the number of processes to schedule on that machine (typically the number of CPU cores). The third column is a 1 if the machine is a Mark5 unit and 0 otherwise. If this column is omitted, the machine will be assumed to be a Mark5 unit if the first 5 characters of the computer name are `mark5`, and will be assumed not to be otherwise. Comments in this file begin with an octothorpe (`#`). Lines with fewer than two columns (after excision of comments) are ignored.

7.17 .dir

Reading directory information off Mark5 modules can take a bit of time (measured in minutes usually). Since the same modules are often accessed multiple times, the directories are cached in `$MARK5_DIR_PATH/`. In this directory, there will be one file per module that has been used, named `VSN.dir`, where `VSN` is the volume serial number of the module, e.g., NRAO-023. The format of these files is as follows: The first line contains three fields: `VSN`, the number of scans on the module, `nScan`, and either `A` or `B` indicating the last bank the module was installed in. At the end of this line the characters `RT` can be added (by hand) which will cause the modules to be accessed using *Real-Time* mode which is tolerant of missing or bad disks within the module. Then there are `nScan` rows containing information about each scan, each with 11 columns. Values are floating point unless otherwise noted.

Key	Units/allowed values	Comments
Start byte	64-bit integer bytes	offset of the scan on the Mark5 module
Length	64-bit integer bytes	length of the scan
Start day	integer MJD	the modified Julian day of the scan start
Start time	integer seconds	the scan start time
Frame num	integer	frame number since last second tick
Frames per sec	integer	number of frames per second
Scan Duration	seconds	the duration of the scan
Frame size	integer bytes	the length of one data frame, including headers
Frame offset	integer bytes	the offset to the start of the first entire frame
Tracks	integer	the number of data tracks
Format	integer	0 for VLBA format, 1 for Mark4 format, 2 for Mark5B
Name	string	scan name, usually including the project code and station
Extra info	string(s)	see below

After the name of the scan additional free-form text can appear. These extra parameters can be machine parsable. The only use of this as of this writing is to indicate the thread ids for VDIF data. This will always be formatted as `Threads=` followed by a comma separated list of thread ids without any spaces. For example: `Threads=0,640,256,896`.

7.18 .filelist

When using the `filelist` parameter in an `ANTENNA` section of a `.v2d` (§7.42), the list of data files to correlate are stored in a text file. This is a text file containing data lines and optionally comments. Any text after the comment character (`#`) is ignored. A data line consists of a filename (must have complete path as can be used to find the file on the datastream node for this antenna) and optionally a start time and stop time. Start and stop times can be expressed in any of the formats supported by `vex2difx`.

7.19 .FITS

The `.FITS` files discussed here are produced by `difx2fits`. They aim to conform to the same table structures as the `FITS-IDI` files produced by the VLBA correlator. The format is described in AIPS Memo 102, “The `FITS` Interferometry Data Interchange Format”, however, this memo is a bit out of date and the data structures described are not in exact agreement with those made by the VLBA correlator; in all cases the format of data produced by the VLBA hardware correlator is favored where the two disagree. The tables in these `FITS` files have a nearly 1 to 1 relationship with the tables that are seen within AIPS, though their two letter abbreviations differ. The following tables are produced by `difx2fits`:

Table	Description
AG	The array geometry table
SU	The source table
AN	The antenna table
FR	The frequency table
ML	The model table
CT	The correlator (eop) table
MC	The model components table
SO	The spacecraft orbit table
UV	The visibility data table
FG	The flag table
TS	The system temperature table
PH	The phase calibration table (pulse cals and state counts)
WR	The weather table
GN	The gain curve table
GM	The pulsar gate model table

Not all of these tables will always be written.

7.20 .fitslist

A `.fitslist` file is written by `makefits` and contains the entire list of `.FITS` files for the correlator pass. Due to the different constraints of the correlation process and the FITS-IDI format, the number of resultant FITS files may be greater or less than the number of jobs. This file type is used by `difxarch` to ensure that all of the correlated output ends up in the archive. The file is composed of two parts: a header line and one line for each `.FITS` file. The header line consists of a series of *key=value* pairs. Each *key* and *value* must have no whitespace and no whitespace should separate these words from their connecting = sign. While any number of key-value pairs may be specified, the following ones (which are case sensitive) are expected to be present:

1. `exper` : the name of the experiment, including the segment code
2. `pass` : the name of the correlator pass
3. `jobs` : the name of the `.joblist` file used by `makefits`
4. `mjd` : the modified Julian day when `makefits` created this file
5. `DiFX` : the version name for the DiFX deployment (the value of `$DIFX_VERSION` when `vex2difx` was run)
6. `difx2fits` : the version of `difx2fits` that was run

Each additional line contains information for one `.FITS` file of the correlation pass. These lines contain three fields:

1. *archiveName* : the name of the file that will get injected into the archive (see §??)
2. *fileSize* : the size of the file in MB
3. *origName* : the name of the file as produced by `difx2fits` (via `makefits`)

7.21 .flag

The program `vex2difx` may write a `.flag` file for each `.input` file it creates. This file is referenced from the `.calc` file. This flag file is used by `difx2fits` to exclude nonsense baselines that might have been correlated. Data from nonsense baselines can occur in DiFX output when multiple subarrays are coming and going. The flag file instructs `difx2fits` to drop these data during conversion to FITS-IDI. The format of this text file is as follows. The first line contains an integer, n , which is the number of flag lines that follow. The next n lines each have three numbers: MJD_1 , MJD_2 and ant . The first two floating point numbers determine the time range of the flag in Modified Julian Days. The last integer number is the antenna number to flag, a zero-based index corresponding to the `TELESCOPE` table of the corresponding `.input` file.

7.22 .<antId>.flag

A series of files called `.<antId>.flag` are created when program `vlog` operates on the `cal.vlba` file. These files contain lists of antenna-based flags generated by the on-line system that should be propagated into the FITS FL table. These flag files contains two kinds of lines. Comment lines begin with an octothorpe (`#`) and contain no vital information. Flag lines always consist of exactly 5 fields:

1. *antId* : Station name abbreviation, e.g., LA; also part of the file name.
2. *start* : Beginning of flagged period (day of year, including fractional portion; or Modified Julian Days)
3. *end* : End of flagged period (day of year, including fractional portion; or Modified Julian Days)
4. *recChan* : Record channel affected; -1 for all record channels, otherwise a zero-based index
5. *reason* : Reason for flag, enclosed in single quotes, truncated to 24 characters

The flag rows are sorted by start time. The flags are propagated into the FL table. Visibility data are not altered. Special VLBA reason codes recognized by `difx2fits` are as follows: `'recorder'`, the flag entry will be ignored and is not propagated into FITS, and `'observingsystemidle'`, the value of MJD_2 is ignored and replaced by the ending MJD of the observation.

7.23 .channelflags

User scripts or the program `vex2difx` may write a `.channelflags` file for each `.input` file. `Difx2fits` will propagate the user flags of bad spectral channels into the FITS flag table. This file contains two kinds of lines. Comment lines begin with an octothorpe (`#`) and contain no vital information. Flag lines always consist of 7 fields:

1. *antId* : Station name abbreviation, e.g., LA
2. *start* : Beginning of flagged period (Day of Year or Modified Julian Day; including fractional portion)
3. *end* : End of flagged period (Day of Year or Modified Julian Day; including fractional portion)
4. *freqIndex* : The DiFX frequency to flag, a zero-based index corresponding to a frequency in the `FREQ TABLE` of the corresponding `.input` file.
5. *startCh* : The first spectral channel to flag, a zero-based index.
6. *endCh* : The last spectral channel to flag, a zero-based index.
7. *reason* : Reason for flag, enclosed in single quotes, truncated to 24 characters.

`Difx2fits` translates the DiFX frequency *freqIndex* to the corresponding FITS IF, and flags the specified channels in *all polarizations* of that IF. These flags are propagated into the FL table. Visibility data are not altered.

7.24 flag

A file called **flag** is created when program **vlog** operates on the **cal.vlba** file. The file contains a list of antenna-based flags generated by the on-line system that should be propagated into the FITS FL table. The file is an experiment-wide flag file and effectively a concatenation of **.<antId>.flag** files. This file contains two kinds of lines. Comment lines begin with an octothorpe (**#**) and contain no vital information. Flag lines always consist of exactly 5 fields:

1. *antId* : Station name abbreviation, e.g., **LA**
2. *start* : Beginning of flagged period (Day of Year or Modified Julian Day; including fractional portion)
3. *end* : End of flagged period (Day of Year or Modified Julian Day; including fractional portion)
4. *recChan* : Record channel affected; -1 for all record channels, otherwise a zero-based index
5. *reason* : Reason for flag, enclosed in single quotes, truncated to 24 characters

The flag rows are sorted first by antenna, and then start time. Special VLBA reason codes recognized by **difx2fits** are as follows: *'recorder'*, the flag entry will be ignored and is not propagated into FITS, and *'observing system idle'*, the value of MJD_2 is ignored and replaced by the ending MJD of the observation.

7.25 .im

The **.im** file contains polynomial models used by **difx2fits** in the creation of FITS files. After a header that is similar to that of a **.rate** file, the contents are organized hierarchically with scan number, sub-scan interval, and antenna number being successively faster-incrementing values. The keys and allowed values in this section are summarized below: Note that the values of the delay polynomials in this file have the opposite sign as compared to those generated by **CALC** and those stored in **.FITS** files. Keys preceded by ***** are optional. Note that all polynomials are expanded about their **MJD**, **SEC** start time and use seconds as the unit of time.

Key	Units/allowed values	Comments
★ CALC SERVER	string	name of the calc server computer used
★ CALC PROGRAM	integer	RPC program ID of the calc server used
★ CALC VERSION	integer	RPC version ID of the calc server used
START YEAR	integer	calendar year of START MJD
START MONTH	integer	calendar month of START MJD
START DAY	integer	day of calendar month of START MJD
START HOUR	integer	hour of START MJD
START MINUTE	integer	minute of START MJD
START SECOND	integer	second of START MJD
POLYNOMIAL ORDER	2, 3, 4 or 5	polynomial order of interferometer model <i>on</i>
INTERVAL (SECS)	integer	interval between new polynomial models
ABERRATION CORR	$\begin{cases} \text{UNCORRECTED} \\ \text{APPROXIMATE} \\ \text{EXACT} \end{cases}$	level of u, v, w aberration correction
NUM TELESCOPES	integer ≥ 1	number of telescopes ($nTelescope$) The row below is duplicated $nTelescope$ times.
TELESCOPE t NAME	string	upper case antenna name abbreviation
NUM SCANS	integer ≥ 1	number of scans ($nScan$). Everything below is duplicated $nScan$ times.
SCAN s POINTING SRC	string	name of source used as pointing center
SCAN s NUM PHS CTRS ≥ 1	number of phase centers this scan (nPC_s)	Everything below is duplicated nPC_s times.
SCAN s PHS CTR p SRC	string	name of source defining this phase center
SCAN s NUM POLY	≥ 1	number of polynomials covering scan ($nPoly_{s,p}$) Everything below is duplicated $nPoly_{s,p}$ times.
SCAN s POLY p MJD	integer ≥ 0	the start MJD of this polynomial
SCAN s POLY p SEC	integer ≥ 0	the start sec of this polynomial
ANT a DELAY (us)	$order+1$ numbers	Everything below is duplicated $nTelescope$ times. terms of delay polynomial
ANT a DRY (us)	$order+1$ numbers	terms of dry atmosphere
ANT a WET (us)	$order+1$ numbers	terms of wet atmosphere
★ ANT a AZ	$order+1$ numbers	azimuth polynomial (deg)
★ ANT a EL GEOM	$order+1$ numbers	geometric (encoder) elevation (deg)
★ ANT a EL CORR	$order+1$ numbers	refraction corrected elevation (deg)
★ ANT a PAR ANGLE	$order+1$ numbers	parallactic angle (deg)
ANT a U (m)	$order+1$ numbers	terms of baseline u
ANT a V (m)	$order+1$ numbers	terms of baseline v
ANT a W (m)	$order+1$ numbers	terms of baseline w

7.26 .input

This section describes the `.input` file format used by `mpifxcorr` to drive correlation. Because NRAO-DiFX 1.0 uses a non-standard branch of `mpifxcorr` some of the data fields will differ from those used in the official version, either in parameter name or in the available range of values. Currently the parameters must be in the order listed here. To get the most out of this section it is advisable to look at an actual file while reading. An example file is stashed at <http://www.aoc.nrao.edu/~wbrisken/NRAO-DiFX-1.1/>. In the tables below, numbers are assumed to floating point unless otherwise stated.

Note that the input file format has undergone a few minor changes since NRAO-DiFX version 1.0.

7.26.1 Common settings table

Below are the keywords and allowed values for entries in the common settings table. This table begins with header

```
# COMMON SETTINGS ##!
```

This is always the first table in a `.input` file.

Key	Units/allowed values	Comments
CALC FILENAME	string	name and full path to <code>.calc</code> file
CORE CONF FILENAME	string	name and full path to <code>.threads</code> file
EXECUTE TIME (SEC)	integer seconds	observe time covered by this <code>.input</code> file
START MJD	integer MJD	start date
START SECONDS	integer seconds	start time
ACTIVE DATASTREAMS	integer ≥ 2	number of antennas ($nAntenna$)
ACTIVE BASELINES	integer ≥ 1	number of baselines to correlate ($nBaseline$)
VIS BUFFER LENGTH	integer ≥ 1	the number of concurrent integrations to allow
OUTPUT FORMAT	boolean	always <code>SWIN</code> here
OUTPUT FILENAME	string	name of output <code>.difx</code> directory

Typically, $nBaseline = nAntenna \cdot (nAntenna - 1)/2$. Autocorrelations are not included in this count.

7.26.2 Configurations table

Below are the keywords and allowed values for entries in the configurations table. This table begins with header

```
# CONFIGURATIONS ###!
```

Two indexes are used for repeated keys. The index over datastream (antenna) is d , running from 0 to $nAntenna - 1$ and the index over baseline is b , running from 0 to $nBaseline - 1$.

Key	Units/allowed values	Comments
NUM CONFIGURATIONS	integer ≥ 1	number of modes in file ($nConfig$)
CONFIG NAME	string	name of configuration
INT TIME (SEC)	seconds	integration time
SUBINT NANOSECONDS	nanosec	amount of time to process as one subintegration
GUARD NANOSECONDS	nanosec ≥ 0	amount of extra data to send for overlap
FRINGE ROTN ORDER	int	0 is post-FFT, 1 is delay/rate, ...
ARRAY STRIDE LENGTH	int	used for optimized fringe rotation calculations
XMAC STRIDE LENGTH	int	number of channels to cross multiply in one batch (must evenly divide)
NUM BUFFERED FFTS	int	number of FFTs to cross-multiply in one batch
WRITE AUTOCORRS	boolean	enable auto-correlations; <code>TRUE</code> here
PULSAR BINNING	boolean	enable pulsar mode
PULSAR CONFIG FILE	string	(only if <code>BINNING</code> is <code>True</code>) see § 7.7
PHASED ARRAY	boolean	set to <code>FALSE</code> (placeholder for now)
DATASTREAM d INDEX	integer ≥ 0	DATASTREAM table index, starting at 0
BASELINE b INDEX	integer ≥ 0	BASELINE table index, starting at 0

7.26.3 Rule table

The rule tables describes which configuration will be applied at any given time. Usually this filters on scan attributes such as source, but can also be done in a time-based manner (start and stop times). An time for

which no configuration matches will not be correlated. If more than one rule matches a given time, they must all refer to the same configuration.

This table begins with header

```
# RULES #####!
```

The table below uses r to represent the rule index, which ranges from 0 to $nRule - 1$.

Key	Units/allowed values	Comments
RULE r CONFIG NAME	string	name to associate with this rule
* SOURCE	string	source to match
* SCAN ID	string	scan name to match
* CALCODE	string	cal code to match
* QUAL	string	source qualifier to match
* MJD START	string	earliest time to match
* MJD STOP	string	latest time to match

7.26.4 Frequency table

Below are the keywords and allowed values for entries in the frequency table which defines all possible sub-bands used by the configurations in this file. Each sub-band of each configuration is mapped to one of these through a value in the datastream table (§7.26.6). Each entry in this table has three parameters which are replicated for each frequency table entry. This table begins with header

```
# FREQ TABLE #####!
```

The table below uses f to represent the frequency index, which ranges from 0 to $nFreq - 1$ and t to represent pulse cal tone index, which ranges from 0 to $nTone_f$.

Key	Units/allowed values	Comments
FREQ ENTRIES	integer ≥ 1	number of frequency setups ($nFreq$)
FREQ (MHZ) f	MHz	sky frequency at band edge
BW (MHZ) f	MHz	bandwidth of sub-band
SIDEBAND f	U or L	net sideband of sub-band
NUM CHANNELS f	integer ≥ 1	initial number of channels (FFT size, $nFFT$, is twice this)
CHANS TO AVG f	integer ≥ 1	average this many channels before generating output spectra)
OVERSAMPLE FAC. f	integer ≥ 1	total oversampling factor of baseband data
DECIMATION FAC. f	integer ≥ 1	portion of oversampling to handle by decimation
PHASE CALS f OUT	integer ≥ 0	number of phase calcs to produce ($nTone_f$)
PHASE CAL f/t INDEX	integer	The row below is duplicated $nTone_f$ times. tone number of band

7.26.5 Telescope table

Below are the keywords and allowed values for entries in the telescope table which tabulates antenna names and their associated peculiar clock offsets, and the time derivatives of these offsets. Much of the other antenna-specific information is stored in the datastream table (§7.26.6). Each datastream of each configuration is mapped to one of these through a value in the datastream table. Each entry in this table has three parameters which are replicated for each telescope table entry. This table begins with header

```
# TELESCOPE TABLE ##!
```

The table below uses a to represent the antenna index, which ranges from 0 to $nAntenna - 1$ and c to represent clock coefficient, ranging from 0 to $nCoeff_a$.

Key	Units/allowed values	Comments
TELESCOPE ENTRIES	integer ≥ 1	number of antennas ($nAntenna$)
TELESCOPE NAME a	string	abbreviation of antenna name
CLOCK REF MJD a	double	date around which the following polynomial is expanded
CLOCK POLY ORDER a	int ≥ 0	polynomial order of telescope clock model ($nCoeff_a$)
CLOCK COEFF a/c	$\mu\text{sec}/\text{sec}^c$	clock model polynomial coefficient

7.26.6 Datastream table

The datastream table begins with header

```
# DATASTREAM TABLE #!
```

The table below uses f to represent recorded frequencies, which ranges from 0 to $nFreq - 1$. A second index, z , is used to iterate over zoom bands, ranging from 0 to $nFreq - 1$. A third index, i , is used to cover the range 0 to $nBB - 1$, where the total number of basebands is given by $nBB \equiv \sum_f nPol_f$. In the DiFX system, all sub-bands must have the same polarization structure, so $nBB = nFreq \cdot nPol$. This index is reused for the zoom bands in an analogous manner.

Key	Units/allowed values	Comments
DATASTREAM ENTRIES	integer ≥ 1	number of antennas ($nDatastream$)
DATA BUFFER FACTOR	integer ≥ 1	
NUM DATA SEGMENTS	integer ≥ 1	
TELESCOPE INDEX	integer ≥ 0	telescope table index of datastream
TSYS	Kelvin	if zero (normal in NRAO usage), don't scale data by $tsys$
DATA FORMAT	string	data format
QUANTISATION BITS	integer ≥ 1	bits per sample
DATA FRAME SIZE	integer ≥ 1	bytes in one frame(or file) of data
DATA SAMPLING	string	REAL or COMPLEX
DATA SOURCE	string	FILE (see §??), MODULE for Mark5 playback, or FAKE for benchmark
FILTERBANK USED	boolean	currently only FALSE
PHASE CAL INT (MHZ)	int	pulse cal comb frequency spacing, or 0 if no pulse cal tones
NUM RECORDED FREQS	integer ≥ 0	number of different frequencies recorded for this datastream
REC FREQ INDEX f	integer ≥ 0	index to frequency table
CLK OFFSET f (us)	μsec	frequency-dependent clock offset
FREQ OFFSET f (us)	μsec	frequency-dependent LO offset
NUM REC POLS f	1 or 2	for this recorded frequency, the number of polarizations
REC BAND i POL	R or L	polarization identity
REC BAND i INDEX	integer ≥ 1	index to frequency setting array above; nBB per entry
NUM ZOOM FREQS	integer ≥ 0	number of different zoom bands set for this datastream
ZOOM FREQ INDEX z	integer ≥ 0	index to frequency table
NUM ZOOM POLS z	1 or 2	for this recorded frequency, the number of polarizations
ZOOM BAND i POL	R or L	polarization identity
ZOOM BAND i INDEX	integer ≥ 1	index to frequency setting array above; nBB per entry

7.26.7 Baseline table

In order to retain the highest level of configurability, each baseline can be independently configured at some level. This datastream table begins with header

```
# BASELINE TABLE ###!
```

The baseline table has multiple entries, each one corresponding to a pair of antennas, labeled **A** and **B** in the table. For each of $nBaseline$ baseline entries, $nFreq$ sub-bands are processed, and for each a total of $nProd$ polarization products are formed. Indexes for each of these dimensions are b , f and p respectively, each starting count at 0. Within the DiFX context, all baselines must have the same $nFreq$ and $nProd$, though this is not a requirement of `mpifxcorr` in general. Each of the $nFreq$ sub-bands specifies a pair of (possibly identical) datastream table bands to correlate. The global frequency table index of that product is by default identical to some index that one band out of the datastream table band pair ultimately refers to. Under DiFX 2.7 that target frequency can be specified explicitly and is interpreted as a spectral placement directive – which target frequency the sub-band shall contribute data to. One or more sub-bands can refer to the same target frequency index.

Key	Units/allowed values	Comments
BASELINE ENTRIES	integer ≥ 1	number of entries in table, $nBaseline$
D/STREAM A INDEX b	integer ≥ 0	datastream table index of first antenna
D/STREAM B INDEX b	integer ≥ 0	datastream table index of second antenna
NUM FREQS b	integer ≥ 1	number of frequencies on this baseline, $nFreq_b$
(TARGET FREQ b	integer ≥ 0	index to frequency table)
POL PRODUCTS b/f	integer ≥ 1	number of polarization products, $nProd_b$
D/STREAM A BAND p	integer ≥ 0	index to frequency array in datastream table
D/STREAM B BAND p	integer ≥ 0	same as above, but for antenna B, not A

7.26.8 Data Table

In the following table, d is the datastream index, ranging from 0 to $nDatastream - 1$ and f is the file index ranging from 0 to $nFile_d$.

Key	Units/allowed values	Comments
D/STREAM d FILES	integer ≥ 1	number of files $nFile_d$ associated with datastream d
FILE d/f	string	name of file or module associated with datastream d

For datastreams reading off Mark5 modules, $nFile$ will always be 1 and the filename is the *VSN* of the module being read.

7.27 .joblist

A single `.joblist` file is written by `vex2difx` (§6.102) as it produces the DiFX `.input` (and other) files for a given correlator pass. This file contains the list of jobs to run and some versioning information that allows improved accountability of the software versions being used. This file is used by `difxqueue` and `makefits` to ensure that a complete set of jobs is accounted for. The file is composed of two parts: a header line and one line for each job. The header line consists of a series of *key=value* pairs. Each *key* and *value* must have no whitespace and no whitespace should separate these words from their connecting = sign. While any number of key-value pairs may be specified, the following ones (which are case sensitive) are expected to be present:

1. **exper** : the name of the experiment, including the segment code
2. **v2d** : the `vex2difx` input file used to produce the jobs of this pass
3. **pass** : the name of the correlator pass
4. **mjd** : the modified Julian day when `vex2difx` created this file
5. **DiFX** : the version name for the DiFX deployment (the value of `$DIFX_VERSION` when `vex2difx` was run)

6. *vex2difx* : the version of *vex2difx* that was run

Each additional line contains information for one job in the pass. The columns are:

1. *jobName* : the name/prefix of the job
2. *mjdStart* : the observe start time of the job
3. *mjdStop* : the observe stop time of the job
4. *nAnt* : the number of antennas in the job
5. *maxPulsarBin* : the maximum number of pulsar bins to come from any scan in this job (usually zero)
6. *nPhaseCenters* : the maximum number of phase centers to come from any scan in this job (usually one)
7. *tOps* : estimated number of trillion floating point operations required to run the job
8. *outSize* : estimated FITS file output size (MB)

Usually the comment character *#* followed by a list of station codes is appended to the end of each line.

7.28 .jobmatrix

As of version 2.6 of *difx2fits* a file with extension *.jobmatrix* will be written for each *.FITS* file created. This file is meant as a summary for human use and as such does not have a format that should be considered fixed. The file contains a 2-dimensional map (antenna number vs. time) of which jobs contributed to the *.FITS* file.

7.29 .lag

Program *zerocorr* (§6.108 produces lag output in a format documented here. There is one line of output per lag. Each line has 7 columns as per the following table:

Line	Contents
1	Channel (spectral point) number (counts from zero)
2	Time lag (sec)
3	Real value of the lag function
4	Imaginary value of the lag function
5	Amplitude
6	Phase
7	Window function (weight at this lag)

7.30 .log

When generation of sniffer output files is not disabled, each *.FITS* file written by *difx2fits* will be accompanied by a corresponding *.log* file. This file contains a summary of the contents of that *.FITS* file. It is analogous to the *logfile.lis* file produced by the old *FITSniffer* program. This file is free-form ASCII that is intended for viewing by human eyes, and is should not be used as input to any software as the format is not guaranteed to remain constant.

7.31 .machines

The `.machines` file is used by `mpirun` to determine which machines will run `mpifxcorr`. This is a text file containing a list of computers, one to a line possibly with additional options listed, on which to spawn the software correlator process. As a general rule the MPI rank, a unique number for each process that starts at 0, are allocated in the order that the computer names are listed. This general rule can break down in cases where the same computer name is listed more than once; the behavior in this case depends on the MPI implementation being used. MPI rank 0 will always be the manager process. Ranks 1 through `nDatastream` will each be a datastream process. Additional processes will be computing (core) processes. If more processes are specified for `mpirun` with the `-np` option than there are lines in this file, the file will be read again from the top, so the processes will be assigned in a cyclic fashion (again, this depends somewhat on the MPI implementation and the other parameters passed to `mpirun`; for DiFX with OpenMPI, this assumes `--bynode` is used). If the program `startdifx` is used to start the correlation process, the number of processes to start is determined by the number of lines in this file. If wrapping to the top of this file is desired, dummy comment lines (beginning with `#`) can be put at the end of the `.machines` file to artificially raise the number of processes to spawn. Within DiFX, this file is typically produced by `genmachines`. Keep in mind that this file is directly read by the MPI execution program `mpirun` and the format of the file may differ depending on the MPI implementation that you are using. With OpenMPI appending `slots=1 max-slots=1` to the end of each line ensures that a single instance of `mpifxcorr` is run on that machine. If both a datastream process and a core process are to be run on the same computer, then using options `slots=1 max-slots=2` might be appropriate.

7.32 .mark4list

A `.mark4list` file is written by `makemark4` and contains the entire list of `gzip` compressed file sets (ending in `.mark4.tar.gz`) for the correlator pass. Due to the different constraints of the correlation process and the Mark4 format, the number of resultant compressed file sets may be greater or less than the number of jobs. This file type is used by `difxarch` to ensure that all of the correlated output ends up in the archive. The file is composed of two parts: a header line and one line for each compressed file set. The header line consists of a series of *key=value* pairs. Each *key* and *value* must have no whitespace and no whitespace should separate these words from their connecting `=` sign. While any number of key-value pairs may be specified, the following ones (which are case sensitive) are expected to be present:

1. `exper` : the name of the experiment, including the segment code
2. `pass` : the name of the correlator pass
3. `jobs` : the name of the `.joblist` file used by `makemark4`
4. `mjd` : the modified Julian day when `makemark4` created this file
5. `DiFX` : the version name for the DiFX deployment (the value of `$DIFX_VERSION` when `vex2difx` was run)
6. `difx2mark4` : the version of `difx2mark4` that was run

Each additional line contains information for one compressed file set of the correlation pass. These lines contain three fields:

1. `archiveName` : the name of the file that will get injected into the archive (see §??)
2. `fileSize` : the size of the file in MB
3. `origName` : the name of the file as produced by `difx2mark4` (via `makemark4`)

7.33 .oms

A `.oms` file is written by `sched` and contains machine (and human) readable information that is useful in setting up correlator jobs. In the case of the VLBA DiFX correlator, program `oms2v2d` (§6.72) uses this file to prepare a template `.v2d` file (§7.42) that contains some information not available in the `vex` file, such as intended integration time and number of channels.

7.34 .params

A file with extension `.params` is written by `vex2difx` (§6.102) when it is provided with the `-o` option. This file is a duplicate of the `.v2d` file that was supplied but with all unspecified parameters listed with the defaults that they assumed. The format is exactly the same as the `.v2d` files; see §7.42 for documentation of the format. The `.params` file can be used as a legal `.v2d` file if necessary.

7.35 pcal

A file called `pcal` is created when program `vlog` operates on the `cal.vlba` file. This file contains three measurements: the cable length calibration, pulse calibration, and state counts. This file contains two kinds of lines. Comment lines begin with an octothorpe (`#`) and contain no vital information. Data lines always begin with 9 fields describing the content of that data line:

1. *antId* : Station name abbreviation, e.g., LA
2. *day* : Time centroid of measurement (MJD or day of year, including fractional portion)
3. *dur* : Duration of measurement (days)
4. *cableCal* : Cable calibration measurement (picoseconds)
5. *nPol* : Number of polarizations with measurements
6. *nBand* : Number of sub-bands with measurements
7. *nTone* : Number of pulse cal tones detected per band per polarization, possibly zero
8. *nState* : Number of state count states measured per band per polarization, possibly zero
9. *nRecChan* : Number of record channels at time of measurement ($\leq nPol * nBand$)

Following these nine fields are two variable-length arrays of numbers. The first variable-length field is the pulse cal data field consisting of $nPol * nBand * nTone$ groups of four numbers. The first member of this group is the recorder channel number (zero-based) corresponding to the measurement. The second member of this group is the tone sky frequency (MHz). The third and fourth are respectively the real and imaginary parts of the tone measured at the given sky frequency. The order in which the groups are presented (in ‘C’ language array syntax, as used throughout this document) is $[nPol][nBand][nTone]$. Note that if there are fewer than $nPol * nBand$ record channels, the record channel will be -1 for some groups. The second variable-length field is the state count data. For each band of each polarization, $nState + 1$ values are listed. The first number is the record channel number or -1 if that polarization/band combination was not observed or monitored. The remainder contain state counts. $nState$ can be either 0 or 2^{nBit} , where $nBit$ is the number of quantization bits. The order in which these groups are listed is $[nPol][nBand]$.

7.36 .polyco

A `polyco` file contains a single polynomial for pulse phase that is valid for a fraction (up to 100%) of a job file. An additional numeric suffix is appended to the filename specifying the polynomial index for a particular `.pulsar` file that shares the same base name. The format of the file is the same as a Tempo pulsar file [8].

7.37 .shelf

The vex file format (see §7.44 and references within) does not have a formal slot to record the shelf location of media, so `db2vex` stashes the shelf location in a separate file. This information is critical for the correlator operators to know where to find modules for a project and for analysts preparing correlator jobs to know if media have arrived. The `.shelf` file is used by `vex2difx` when writing `.calc` files. It can also be used as input to `getshelf`. The file format is very simple. One row is used for each module that was used in the observation. Typically rows are sorted in the same order as antennas in the `.input` file. The comment character is `#` – any text following this character on a line is ignored. Each line contains 3 white-space separated columns:

1. *antId* : The typically 2 letter station abbreviation
2. *vsu* : The volume serial number of the media (e.g., NRAO-123)
3. *shelf* : The shelf location, which can be any string without whitespace (e.g., BD89), or `none` if the media is not at the correlator

7.38 .threads

The `.threads` file tells `mpifxcorr` how many threads to start on each processing node. Within DiFX, this file is typically produced by `genmachines`. The `.threads` file has a very simple format. The first line starts with `NUMBER OF CORES:`. Starting at column 21 is an integer that should be equal to the number of processing nodes (*nCore*) specified in the corresponding `.machines` file. Each line thereafter should contain a single integer starting at column 1. There should be *nCore* such lines.

7.39 tsys

A file called `tsys` is created when program `vlog` operates on the `cal.vlba` file. This text file contains measurements of the system temperature and name of receiver for each baseband channel. This file contains two kinds of lines. Comment lines begin with an octothorpe (`#`) and contain no vital information. Data lines always contain 4 fixed-size fields:

1. *antId* : Station name abbreviation, e.g., LA
2. *day* : Time centroid of measurement (day of year or mjd, including fractional portion)
3. *dur* : Duration of measurement (days) or zero if not known
4. *nRecChan* : Number of baseband channels recorded

Following these 4 fields are *nRecChan* pairs of values, one for each baseband channel. The first element of each pair is the system temperature (in K) and the second is the receiver name (e.g., 4cm, or 7mm).

This format should not be confused with switched power files produced by `mpifxcorr` (see §7.12.3).

7.40 weather

A file called `weather` is created when program `vlog` operates on the `cal.vlba` file. This file contains tabulated values of various meteorological measurements. This file contains two kinds of lines. Comment lines begin with an octothorpe (`#`) and contain no vital information. Data lines always contain 9 fixed-size fields:

1. *antId* : Station name abbreviation, e.g., LA
2. *day* : Time of measurement (MJD or day of year, including fractional portion)
3. *T* : Ambient temperature (Centigrade)

4. *P* : Pressure (mbar)
5. *dewPoint* : Dew point (Centigrade)
6. *windSpeed* : Wind speed (m/s)
7. *windDir* : Wind direction (degrees E of N)
8. *precip* : Accumulated rain since UT midnight (cm)
9. *windGust* : Maximum wind gust over collection period (m/s)

7.41 .wts

When generation of sniffer output files is not disabled, each `.FITS` file written by `difx2fits` will be accompanied by a corresponding `.wts` file. This file contains statistics of the data weights, typically dominated by the completeness of records as determined by the data transport system, over a typically 30 second long period.

The first line is simply the observe code, e.g., `MT831` .

Each additional line in the file is a complete record for a given antenna for a given interval, containing information for each baseband channel separately. The format of these lines is as follows:

Key	Units/allowed values	Comments
<i>MJD</i>	integer ≥ 1	MJD day number corresponding to line
<i>hour</i>	≥ 0.0 , < 24.0	hour within day
<i>antenna number</i>	≥ 1	antenna table index
<i>antenna name</i>	string	
<i>n_{BBC}</i>	≥ 1	Number of baseband channels
<i>mean weight</i>	≥ 0.0	This column repeated <i>n_{BBC}</i> times
<i>min weight</i>	≥ 0.0	This column repeated <i>n_{BBC}</i> times
<i>max weight</i>	≥ 0.0	This column repeated <i>n_{BBC}</i> times

This file can be used directly with plotting program `plotwt` or used more automatically with `difxsniff`.

7.42 .v2d

The `.v2d` file is used to specify correlation options to `vex2difx` and adjust the way in which it forms DiFX input files based on the `.vex` file. The `.v2d` file consists of a number of global parameters that affect the way that jobs are created and several sections that can customize correlation on a per-source, per mode, or per scan basis. All parameters (those that are global and those that reside inside sections) are specified by a parameter name, the equal sign, and one value, or a comma-separated list of values, that cannot contain whitespace. Whitespace is not required except to keep parameter names, values, and section names separate. All parameter names and values are case sensitive except for source names and antenna names. The `#` is a comment character; any text after this on a line is ignored.

Most parameters are one of the following types:

- **bool** : A boolean value that can be True or False. Any value starting with 0, f, F, or - will be considered False and otherwise True.
- **float** : A floating point number. Can be of the forms: 1.23, 1.2e-4, -12.6, 4
- **int** : An integer.
- **string** : Any sequence of printable(non-whitespace) characters. Certain fields require strings of a maximum length or certain form.

- **date** : A date field; see below.
- **array** : Array can be of any of the four above types and are indicated by enclosing brackets, e.g., [int]. The empty list is indicated with [] which is usually implied to be all-inclusive.

All times used in `vex2difx` are in Universal Time and are internally represented as a double precision value. The integer part of this value is the date corresponding to 0^h UT. The fractional part, when multiplied by 86400, gives the number of seconds since 0^h UT. Note that this format does not allow one to specify the actual leap second if one occurs on that day. Several date formats are supported:

- **Modified Julian Day** : A decimal MJD possibly including fractional day. E.g.: 54345.341944
- **Vex time format** : A string of the form: 2009y245d08h12m24s
- **VLBA-like format** : A string of the form: 2009SEP02-08:12:24
- **ISO 8601 format** : A string of the form: 2009-09-02T08:12:24

Global parameters can be specified one or many per line such as:

`maxGap = 2000 # seconds`

or

`mjdStart = 52342.522 mjdStop=52342.532`

The following parameter names are recognized:

Name	Type	Units	Defaults	Comments
vex	string			filename of the vex file to process; this is required
mjdStart	date		obs. start	discard any scans or partial scans before this time
mjdStop	date		obs. stop	discard any scans or partial scans after this time
break	date			list of MJD date/times where jobs are forced to be broken
minSubarray	int		2	don't make jobs for subarrays with fewer antennas than this
maxGap	float	sec	180	split an observation into multiple jobs if there are correlation gaps longer than this number
tweakIntTime	bool		False	adjust (up to 40%) int. time to ensure int. blocks per send
singleScan	bool		False	if True, split each scan into its own job
singleSetup	bool		True	if True, allow only one setup per job; True is required for FITS-IDI conversion
maxLength	float	sec	7200	don't allow individual jobs longer than this amount of time
minLength	float	sec	2	don't allow individual jobs shorter than this amount of time
maxSize	float	MB	2000	max FITS-IDI file size to allow
dataBufferFactor	int		32	the mpifxcorr DATABUFFERFACTOR parameter
nDataSegments	int		8	the mpifxcorr NUMDATASEGMENTS parameter
jobSeries	string			the base filename of .input and .calc files to be created; defaults to the base name of the .v2d file
startSeries	int		20	the default starting number for jobs created
sendLength	float	sec	0.262144	roughly the amount of data to send at a time from datastream processes to core processes
antennas	[string]		[] = all	a comma separated list of antennas to include in correlation
baselines	[string]		[] = all	a comma separated list of baselines to correlate; see below
padScans	bool		True	insert non-correlation scans in recording gaps to prevent mpifxcorr from complaining
invalidMask	int		0xFFFF	this bit-field selects which flag conditions are considered when writing flag file: 1=Recording, 2=On source, 4=Job time range, 8=Antenna in job
visBufferLength	int		32	number of visibility buffers to allocate in mpifxcorr
overSamp	int			force all basebands to use the given oversample factor
mode	string		normal	mode of operation; see below
threadsFile	string			overrides the name of the threads file to use
nCore	int			with nThread and machines, cause a .threads file to be made
nThreads	int			number of threads per core in .threads file
machines	[string]			comma separated list of machines to use as processors
				first is head node, then datastreams, then cores
maxReadSize	int	bytes	25000000	maximum number of bytes to read at a time
minReadSize	int	bytes	10000000	minimum number of bytes to read at a time

The *baselines* parameter supports the wildcard character * an individual antenna name, or lists of antenna names separated by + on each side of a hyphen (-). Multiple baseline designators can be listed. Examples:

- A1-A2 : Only correlate one baseline
- A1-A2, A3-A4 : Correlate 2 baselines
- *-* : Correlate all baselines
- A1-* or *-A1 : Correlate all baselines to antenna A1
- A1+A2-* : Correlate all baselines to antenna A1 or A2
- A1+A2-A3+A4+A5 : Correlate 6 baselines

A source section can be used to change the properties of an individual source, such as its position or name. In the future this is where multiple correlation centers for a given source will be specified. A source section is enclosed in a pair of curly braces after the keyword SOURCE followed by the name of a source, for example

```
SOURCE 3C273
{
    source parameters go here
}
```

or equivalently

```
SOURCE 3c273 { source parameters go here }
```

Name	Type	Units	Defaults	Comments
ra		J2000		right ascension, e.g., 12h34m12.6s or 12:34:12.6
dec		J2000		declination, e.g., 34d12'23.1" or 34:12:23.1
name	string			new name for source
calCode	char		' '	calibration code, typically A, B, C for calibrators, G for a gated pulsar, or blank for normal target
naiffFile	string			name of leap seconds file (e.g., <code>naif0010.tls</code> for ephemeris)
ephemObject	string			name or number of object in ephemeris file
ephemFile	string			path of ephemeris file (either <code>.bsp</code> or <code>.tle</code> format)
doPointingCentre	bool		true	Whether the pointing centre should be correlated (only ever turned off for multi-phase center)
addPhaseCentre	string			contains info on a source to add; see below

To add additional phase centers, add one or more “addPhaseCentre” parameters to the source setup. In the parameter, the RA and dec must be provided. A name and/or calibrator code can be added as well. For example: `addPhaseCentre=name@1010-1212/RA@10:10:21.1/Dec@-12:12:00.34` .

An antenna section allows properties of an individual antenna, such as position, name, or clock/LO offsets, to be adjusted. Note that the “late” convention is used in `clockOffset` and `clockRate`, unlike the “early” convention used in the `.vex` file itself (see §5.1).

Name	Type	Units	Defaults	Comments
name	string			new name to assign to this antenna
polSwap	bool		False	swap the polarizations (i.e., $L \leftrightarrow R$) for this antenna
clockOffset	float	us	vex value	overrides the clock offset value from the vex file
clockRate	float	us/s	vex value	overrides the clock offset rate value from the vex file
clockEpoch	date		vex value	overrides the epoch of the clock rate value; must be present
deltaClock	float	us	0.0	adds to the clock offset (either the vex value or the clockOffset above)
deltaClockRate	float	us/s	0.0	adds to the clock rate (either the vex value or the clockRate above)
X	float	m	vex value	change the X coordinate of the antenna location
Y	float	m	vex value	change the Y coordinate of the antenna location
Z	float	m	vex value	change the Z coordinate of the antenna location
format	string			force format to be one of VLBA, MKIV, Mark5B, S2, or one of the VDIF types
file	[string]		(none)	a comma separated list of data files to correlate
filelist	string			a filename listing files for the DATA TABLE (see §7.18)
networkPort	int			The eVLBI network port to use for TCP/UDP. A non-number indicates raw mode attached to an ethernet interface. Both force NETWORK type in <code>.input</code> file.
windowSize	int			TCP window size in kilobytes for eVLBI. Set to < 0 in bytes for UDP
UDP_MTU	int			Same as setting windowSize to negative of value.
vsu	string			For raw mode, the number of bytes to strip from ethernet frame.
zoom	string			override the Mark5 Module to be used
addZoomFreq	string			uses the global zoom configuration with matching name for this antenna; <code>zoom=Zoom1</code> will match ZOOM block called <code>Zoom1</code>
freqClockOffs	[float]	microsec		adds a zoom band with specified freq/bw as shown: <code>freq@1810.0/bw@4.0[/specAvg@4][/noparent@false]</code>
loOffsets	[float]	Hz		adds clock offsets to each recorded frequency using the format: <code>freqClockOffs=f1,f2,f3,f4</code> ; must be same length as number of recorded freqs, first value must be zero
tcalFreq	int	Hz	0	adds LO offsets to each recorded frequency using the format: <code>loOffsets=f1,f2,f3,f4</code> ; must be same length as number of recorded freqs.
phaseCalInt	int	MHz	1	enables switched power detection at specified frequency
toneGuard	float	MHz	0.125 of bw	zero turns off phase cal extraction, positive value is the interval between tones to be extracted
toneSelection	string		smart	when using toneSelection <i>smart</i> or <i>most</i> don't select tones within this range of band edge, if possible
sampling	string		REAL	tone selection algorithm; see below
fake	bool		False	set to COMPLEX for complex sampled data
				enable a fake data source

Possible values of “tone Selection” are:

smart	write the 2 most extreme tones at least toneGuard from band edge (default)
vex	write the tones listed in the vex file to FITS
none	don't write any tones to FITS
all	write all extracted tones to FITS
ends	write the 2 most extreme tones to FITS
most	write all tones not closer than toneGuard to band edge

Setup sections are enclosed in braces after the word **SETUP** and a name given to this setup section. The setup name is referenced by a **RULE** section (see below). A setup with the special name **default** will be applied to any scans not otherwise assigned to setups by rule sections. If no setup sections are defined, a setup called **default**, with all default parameters, will be implicitly created and applied to all scans. The order of setup sections is immaterial.

Name	Type	Units	Defaults	Comments
tInt	float	sec	2	integration time
FFTSpecRes	float	MHz	0.125	frequency resolution of FFT
specRes	float	MHz	0.5	output freq res (must be mult. of FFTSpecRes)
nChan	int		16	number of channels per spectral window; must be $5^m \cdot 2^n$
specAvg	int		1	how many channels to average together after correlation
fringeRotOrder	int		1	the fringe rotation order: 0=post-F, 1=linear, 2=quadratic
strideLength	int		16	number of channels to “stride” for fringe rotation, etc.
xmacLength	int		128	number of channels to “stride” for cross-multiply accumulations
numBufferedFFTs	int		1	number of FFTs to do in a row for each datastream, before XMAC
doPolar	bool		True	correlate cross hands when possible
postFFFringe	bool		False	do fringe rotation after FFT?
binConfig	string		none	if specified, apply this pulsar bin config file to this setup
freqId	[int]		[] = all	frequency bands to correlate
outputBandwidth	float 'auto'	MHz	n/a	Target bandwidth when auto determining outputbands, or,
addOutputBand	[string]		n/a	Add an outputband with explicit placement and bandwidth freq@1810.0/bw@4.0

Note that either “FFTSpecRes” and “specRes” can be used, or “nChan” and “specAvg” can be used, but the two sets cannot be mixed.

Zoom channels can be configured in a special section and referenced from ANTENNA sections to minimize complexity of the .v2d file. Each ZOOM section has a name and one or more “addZoomFreq” parameters, with the same format as they would have in the ANTENNA block.

Output bands can be defined in the SETUP block in DiFX 2.7 (OUTPUTBAND block in DiFX 2.8) via “outputBandwidth” or alternatively “addOutputBand”. Permitted “outputBandwidth” values are ‘auto’ to auto-determine the bandwidth, or, an explicit bandwidth as a decimal value in MHz. Band placement is automatic. For explicit placement and bandwidth instead use “addOutputBand”.

Earth Orientation Parameter (EOP) data can be provided via one or more EOP sections. EOP data should be provided either in the .v2d file or in the vex file, but not both. Normally the vex file would be used to set EOP values, but there may be cases (eVLBI?) that want to use the vex file from **sched** without any modification. Like ANTENNA and SOURCE sections, each EOP section has a name. The name must be in a form that can be converted directly to a date (see above for legal date formats). Conventional use suggests that these dates should correspond to 0 hours UT; deviation from this practice is at the users risk. There are four parameters that should all be set within an EOP section:

Name	Type	Units	Defaults	Comments
tai_utc	float	sec		TAI minus UTC; the leap-second count
ut1_utc	float	sec		UT1 minus UTC; Earth rotation phase
xPole	float	arcsec		X component of spin axis offset
yPole	float	arcsec		Y component of spin axis offset

A rule section is used to assign a setup to a particular source name, calibration code (currently not supported), scan name, or vex mode. The order of rule sections *does* matter as the order determines the priority of the rules. The first rule that matches a scan is applied to that scan. The correlator setup used for scans that match a rule is determined by the parameter called “setup”. A special setup name **SKIP** causes

matching scans not to be correlated. Any parameters not specified are interpreted as fully inclusive. Note that multiple rule sections can reference the same setup section. Multiple values may be applied to any of the parameters except for “setup”. This is accomplished by comma separation of the values in a single assignment or with repeated assignments. Thus

```

RULE rule1
{
    source = 3C84,3C273
    setup = BrightSourceSetup
}

```

is equivalent to

```

RULE rule2
{
    source = 3C84 3C273
    setup = BrightSourceSetup
}

```

is equivalent to

```

RULE rule3
{
    source = 3C84
    source = 3C273
    setup = BrightSourceSetup
}

```

The names given to rules (e.g., rule1, rule2 and rule3 above) are not used anywhere (yet) but are required to be unique.

Name	Type	Units	Comments
scan	[string]		one or more scan name, as specified in the vex file, to select with this rule
source	[string]		one or more source name, as specified in the vex file, to select with this rule
calCode	[char]		one or more calibration code to select with this rule
mode	[string]		one or more modes as defined in the vex file to select with this rule
setup	string		The name of the SETUP section to use, or SKIP if this rule describes scans not to correlate

Note that source names and calibration codes reassigned by source sections are not used. Only the names and calibration codes in the vex file are compared.

There are currently two modes of operation supported by `vex2difx`. The mode used in the vast majority of situations is called **normal** and is the default if none is specified. Currently one alternative mode, **profile**, is supported. This mode is useful for generating pulse profiles that would be useful for pulsar gating, scrunching, and binning. The difference compared to normal mode is that the standard autocorrelations are turned off and instead are computed as if they are cross correlations. This allows multiple pulsar bins to be stored. No formal cross correlations are performed. To be useful, one must create and specify a `.binconfig` file and select only the pulsar(s) from the experiment.

See <http://www.atnf.csiro.au/vlbi/dokuwiki/doku.php/difx/vex2difx> for more complete information and examples.

7.43 .xcb

When generation of sniffer output files is not disabled, each .FITS file written by `difx2fits` will be accompanied by a corresponding .xcb file. This file contains cross-correlation spectra for each antenna for each baseline. In order to minimize the output data size, spectra for the same source will only be repeated once per 15 minutes. The file contains many concatenated records. Each record has the spectra for all baseband channels for a particular baseline and has the following format which is very similar to that of the .acb files. Note that no spaces are allowed within any field. Values in **typewriter** font without comments are explicit strings that are required.

Line(s)	Value	Units	Comments
1	timerange: <i>MJD</i> <i>start time</i> <i>stop time</i> obscode: <i>observe code</i> chans: <i>n_{chan}</i> x <i>n_{BBC}</i>	 integer ≥ 1 string string string ≥ 1 ≥ 1	 MJD day number corresponding to line e.g., 13h34m22.6s e.g., 13h34m52.0s e.g., MT831 number of channels per baseband channel number of baseband channels
2	source: <i>source name</i> bandw: <i>bandwidth</i> MHz	 string MHz	 e.g., 0316+413 baseband channel bandwidth
3 to 2+n _{BBC}	bandfreq: <i>frequency</i> GHz GHz polar: <i>polarization</i> side: <i>sideband</i> bbchan: <i>bbc</i>	 GHz 2 chars U or L 0	 band edge (SSLO) frequency of baseband channel e.g., RR or LL for upper or lower sideband Currently not used but needed for conformity
3+n _{BBC} to 2+n _{BBC} (n _{chan} + 1)	<i>ant1 number</i> <i>ant2 number</i> <i>ant1 name</i> <i>ant2 name</i> <i>channel number</i> <i>amplitude</i> <i>phase</i>	 ≥ 1 ≥ 1 string string ≥ 1 ≥ 0.0 degrees	 number of first antenna = chan + (bbc - 1) · n _{chan} for chan, bbc ≥ 1

The above are repeated for each cross correlation spectrum record. This file can be plotted directly with `plotbp` or handled more automatically with `difxsniff`.

7.44 .vex, .skd, .vex.obs, & .skd.obs

The vex (Vlbi EXperiment) file [11] format is a standard observation description format used globally for scheduling observations and for driving the correlation thereof. The original vex file for an experiment is typically created by `sched` or `sked`. In the former case (the case used by most astronomical VLBI), the vex file has the unfortunate file extension .skd; in the later, the file extension is usually the less confusing .vex. These two vex formatted files contain only observation-scheduling based information. A small amount of information based on the actualities of the observation are added by `db2vex`, producing a new vex file

with an additional file extension `.obs` . Please see vex documentation external to this manual for more information.

7.45 `.vis`

Program `zerocorr` (§6.108 produces visibility output in a format documented here. There is one line of output per generated spectral point. Each line has 8 columns as per the following table:

Line	Contents
1	Channel (spectral point) number (counts from zero)
2	Frequency relative to first spectral channel (Hz)
3	Real value of the visibility
4	Imaginary value of the visibility
5	Amplitude
6	Phase
7	Autocorrelation of the first datastream (real only)
8	Autocorrelation of the second datastream (real only)

7.46 `.zerocorr`

Program `zerocorr` (§6.108 is a simple cross correlator. It is limited to correlating one visibility spectrum from one baseband channel and can only make use of a constant offset delay model. The section documents the file that drives this program. This file consists of 17 lines of text. The first 7 lines describe properties of the data from the first antenna and the following (7) lines describe properties of the second antenna as follows:

Line	Contents
1 (8)	Input baseband data file name
2 (9)	Data format (e.g., Mark5B-2048-16-2)
3 (10)	Input sub-band to process (0-based index)
4 (11)	Offset into the file (bytes)
5 (12)	Size of FFT to perform over input bandwidth ($2 \times n_{\text{chan}}$)
6 (13)	First channel (spectral point) to correlate
7 (14)	Number of channels to correlate (negative indicates LSB)

The last three lines dictate the output data files and the number of FFTs to process:

Line	Contents
15	Name of output visibility (<code>.vis</code> ; §7.45) file
16	Name of output <code>.lag</code> (§7.29) file
17	Number of FFTs to process (if -1, run on entire input files)

8 XML message types

The `difxmessage` library (§??) implements a system for sending and receiving messages using XML format. This section documents the “difxMessage” XML document type that is used for interprocess communication during correlation within DiFX. These messages are sent via UDP multicast and are thus restricted to fit within one standard-sized Ethernet packet (~ 1500 bytes). Various logging and monitoring programs (`mk5mon`, `cpumon`, and `errormon`, all eventually to be replaced by a single interactive operator interface) can accept these messages and perform actions based on their content. Several different message types are derived from the following XML base type:

```

<?xml version="1.0" encoding="UTF-8"?>
<difxMessage>
  <header>
    <from>from</from>
    <to>to</to>
    <mpiProcessId>mpiId</mpiProcessId>
    <identifier>idString</identifier>
    <type>messageType</type>
  </header>
  <body>
    <seqNumber>seqNum</seqNumber>
    body
  </body>
</difxMessage>

```

The italicized fields are as follows:

from the hostname of the sender.

to the intended recipient of the XML document. Note that this field is typically not included for report-only messages as it's intended purpose is for directing commands to particular recipients. Also note that multiple *to* fields can be present in a message. Three “shortcuts” are currently allowed: **all** causes all receiving programs (such as **mk5daemon**) on all software correlator cluster members to respond; **mark5** causes all Mark5 units to respond; and **swc** causes all non-Mark5 units to respond.

mpiId the MPI process id of the sender. If there are D (typically 10) datastream processes (i.e., Mark5 units playing back), then *mpiId* takes on the following numbers:

value	mpifxcorr process type
< 0	a process not associated with mpifxcorr
0	the manager process of mpifxcorr
1 to D	one of the datastream processes
$\geq D + 1$	one of the core (computing) processes

idString an additional string identifying the source of the message. For messages sent from **mpifxcorr**, this will be the job id, for example **job3322.000**. Other programs will typically set this field to the name of the program sending the message.

messageType the type of message being sent:

value	description of message type
<code>DifxAlertMessage</code>	an error message.
<code>DifxCommand</code>	a command message.
<code>DifxDatastreamMessage</code>	<i>Not yet implemented</i>
<code>DifxDiagnosticMessage</code>	used by <code>mpifxcorr</code> to pass out diagnostic-type info such as buffer states
<code>DifxFileTransfer</code>	used by the new (USNO) GUI to cause a file to be sent to or from a specified host
<code>DifxFileOperation</code>	used by the new (USNO) GUI to cause some operation to a file (such as <code>mkdir</code> , <code>mv</code> , <code>rm</code>)
<code>DifxGetDirectory</code>	used by the new (USNO) GUI to request a Mark5 <code>.dir</code> file
<code>DifxInfoMessage</code>	<i>not used?</i>
<code>DifxLoadMessage</code>	CPU and memory usage (usually sent by <code>mk5daemon</code>).
<code>DifxMachinesDefinition</code>	used by the new (USNO) GUI to remotely write a <code>.machines</code> file
<code>DifxParameter</code>	specify new parameter value to an <code>mpifxcorr</code> process.
<code>DifxSmartMessage</code>	contains smart data for one drive of a module; usually sent by <code>mk5daemon</code>
<code>DifxStart</code>	tell head node to start a difx job.
<code>DifxStop</code>	tell <code>mk5daemon</code> to stop a particular instance of <code>mpifxcorr</code> .
<code>DifxStatusMessage</code>	status of the <code>mpifxcorr</code> program.
<code>DifxTransientMessage</code>	used by the VFASTR project to indicate an event of interest
<code>DifxVex2DifxRun</code>	used by the new (USNO) GUI to remotely run <code>vex2difx</code>
<code>Mark5DriveStatsMessage</code>	Mark5 module conditioning statistics for one disc.
<code>Mark5StatusMessage</code>	status of the mark5 unit and modules.
<code>Mark5VersionMessage</code>	versions, board types and serial numbers of a Streamstor card.

Many of these message types are described in sections that follow. For those without documentation, the source file `difxmessage.h` in the `difxmessage` package contains the full set of parameters.

seqNum the sequence number (starting at 0) of messages coming from the particular program running on the particular host. The number advances by 1 for each sent message and can be used to detect lost packets.

body message contents that are specific to the particular *messageType*. See sections that follow.

A “C” language library for generating, multicasting, receiving, and parsing XML documents of this type is used within some of the programs, including `mpifxcorr` (the core of the DiFX [2] software correlator) and `mk5daemon` (a program that runs on each Mark5 unit that is responsible for multicast communication when `mpifxcorr` is not running), that transact these XML documents. The default multicast group to be used is 224.2.2.1 and the default port is 50200, though these can be overridden with environment variables `DIFX_MESSAGE_GROUP` and `DIFX_MESSAGE_PORT` respectively.

8.1 DifxAlertMessage

This section describes messages with *messageType* = `DifxAlertMessage`. These messages come from `mpifxcorr` or the head node agent and contain an error message string and severity code that should be displayed to the operator and logged.

The *body* of the message contains:

```
<difxAlert>
  <alertMessage>message</alertMessage>
  <severity>severity</severity>
</difxAlert>
```

The italicized fields are as follows:

message a string containing the error message.

severity an integer indicating the severity. The severity scale is based on that from the EVLA and has values with the following meanings:

value	name	meaning
0	FATAL	processing has failed; a restart is needed
1	SEVERE	data from one or more station is affected badly
2	ERROR	data from one or more station may be affected; e.g., low weights
3	WARNING	minor error of no consequence to ongoing processing
4	INFO	informational only
5	VERBOSE	overly verbose information
6	DEBUG	debugging information

8.2 DifxCommand

This section describes messages with *messageType* = **DifxCommand**. These messages require the *to* field to be set and cause the intended recipient to take an action.

The *body* of the message contains:

```
<difxCommand>
  <command>command</command>
</difxCommand>
```

The italicized field is as follows:

command the command to execute. Commands are not case sensitive and should be among the following:

command	action
GetVSN	cause the mark5 unit to multicast loaded VSNs if possible.
GetLoad	request CPU and memory usage to be reported.
ResetMark5	cause SSReset and ssopen to be run to reset Streamstor.
StartMark5A	start the Mark5A program.
StopMark5A	stop the Mark5A program.
KillMpifxcorr	kill with signal 9 (sigkill) any process with name mpifxcorr .
Clear	reset the mk5daemon; useful sometimes if mpifxcorr crashes.
Reboot	causes machine to reboot.
Poweroff	causes machine to power down.
Copy <bank> <vsn> <scans>	causes scans to be copied to local disk.

8.3 DifxLoadMessage

This section describes messages with *messageType* = **DifxLoadMessage**. These messages contain CPU and memory utilization information and are voluntarily sent by various nodes of the cluster, to be received by the operator interface.

The *body* of this message type contains:

```
<difxLoad>
  <cpuLoad>cpuLoad</cpuLoad>
  <totalMemory>totalMemory</totalMemory>
  <usedMemory>usedMemory</usedMemory>
</difxLoad>
```

The italicized fields are as follows:

cpuLoad CPU utilization on the cluster node. It is a floating point value containing the average number of processes scheduled at one time.

totalMemory total memory on node, in kiB.

usedMemory used memory on node, in kiB.

8.4 DifxParameter

The structure of a DifxParameter message *body* is as follows:

```
<difxParameter>
  <targetMipId>id</targetMipId>
  <name>name</name>
  <index1>index1</index1>
  .
  .
  .
  <indexN>indexN</indexN>
  <value>value</value>
</difxParameter>
```

Such a message is intended to allow a parameter, possibly qualified with array indices, to be set to a particular value. A possible use, for example, is to change a station clock value within a running DiFX instantiation or to enable generation of fast dump spectra for transient searching.

The italicized fields are as follows:

targetMipId the MPI process Id (rank) to target with this message. Values zero and greater target specific MPI processes, with zero always being the manager process. Other special values include:

value	meaning
-1	all MPI processes
-2	all core (processing) processes
-3	all datastream processes

name the name of the parameter to set. For array types, the following *index* values specify the element to set.

indexN an integer specifying the index of the particular array axis.

value a string containing the value.

8.5 DifxSmartMessage

The Self-Monitoring, Analysis and Reporting Technology (SMART) protocol is used by hard drives for health monitoring. The Mark5 units support this and make the SMART values available via the Streamstor API. *mk5daemon* accesses this information periodically and on module insertion so that operators can be made aware of obvious and potential module problems as early as possible. The DifxSmartMessage multicast message is used to convey such SMART information for individual drives in a module. It should be expected that when such messages are sent a separate message will come for each drive (typically 8) of a module.

The *body* of this message type contains:

```

<difxSmart>
  <mjd>mjd</mjd>
  <vsn>vsn</vsn>
  <slot>slot</slot>
  <smart id=smartId value=value />
  .
  .
  .
</difxSmart>

```

Multiple **smart** tags (8 to 16 are usual) will usually be present in each smart message.

The italicized fields are as follows:

mjd The time at which the SMART value was extracted.

vsn The module number.

slot The slot of the hard drive in the module (0 to 7).

smartId The identifier for the value being represented. This is usually a small (< 300) positive integer.

More information can be found at <http://en.wikipedia.org/wiki/S.M.A.R.T.>

value The value of the monitor point corresponding to the *smartId*.

8.6 DifxTransientMessage

This section describes messages with *messageType* = **DifxTransientMessage**. This message is related to a commensal transient search project. A message of this type should be sent as soon as possible after detection; it is likely that no provisions will be made for data copying that does not start before resources assigned to the job are released. When a possible transient is identified by a detecting program which looks at autocorrelations it sends this message to all Mark5 units. Once correlation is complete, the **mk5daemon** program on appropriate Mark5 units will take a few seconds to copy data from the time range(s) of interest.

The *body* of this message type contains:

```

<difxTransient>
  <jobId>jobId</jobId>
  <startMJD>startMJD</startMJD>
  <stopMJD>stopMJD</stopMJD>
  <priority>priority</priority>
  <destDir>destDir</destDir>
  <comment>comment</comment>
</difxTransient>

```

The italicized fields are as follows:

jobId The job identification string. This is required so that only relevant Mark5 units take action on the received message.

eventId A string containing the name of the event as declared by the transient detector.

startMJD The start time (MJD) of the segment of data to copy.

stopMJD The stop time (MJD) of the segment of data to copy. Note that the total amount of data to be copied should be low enough to have an inconsequential impact on correlation throughput.

priority A floating point value indicating the relative importance of capturing this event. In jobs where many triggers occur this field will be used to select the most important ones to save to disk. Higher numbers indicate higher priority.

destDir (optional) A final directory to store the baseband data. If not provided, a default will be assumed. Note that behavior is undefined if different destination directories are provided within a single job.

classification (optional) A string provided by the transient detector containing a tentative classification.

comment (optional) A comment that could be appended to a log file.

All data and a log file will be stored in a subdirectory of the data staging area named *jobId*. The log file, at a minimum, will contain the list of events sent by the transient detection program and a log of the copying process, indicating any errors that may have occurred. The subdirectory will have a temporary name starting with a period (.) until all data copy for the job in question is complete. The use of this message type is demonstrated in Fig. 2.

8.7 DifxStart

This document type causes the head node to spawn a correlator job. The document contents describe which resources to use and which *.input* file to use.

The *body* of the message contains:

```
<difxStart>
  <input>input_file</input>
  <force>forceOverwrite</force>
  <manager node="node" />
  <datastream nodes="nodes" />
  <process nodes="nodes" threads="count" />
  <env>envvar=value</env>
  <difxProgram>program</difxProgram>
  <difxVersion>version</difxVersion>
  <mpiWrapper>mpiWrapper</mpiWrapper>
  <mpiOptions>options</mpiOptions>
</difxStart>
```

In the above XML file, exactly one manager node must be supplied. There must be at least one datastream node (one per antenna being correlated). There must be at least one process node. Zero or more (up to a maximum of 8) environment variables may be set. The italicized fields are as follows:

input file complete path to the *.input* file for this correlation.

forceOverwrite cause any previous correlator output of this job to be deleted before starting the correlation. A value of 1 or **True** will enable overwrite.

value the value of the environment variable.

node the name of the node being assigned, e.g. *mark5fx02* or *swc000*.

nodes a list of node names. The list members should be space or comma separated.

count the maximum number of threads to schedule. If not specified, 1 will be assumed. This applies only to process nodes.

envvar an environment variable to set before running *mpifxcorr*.

program the name of the software correlator executable. This is optional and defaults to `mpifxcorr` if not set.

version the version (e.g., *DIFX-1.5.4*) of DiFX to run.

mpiWrapper the name of the program used to start the MPI processes. This field is optional and defaults to `mpirun` if none is provided.

options extra options to pass to `mpirun`. This is optional; sensible defaults are assumed if not explicitly set.

Note that multiple `<process />` tags can be specified, each with its own thread count. Each tag's thread count only affects those nodes specified in that tag. If *version* is provided, then a wrapper script called `runmpifxcorr.version` is expected to be in the default path which sets the environment for the version of DiF to actually run.

8.8 DifxStatusMessage

This section describes messages with *messageType* = `DifxStatusMessage`. This message type is only produced by `mpifxcorr` or the programs immediately responsible for starting and stopping it.

The *body* of the message contains:

```
<difxStatus>
  <state>state</state>
  <message>message</message>
  <visibilityMJD>visibilityMJD</visibilityMJD>
  <weight ant=antId wt=weight>
</difxStatus>
```

The italicized fields are as follows:

state the state of `mpifxcorr`, which must be one of the following:

state	meaning
Spawning	the <code>mpifxcorr</code> processes are being started (not sent by <code>mpifxcorr</code>).
Starting	all the processes are ready to begin.
Running	the correlator is running.
Ending	the correlator has reached the end of the job.
Done	the correlation has completed.
Aborting	correlation is stopping early due to an error.
Terminating	correlation is stopping early due to signal.
Terminated	correlation has stopped early.
MpiDone	all of the MPI processes have ended (not sent by <code>mpifxcorr</code>).
Crashed	<code>mpifxcorr</code> crashed; usually sent by <code>mk5daemon</code> .

message a string containing information for the operator.

visibilityMJD the time-stamp (MJD + fraction) of last completed visibility record.

antId the antenna id for the associated weight, ranging from 0 to $N_{\text{ant}} - 1$.

weight the data weight for the associated antenna, ranging from 0 to 1. Note that in each XML document of this type there will in general be one *weight* value for each antenna being correlated.

8.9 DifxStop

Messages with *messageType* = **DifxStop** are typically sent by the DiFX Operator Interface to the **mk5daemon** running on the correlator head node to cause a particular instance of DiFX to be killed.

The *body* of the message contains:

```
<difxStop>
</difxStop>
```

8.10 Mark5DriveStatsMessage

Mark5 module conditioning is done periodically to ensure top performance of Mark5 modules. Each disk in the module gets written across its whole surface to identify bad areas and to *calibrate* the electronics. One message applies to one disk of the module

The *body* of the message contains:

```
<difxDriveStats>
  <serialNumber>serial</serialNumber>
  <modelName>model</modelName>
  <size>size</size>
  <moduleVSN>vsN</moduleVSN>
  <moduleSlot>slot</moduleSlot>
  <startMJD>startMJD</startMJD>
  <stopMJD>stopMJD</stopMJD>
  <binN>statsN</binN>
  <type>statsType</type>
  <startByte>startByte</startByte>
</difxDriveStats>
```

The italicized fields are as follows:

serial the serial number of the disk.

model the model number of the disk.

size the size of the disk, in GB.

vsN the module Volume Serial Number (VSN).

slot the location of the disk within the module, from 0 to 7.

startMJD the time when conditioning began.

stopMJD the time when conditioning ended.

statsN the histogram count for bin *N* for *N* in the range 0 to 7.

statsType type of operation which is one of **condition**, **condition.read**, **condition.write**, **read**, **write**, **unknown**, **test**.

startByte if not present, assumed to be zero; only relevant for some types of operations.

8.11 Mark5StatusMessage

This section describes messages with *messageType* = **Mark5StatusMessage**. This message type comes from either **mpifxcorr** or **mk5daemon** (or perhaps another program that makes heavy use of Mark5 units and wishes to volunteer status information).

The *body* of the message contains:

```

<mark5Status>
  <bankAVSN>vsnA</bankAVSN>
  <bankBVSN>vsnB</bankBVSN>
  <statusWord>statusWord</statusWord>
  <activeBank>activeBank</activeBank>
  <state>state</state>
  <scanNumber>scanNumber</scanNumber>
  <scanName>scanName</scanName>
  <position>position</position>
  <playRate>playRate</playRate>
  <dataMJD>dataMJD</dataMJD>
</mark5Status>

```

The italicized fields are as follows:

vsnA the VSN of the module in bank A.

vsnB the VSN of the module in bank B.

statusWord a hexadecimal number with the following bits: *TBD*

activeBank the active bank, either A or B for banks A and B respectively, N if the unit is in non-bank mode, or blank if no modules are active.

state the state of the Mark5 unit:

state	meaning
Opening	the Streamstor card is being opened.
Open	the Streamstor was successfully opened and is ready for use.
Close	the Streamstor has been closed.
GetDirectory	the unit is recovering the directory or finding data.
GotDirectory	the unit successfully found needed data on the module.
Play	the unit is playing back data.
PlayStart	the unit is about to start playback.
PlayInvalid	the unit is playing data, but the data is invalid.
Idle	the unit is not doing anything; no process has control of it.
Error	the unit is unusable due to an error.
Busy	the unit is busy and cannot respect commands.
Initializing	the Streamstor card is initializing.
Resetting	the unit is resetting the Streamstor card.
Rebooting	the unit is about to reboot.
Poweroff	the unit is about to turn off.
NoData	the unit is not playing data since there is none that is appropriate.
NoMoreData	the unit has played all the data for the job and is stopped.
Copy	data is being copied off the module to a local disk.

scanNumber the directory index number for the current scan. This number starts at 1.

scanName the name associated with the current scan.

position the byte position being accessed. Note that this number can be very large ($> 2^{46}$).

playRate the time-averaged playback rate in Mbps.

dataMJD the date stamp (MJD + fraction) of the most recently read data.

8.12 Mark5VersionMessage

This section describes messages with *messageType* = *Mark5VersionMessage*. This message comes from *mk5daemon*. It is typically broadcast once upon the start of *mk5daemon* and when requested.

The *body* of the message contains:

```
<mark5Version>
  <ApiVer>ApiVer</ApiVer>
  <ApiDate>ApiDate</ApiDate>
  <FirmVer>FirmVer</FirmVer>
  <FirmDate>FirmDate</FirmDate>
  <MonVer>MonVer</MonVer>
  <XbarVer>XbarVer</XbarVer>
  <AtaVer>AtaVer</AtaVer>
  <UAtaVer>UAtaVer</UAtaVer>
  <DriverVer>DriverVer</DriverVer>
  <BoardType>BoardType</BoardType>
  <SerialNum>SerialNum</SerialNum>
  <DaughterBoard>
    <PCBType>PCBType</PCBType>
    <PCBSubType>PCBSubType</PCBSubType>
    <PCBVer>PCBVersion</PCBVer>
    <FPGAConfig>FPGAConfig</FPGAConfig>
    <FPGAConfigVer>FPGAConfigVersion</FPGAConfigVer>
  </DaughterBoard>
</mark5Version>
```

Note that the `DaughterBoard` tag and its subtags are optional and are not broadcast if a daughter board is not detected on the Mark5C unit. The italicized fields are as follows:

ApiVer The software API version of the Streamstor API.

ApiDate Date associated with the above.

FirmVer The version of the firmware that is loaded.

FirmDate Date associated with the above.

MonVer The version of the Monitor FPGA code.

XbarVer The version of the cross bar FPGA code

AtaVer The version of the ATA disk controller FPGA code.

UAtaVer The version of the UATA disk controller FPGA code.

DriverVer The version number of the driver code.

BoardType The type of Streamstor board.

SerialNum The serial number of the Streamstor board.

PCBType The type of Streamstor daughter board.

PCBSubType Subtype of the above, if any.

PCBVersion The version of the daughter board.

FPGAConfig Name of the FPGA configuration.

FPGAConfigVersion Version number of FPGA configuration.

9 DiFX alert messages

This section attempts to list all of the messages you might see coming from the software correlator with some explanation about their meaning. For some messages, certain actions to be taken are suggested. Each subsection below contains descriptions of the messages for a particular severity level, ordered from most severe to least severe. There are almost 500 distinct messages that could be produced by `mpifxcorr` as of DiFX version 1.5.2, thus no effort has been made to be complete in the descriptions here. Effort has been made to document the most important ones in detail. Messages are sorted first by their severity level and are then roughly alphabetical within each subsection.

9.1 Fatal

All fatal errors will cause immediate termination of the correlation project. Most such errors would occur at the very start of a job as the input files are being processed.

- **Bin phase breakpoints are not in linear ascending order!!!:** The pulsar bins are not listed in phase order in the `.binConfig` file. Each `BIN PHASE END` entry must be greater than the previous and they must all be in the range 0 to 1.
- **Cannot create output directory *outDir*: flag - aborting!!!:** The output directory could not be created. The most common cause is an existing output file from a previous attempt to correlate the job in question. Other possible causes include: permission issues, inaccessibility of output directory to the DiFX head node, and output filesystem being full.
- **Cannot locate *stationName* in delay file *delayFile* - aborting!!!:** The specified delay file does not contain needed information for a station called *stationName*. This should never happen for data going through the correlator in a standard way.
- **Cannot open Streamstor device. Either this Mark5 unit has crashed, you do not have read/write permission to /dev/windrvr6, or some other process has full control of the Streamstor device.:** This message will only come from a Mark5 unit that is requested to play back data. The most likely cause of such a problem is the Streamstor card getting stuck in a compromised state, although fresh correlator installations that may not have left `/dev/windrvr6` with global read/write permission is a second likely cause of this problem. The fix likely requires a reboot of the Mark5 unit. On occasion a full power cycle of the Mark5 unit (not just a soft reboot) is required.
- **Cannot open file *inputFile* - aborting!!!:** The `.input` file named *inputFile* is not readable by the software correlator. This could be due to one of a number of issues, including: read permission problems, the request file not existing, or the file existing but not visible from one or more of the software correlator nodes.
- **Cannot open delay file *delayFile* - aborting!!!:** The `.delay` file named *delayFile* is not readable by the software correlator. This could be due to one of a number of issues, including: read permission problems, the request file not existing, or the file existing but not visible from one or more of the software correlator nodes.
- **Cannot open output file *outputFile* - aborting!!!:** The output file could not be created. The most common cause is an existing output file from a previous attempt to correlate the job in question. Other possible causes include: permission issues, inaccessibility of output directory to the DiFX head node, and output filesystem being full.
- **Cannot open pulsar config file *binConfigFile* - aborting!!!:** The specified pulsar bin file, *binConfigFile* cannot be opened. This could be due to one of a number of issues, including: read permission problems, the request file not existing, or the file existing but not visible from one or more of the software correlator nodes.

- **Cannot open uvw file *uvwFile* - aborting!!!:** The specified `.uvw` file, *uvwFile* cannot be opened. This could be due to one of a number of issues, including: read permission problems, the request file not existing, or the file existing but not visible from one or more of the software correlator nodes.
- **Cannot quad interpolate delays with post-f fringe rotation - aborting!!!:** Two mutually exclusive options (QUAD DELAY INTERP and POST-F FRINGE ROT) were both enabled in the `.input` file.
- **Caught an MPI exception!!! *errorString*:** A process communication error has occurred causing the correlator to terminate. The outcome of such an event cannot be good; contact a DiFX developer.
- **Config encountered inconsistent setup in config file - aborting!!!:** One or more of the configurations (group of settings defined in the `.input` file) is either illegal or incompletely defined. Usually this message will come with another more detailed error message. In any case, either there is a correlator version mismatch or there is something wrong with the `.input` file.
- **Core received a request to process data from time *time* which does not have a config - aborting!!!:** This is likely due to failed consistency check in the software correlator that resulted from a logic error in the code. This should be reported to a DiFX developer.
- **Could not find station *stationName* in the uvw file when making rpfits header!!! This station is used in the correlation so I will abort!!!:** This error only occurs with RPFITS output format which is not supported by this documentation – please seek other sources of assistance if needed.
- **Could not locate any of the specified sources in the specified time range - aborting!!!:** The correlator gave up since none of the sources to be correlated appeared in the `.uvw` file. This should never happen for data going through the correlator in a standard way.
- **Could not locate a polyco to cover the timerange ...:** A pulsar polynomial for a certain time period could not be found. The person supplying the pulsar polynomial should be contacted.
- **DataStream assumes long long is 8 bytes, it is *x* bytes - aborting!!!:** This should not occur on any modern operating system. If this message is seen, please contact a DiFX developer and be sure to indicate exactly which operating system and computer type you are using.
- **DataStream assumes int is 4 bytes, it is *x* bytes - aborting!!!:** This should not occur on any modern operating system. If this message is seen, please contact a DiFX developer and be sure to indicate exactly which operating system and computer type you are using.
- **DataStream *mpiid*: expected *x* bytes, got *y* bytes - aborting!!!:** In an eVLBI read, the wrong number of bytes was received, indicating a mismatch in send/receive setups. Check the setup at both ends and try again.
- **Datastream *mpiid*: implied UDP packet size is negative - aborting!!!:** When considering the size of a UDP header, an unphysical packet size results. This should only occur when attempting UDP based eVLBI.
- **Datastream *mpiid*: read too few UDP packets: bytestoread=*x* udp_offset=*y* bytes=*z* - aborting!!!:** Fewer than expected UDP packets were received in an eVLBI transfer (FIXME – more details please!)
- **Datastream *mpiid*: read too many UDP packets: bytestoread=*x* udp_offset=*y* bytes=*z* - aborting!!!:** More than expected UDP packets were received in an eVLBI transfer (FIXME – more details please!)

- **Datastream *mpiid*: could not allocate databuffer (length *size*) - aborting!!!:** A memory allocation failed. This would probably be due to either a developer error or attempt to run a job on an underpowered computer. In any case, the developer should be contacted.
- **Developer error: Cannot handle delays more negative than *maxDelay*. Need to unimplement the datastream check for negative delays to indicate bad data - aborting!!!:** The maximum negative delay (which is quite large) has been exceeded. This is almost certainly a logic error in the software and should be reported to a DiFX developer.
- **Developer error: in `Mk5Mode::Mk5Mode`, `mark5stream` is null:** An internal error related to initializing a Mark5 decoder has occurred. Contact a DiFX developer.
- **Developer error: in `Mk5Mode::Mk5Mode`, `framesamples` is inconsistent (*x/y*):** An internal inconsistency has been found in the Mark5 data frame size that would lead to downstream errors. This could only be caused by a software logic error. Contact a DiFX developer.
- **Developer error: UVW has not been created!!!:** This message would come from an internal consistency check that failed. If this occurs, the DiFX developers should be notified as this indicates a logic error inside the software correlator.
- **`genMk5FormatName : formatType format : framebytes = frameBytes` is not allowed:** An illegal frame size was specified in the `.input` file. This should never happen for data going through the correlator in a standard way. If you see this message it is probably due to a programming error and the DiFX developers should be notified.
- **`genMk5FormatName : unsupported format encountered`:** A data format not handled by the software correlator has been requested in the `.input` file.
- **Input file out of order!:** There is an error in the input file. This might be caused by using an input file formatted for one version of `mpifxcorr` on a different version or by an incorrectly written file.
- **Manager aborting correlation!:** The configuration of a visibility buffer was not OK. There are many possible causes for this, but is most likely due to a logic error in the software. Contact a DiFX developer if this is encountered.
- **`Mk5DataStream::readnetwork` bytestoread too large (*x/y*) - aborting!!!:** An eVLBI read size is too large. (FIXME – more details here please!)
- **`mpifxcorr` must be invoked with at least *x* processors (was invoked with *y* processors) - aborting!!!:** For a correlation of *N* datastreams (usually equal to the number of antennas), at least *N* + 2, but preferably even more, processes must be started.
- **`NativeMk5DataStream::NativeMk5DataStream` stub called, meaning `mpifxcorr` was not compiled for `nativemk5` support, but it was requested (with `MODULE` in `.input` file) - aborting!!!:** Correlation directly off a Mark5 module requires compiling against the Streamstor libraries which was apparently not done but requested by the `.input` file. Contact the person responsible for your correlator setup and ask them to properly link the Streamstor libraries to the `mpifxcorr` executable.
- **No config section in input file:** The input file is missing its config section and hence correlation cannot proceed. This should never happen for data going through the correlator in a standard way.
- **Not enough baselines are supplied in the baseline table (*x*) compared to the number of baselines (*y*)!!!:** The `.input` file requests more baselines in the common table than there are enumerated in the baseline table. This should never happen for data going through the correlator in a standard way.

- **Not enough datastreams are supplied in the datastream table (x) compared to the number of datastreams (y)!!!** The `.input` file requests more datastreams (nominally equal to the number of antennas) in the common table than there are enumerated in the baseline table. This should never happen for data going through the correlator in a standard way.
- **Please invoke with `mpifxcorr ...`:** The correlator was not started according to its usage. See §6.71 for more details.
- **Polyco `polycoId` / `subcount` is malformed** The pulsar polynomial file is not compliant (possibly due to manual editing).
- **RPFITS not compiled in - aborting:** RPFITS output format is requested, but support for RPFITS is not compiled into `mpifxcorr`. Either requested output format, or recompile `mpifxcorr` with RPFITS. Note: This document does not contain instructions for RPFITS installations – you are on your own!
- **Samplesperblock is less than 1, current implementation cannot handle this situation - aborting!!!:** An illegal data sub-mode has been requested. Contact a developer.
- **Unknown data format `formatName`:** A data format (e.g., VLBA, Mark4, LBA) has been requested that cannot be processed.
- **Unknown data source `sourceName`:** A data source (e.g., FILE, MODULE, EVLBI) has been requested that cannot be used.

9.2 Severe

Severe errors typically reflect an unexpected software error. All severe errors are related to a process control (threading) failure or a failure in a numerical routine. Errors of the severe type are likely to cause widespread erroneous results or complete failure of correlation. Except during periods of software development errors of these two types are unexpected. If encountered, they should be reported to a DiFX developer and the correlation should be reattempted once. Since there are many errors in the “severe” class, all unlikely and with the same procedure for working around, individual errors of this type are not listed below.

9.3 Error

Messages in the “Error” class are typically fairly significant, often cascading to fatal messages and termination of the correlation. Many errors of the Mark5 variety result in data loss ranging from fractions of a second to the full job in length. For errors of this type, operator judgement is needed: whether to restart correlation or keep going will depend on many circumstances. Note that many of the eVLBI and real-time monitor errors are not documented here yet.

- **All data from this module was discarded: ...:** Due either to malformed data or a directory file with incorrect values, no data was deemed suitable for correlation for a particular Mark5 module. It is probably worth trying to extract the module directory structure again and trying the job again and/or moving the module to a different unit for playback. It is likely that other jobs using this module will face a similar problem.
- ***n* consecutive sync errors. Something is probably wrong!:** A large number of sync errors were seen. This probably means that the Mark5 data being read is somehow corrupted.
- **All bandwidths for a given datastream must be equal:** Currently, `mpifxcorr` does not allow different bandwidths on different frequency bands. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.

- **All configs must have the same telescopes! Config m datastream n refers to different telescopes:** The same set of telescopes (antennas) must be used throughout a job. Two configurations have been found that use different antenna subsets. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **All LBASTD Modes must have 2 bit sampling - overriding input specification!!!:** The Australian LBASTD data format modes can only handle 2-bit (4 level) quantization at the moment. The `.input` file value for QUANTISATION BITS is being ignored here. Probably the maker of the `.input` file made a mistake.
- **Attempting to get a delay from offset time $time$, will take first—last source:** Correlation is requested for a time either before or after the list of scans. Data affected by this is likely to have the wrong delay applied and is unlikely to be useful. The creator of the input files should be contacted.
- **Attempting to refer to freq outside local table!!!:** The frequency index of a datastream table exceeds the length of the frequency table in the `.input` file. This will most certainly cause this frequency band to be incorrectly correlated.
- **Baseline table entry m , frequency n , polarisation product p for datastream q refers to a band outside datastream q 's range (r):** The band index for a baseline table exceeds its legal range. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **Baseline table entry m has a datastream index outside the datastream table range! Its two indices are n , p :** The baseline table has a datastream index that exceeds the number of datastreams. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **Cannot clear Mark5 write protect:** The software correlator attempts to set the Disk Module State to PLAYED. This requires temporarily disabling write protection. If this fails, then either the Mark5 unit is in a bad state or the module has a problem.
- **Cannot open data file $dataFile$:** The datastream process cannot open the specified file containing baseband data to be correlated. Correlation will proceed, but no data for an antenna for the duration of this file will be correlated. Usually this error should be a cause for concern.
- **Cannot open polyco file $polycoFile$:** Either file permissions, file location, or non-existence is preventing the file called `polycoFile` from being opened. This will likely end badly.
- **Cannot put Mark5 unit in bank mode:** The command to put the Mark5 unit in single bank mode failed and further commands to the Mark5 unit will probably fail as well. Usually this only happens if the Mark5 unit is in a bad state. The Mark5 unit probably needs a reboot.
- **Cannot read data from Mark5 module...:** Read from a Mark5 module failed. No further attempt to read data from the module will be made. It is strongly recommended that the Mark5 unit be rebooted and the correlation be reattempted. If the error is reproducible, the additional information contained at the end of this error message may help diagnose the problem.
- **Cannot read the Mark5 module label:** The command to retrieve the module label, which includes the volume serial number (VSN) and the previous state, has failed. This implies trouble with the Mark5 unit or module. First the module should be moved to a different Mark5 unit and the correlation reattempted. Upon further failure, the module should be checked for problems
- **Cannot set Mark5 data replacement mode / fill pattern:** Either the command to tell the Mark5 unit to enter real-time playback mode or the command to set the fill pattern have failed; expect more problems with this Mark5 unit. The Mark5 unit probably needs a reboot.
- **Cannot set the Mark5 module state:** The software correlator failed to set the Disk Module State to PLAYED. If this fails, then either the Mark5 unit is in a bad state or the module has a problem.

- **Cannot set Mark5 write protect:** The software correlator attempts to set the Disk Module State to PLAYED. This requires temporarily disabling write protection. If this fails, then either the Mark5 unit is in a bad state or the module has a problem.
- **Cannot unpack Mark5 format data at sampleoffset n from buffer *time*:** An error of this kind likely represents a logic error in the software correlator and should be reported to a DiFX developer.
- **Config m baseline index n refers to baseline p which is outside the range of the baseline table:** The baseline index of a config table exceeds the number of baselines in the baseline table. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **Config m baseline index n refers to baseline p which is out of order with the previous baseline ...:** Entries in the datastream table are not in the expected order. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **Could not find a polarisation pair, will be put in position $x!!!$:** This is due either to an uncaught inconsistency in the `.input` file or a logic error in the software correlator. Contact a DiFX developer.
- **Could not find any bands for frequency m of datastream n :** This is due either to an uncaught inconsistency in the `.input` file or a logic error in the software correlator. Contact a DiFX developer.
- **Could not open *threadFile* - will set all numthreads to 1!!!:** The requested `.threads` file could not be opened. Possible causes include: permission issues, inaccessibility of directory to the DiFX head node, and the file simply not existing. The impact of this is that correlation will proceed at a potentially much reduced speed since each CPU will use only a single processing thread. Accuracy of the results will not be affected.
- **Could not parse LBA file header:** An LBA format file appears corrupt. Data for one antenna for the duration of this file will not be correlated.
- **Datastream table entry m has a frequency index (freq n) that refers outside the frequency table range (p):** The `.input` file has a frequency index error. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **Datastream table entry m has an input band local frequency index (band n) that refers outside the local frequency table range (p):** The `.input` file has a frequency index error. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **Datastream table entry m has a telescope index that refers outside the telescope table range (n):** The `.input` file has a telescope index error. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **FFT chunk time for config m , datastream n is not a whole number of nanoseconds (p):** In order to keep track of time properly, each FFT must start on an integer nanosecond boundary. There are currently no modes supported where this should be a problem, so if you see this message, there is a more serious problem. Contact a DiFX developer!
- **First datastream for baseline n has a higher number than second datastream - reversing!!!:** The entries of the baseline table (indicating which antenna pairs to correlate) should always have the datastream indices in ascending order. If you see this warning, the correlator is correcting your baseline ordering, but be aware that this might be hinting that something else might be awry with the `.input` file. The creator of the `.input` file should be contacted.
- **Hit the end of the file! Setting the numthread for Core n to 1:** The `.threads` file ended unexpectedly early and one or more core processes will be forced to use a single processing thread, with potentially crippling performance penalty. Accuracy of the results will not be affected.

- **Increment per read in nanoseconds is x - too large to fit in an int:** The way timekeeping works in DiFX, data chunks larger than $2^{31} - 1$ nanoseconds in length are not possible at the moment. The maker of the `.input` file needs to reduce the read size by changing some of the parameters (such as `NUM CHANNELS`, `BLOCKS PER SEND`, or possibly others).
- **lastoffsetns less than 0 still! = x :** This message probably indicates a logic error in the correlator program so should be reported to a DiFX developer.
- **Mk5DataStream::calculateControlParams : vlbaoffset= x bufferindex= y atsegment= z :** A Mark5 data frame was found unaligned. A corresponding subintegration of data will be invalidated. A large number of messages of this type probably indicates corrupted data.
- **Module VSN contains undecoded scans!:** The module directory for module *VSN* has problems. Please correct the problem with the directory (as stored in `$MARK5_DIR_PATH`) and try again.
- **Module VSN not found in unit!:** The `.machines` file suggested that a specified Mark5 module would be found this this Mark5 unit but it was not. Check to make sure the module is in the unit.
- **Most of the data from this module was discarded: ...:** Due either to malformed data or a directory file with incorrect values, a large fraction of the data for this job from a particular module was not decodable. It is probably worth trying to extract the module directory structure again and trying the job again and/or moving the module to a different unit for playback. It is likely that other jobs using this module will face a similar problem.
- **No valid data found. Stopping playback!:** According to the directory for the module in question there is no valid data available to correlate. It is possible that a directory reconstruction will allow some or all of the data on the module to be retrieved.
- **Not all datastreams accounted for in the datastream table for config m :** All datastreams (roughly equivalent to antennas) must be represented in each configuration within the `.input` file. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **Number of input bands for datastream m (n) does not match with Mark5 file `fileName` (p), will be ignored!!!:** The description of Mark5 baseband data (VLBA or Mark4 format) in the `.input` file is inconsistent with the actual content. The creator of the `.input` file should verify that the format is correctly specified. All data related to this condition will be left uncorrelated.
- **Oversamplefactor (m) is less than decimation factor (n):** The oversample factor must be an integer r times the decimation factor. Essentially oversampling is handled by two mechanisms: decimation of the input data stream in the data unpacker and through spectral selection at the time of FITS file creation. The oversampling factor of these two approaches must multiply to be the total *Oversamplefactor*. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **Pulsar phase as calculated fromt the polyco will not be accurate over entire range of *time* as the frequency is changing too rapidly. The maximum safe calc length would be x - try reducing blocks per send or numchannels ...:** A detail of the way the pulsar polynomial is used is likely to cause imperfect binning. Please consult with the PI and the creator of the `.input` file.
- **Stale data was received from core n regarding time *time* seconds - it will be ignored!!!:** One subintegration of data is being discarded as it did not arrive in time for the data to be written to disk. If this is a chronic problem, increasing the value of `VIS BUFFER LENGTH` in the `.input` might help. A small number of errors of this type is not a problem.
- **Telescope *antName* could not be found in the uvw file!!!:** Data for this antenna will not have useful UVW values in its output file. Contact the creator of the `.input` file.

- **There must be an integer number of sends per dataset.** Presently `databufferfactor` is m , and `numdatasetsegments` is n : The `.input` file contains an inconsistency in its parameterization of the various buffer sizes. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **Trying to read past the end of file!!!:** One of the files read by `mpifxcorr` ended prematurely. The maker of the `.input` file should be contacted.
- **Unknown output format *outFormat* (case sensitive choices are RPFITS, SWIN and ASCII):** `mpifxcorr` currently supports 3 output formats (ASCII, DIFX and RPFITS) and the requested one does not match one of these. This will lead to termination of the correlator. The creator of the `.input` file should be contacted.
- **Unsupported format or mode requested!!!:** A data format has been requested that is recognized but not supported by the correlator. Scans using this format (probably all scans for the antenna in question) will produce no data. The maker of the `.input` file should be contacted.
- **Waited 6 seconds for a Mark5 read and gave up.:** A Mark5 read timed out. A small number of such errors can be tolerated, especially for a module that is known to be bad, but many such messages should prompt a second correlation attempt after module move / unit reboot.
- **We thought we were reading something starting with '*something*', when we actually got '*somethingElse*':** The ordering or content of one of the files being read by `mpifxcorr` does not match expectations and is thus non-conformant. The maker of the file should be contacted. Note that this could be caused by a version mismatch between in the input files and the correlator.
- **XLRCardReset() failed. Remainder of data from this antenna will not be correlated and a reboot of this Mark5 unit is probably needed.:** It is probably worth reattempting correlation after Mark5 reboot and possible module move.
- **XLROpen() failed. Remainder of data from this antenna will not be correlated and a reboot of this Mark5 unit is probably needed.:** After a successful Streamstor card reset, the card was not able to be accessed. It is probably worth reattempting correlation after Mark5 reboot and possible module move.

9.4 Warning

Warning level messages typically relate to a situation that may result in the loss of a very small amount of data or indicate some other irregularity to the operator. A single warning should not be of concern, but a large number of warnings should be noted.

- **n consecutive fill patterns at time *time*:** Instead of reading data from the module, the Streamstor card returned fill pattern for this many consecutive data frames. If the module is flakey, it may be possible to recover more data after moving the module to a different Mark5 unit, but usually warnings of this type indicate inevitable data loss due to fill pattern replacement.
- **n consecutive sync errors starting at readpos p ...:** Data was read off the module, but the sync word was not found. This either indicates attempted playback of the wrong data format, completely corrupted data, or valid data that has slipped samples and is thus unfortunately not processable by the software correlator. It might be worth a recorrelation attempt after moving the module, but not likely.
- **Cannot find a valid configindex to set Mk5-related info. Flagging this subint:** Application of the delay model caused a small amount of data to cross a scan boundary in a manner that required the flagging of an entire subintegration. Such an event should be very rare.

- **Connection to monitor socket still pending:** When real-time monitoring of the output visibilities is requested but no connection to a monitor program has been established this warning may appear. A warning of this type is not associated with any data loss.
- **Copying a polyco with no frequency information!:** A pulsar polynomial without frequency information has been found. This could present problems in setting the gate properly across frequency channels, but could be intentional.
- **Could not find station *stationName* in the uvw file when making rpfits header!!! This station is not used in this correlation so its parameters will be initialised to 0!!!:** The `.input` file included a station not in the `.uvw` file, but this station is not used so the results won't be affected. This warning will only be issued for data being written in RPFITS format. The resultant RPFITS file will still list the missing antenna, but its information will be bogus. This could confuse downstream software.
- **Could not open command monitoring socket! Aborting message receive thread.:** Real-time monitoring was wanted, but a problem at the operating system level (perhaps permissions?) prevented the needed socket from being created. This has no bearing on the quality of the output data.
- **Data was received which is too recent ($x\text{sec} + y\text{ns}$)! ...:** A portion of one visibility record will have incomplete weight due to forced flushing of the long term accumulator buffer for that visibility record. A single warning of this type on source change is nothing to worry about, but constant warnings of this should be reported as this may indicate a different problem.
- **DataStream *mpiid*: could not identify Mark5 segment time (*formatName*):** An eVLBI data packet could not be decoded and hence its time cannot be determined. This is usually not a problem since time can be dead-reckoned from previous data and the corrupt packet won't be used anyway.
- **Filterbank not yet supported!!!!:** The filterbank mode, enabled with the `FILTERBANK USED` option in the `.input` file, was requested, but is currently not supported. When supported, the filterbank mode will offer options for *crisper* spectral channels.
- **Fractional start time of x seconds plus y ns was specified, but the start time corresponded to a configuration not specified in the input file and hence we are skipping z seconds ahead! The ns offset will be set to 0!!!:** This warning indicates that the start time of the job is not contained within a scan to be correlated and the subsequent start time will be rounded to the next integer second. Except in cases where exact timestamp matching is needed this is not a problem. The warning is issued as it is not normal for the start time of a job to be outside a scan. If you see this a lot, contact the person generating your jobs and let them know.
- **FXMANAGER caught a signal and is going to shut down the correlator:** A terminate signal was sent to the manager process indicating that the correlator should be stopped immediately. The resultant output file will be incomplete, but the early stop was probably on purpose so that should be expected.
- **Hit end of first line prematurely - check your polyco conforms to standard! Some values may not have been set properly, but likely everything is ok:** The pulsar polynomial file had an unexpectedly short first line. Some parameters may not have been properly set so the pulsar gating may not work. Retrying correlation will not help! If you are concerned, contact the producer of the polynomial.
- **Hit end of second line prematurely - check your polyco conforms to standard! Some values may not have been set properly. This often happens for non-binary pulsars. Likely everything is ok.** It is possible that the pulsar polynomial file is incomplete, but more likely nothing is wrong.

- **Incorrectly set config index at scan boundary! Flagging this subint:** A subintegration that crosses the end of a scan got flagged as a result of a potential format change at this time. This should happen no more than once per scan and affects only a small fraction of one integration period in most cases.
- **Internal Error, trying to copy pass buffer size:** This is an eVLBI related error. (FIXME – please add more description here)
- **Mk5DataStream::calculateControlParams : bufferindex= x >= bufferbytes= y :** A fraction of a dataframe is being discarded due to a frame misalignment. This affects a very small amount of data and should be very rare.
- **Module label is not terminated!:** A Mark5 module has an oversized extended serial number label. This in itself is not a problem but may indicate either corruption on the module or poor seating of either the Mark5 module or one of the cables/cards inside the Mark5 unit.
- **Module label record separator not found!:** No “Disk Module State” was stored on this module. This is probably due to recording on a very old version of Mark5A or is possibly the result of some other incompatibility.
- **MPI Id *mpiid*: warning - received a parameter instruction regarding *paramName* which cannot be honored and will be ignored!** The software correlator can receive a number of different commands that can affect its dumping of its long-term accumulator to an external piggy-back processor. If an unrecognized parameter is received a warning of this type will be issued. This is completely unrelated to correlator data quality.
- **No more data for project on module *mpiid*:** Although data on the module extends past the end of the current job, none of the remaining data is relevant for this job so playback is stopping early. This is only a problem if it is due to an incorrect transcription of the module directory. It might be prudent to look at the observe log to see if data is expected.
- **No more data on module *mpiid*:** The Mark5 unit sending this message is stopping its reading before the end of the job since no more data exists. This is only a problem if it is due to an incorrect transcription of the module directory. It might be prudent to look at the observe log to see if data is expected.
- **Post-f fringe rotation not yet tested - use at your own risk!!!:** Post FFT fringe rotation (enabled with the POST-F FRINGE ROT .input file option) has not been extensively tested. Results may be OK, but you are on your own!
- **Trying to read s seconds past the end of the UVW array!:** The .uvw file does not cover an entire scan. If more than a couple such warnings are seen, the creator of the .input and .calc file should be notified.
- **Waited s sec state=*state*:** A Mark5 module is taking longer than expected to respond. Additional messages will follow if this situation is serious.
- **XLRCardReset() being called!:** If a read timeout occurs, the Streamstor card in the Mark5 unit will be reset. This message indicates that this non-standard procedure is occurring. The result of this reset will usually either be success (in which case a message will indicate such), failure (in which case an error message will state that no more data will be read from the module), or a Mark5 unit hang, and hence the correlation will stop. If this is not successful, it is recommended that the correlation be retried after rebooting the affected Mark5 unit and possibly moving the module to a different Mark5 unit.

9.5 Info

Info messages convey normal messages to the operator.

9.6 Verbose

Verbose messages convey normal messages to the operator that are either not very important or come in vast quantities; they are typically filtered out of data logging to prevent unnecessary bloat.

9.7 Debug

Debug messages are useful only to developers and don't usually indicate error conditions. In some cases they might be useful in diagnosing a problem. None of the debug messages are explicitly documented here because they will be very version dependent and in general provide little value to the operator.

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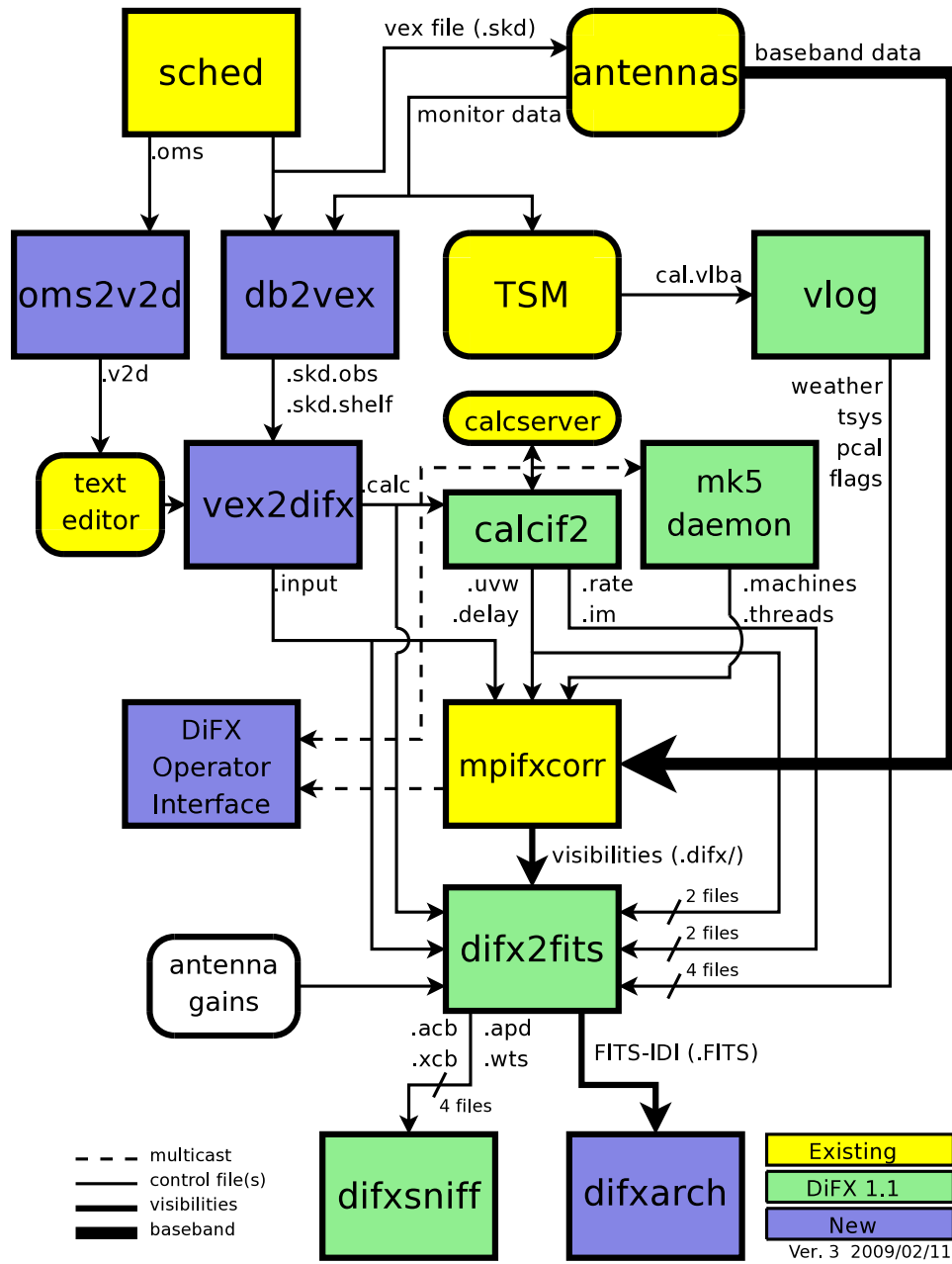


Figure 1: The DiFX software correlator block diagram as implemented for the VLBA.

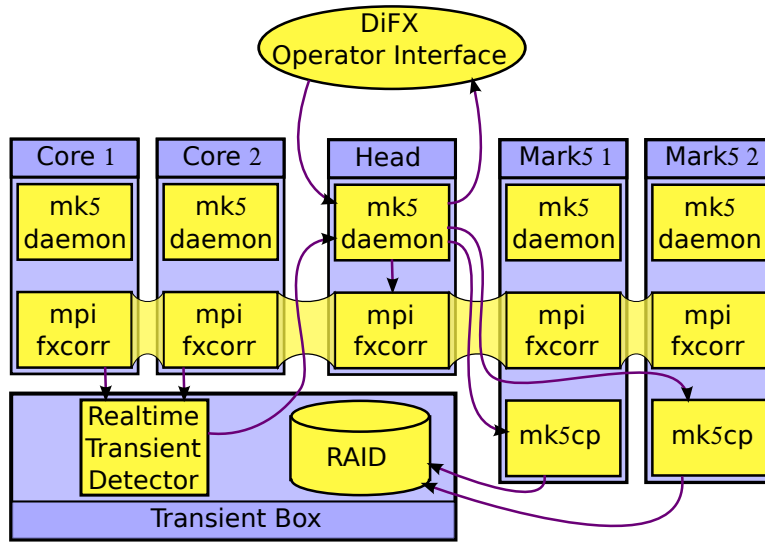


Figure 2: *The sequence of events leading to transient data capture.* First the DiFX Operator Interface sends a `DifxStart` message to the head node, causing the `mpifxcrr` process to be started. The core nodes of `mpifxcrr` will transmit autocorrelation data to a real-time transient detection algorithm. As interesting events are identified, `DifxTransientMessage` documents are sent to the head node; the `mk5daemon` process keeps a prioritized list of events. After correlation completes but before `mk5daemon` on the head node starts a series of `mk5cp` processes to run to copy the baseband data from the Mark5 modules before the next job starts.